

FROM RESEARCH TO INDUSTRY

cea

GRENOBLE



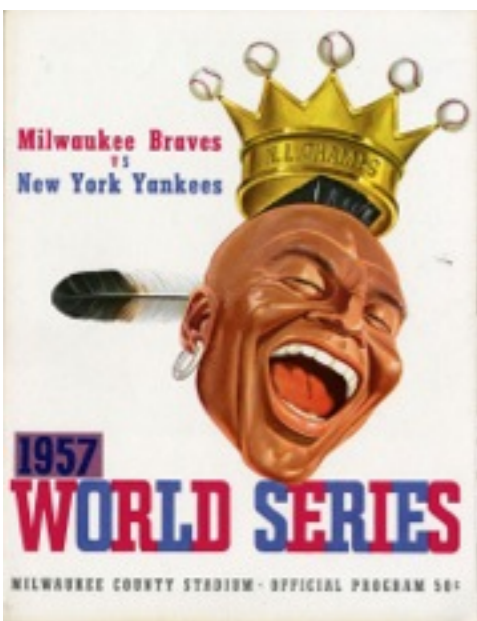
www.cea.fr

SPACE CRYOCOOLER DEVELOPMENTS

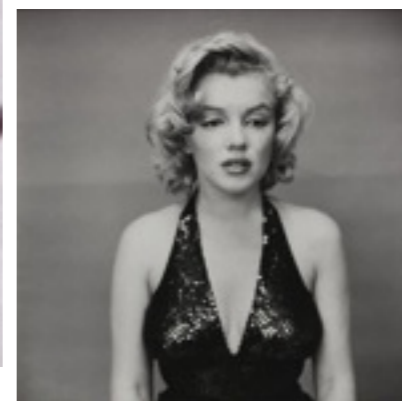


Lionel DUBAND
Univ. Grenoble Alpes, CEA-INAC-SBT, F-38000 Grenoble

WHEN IT ALL STARTED



'57 CHEVROLET! SWEET, SMOOTH AND SASSY!



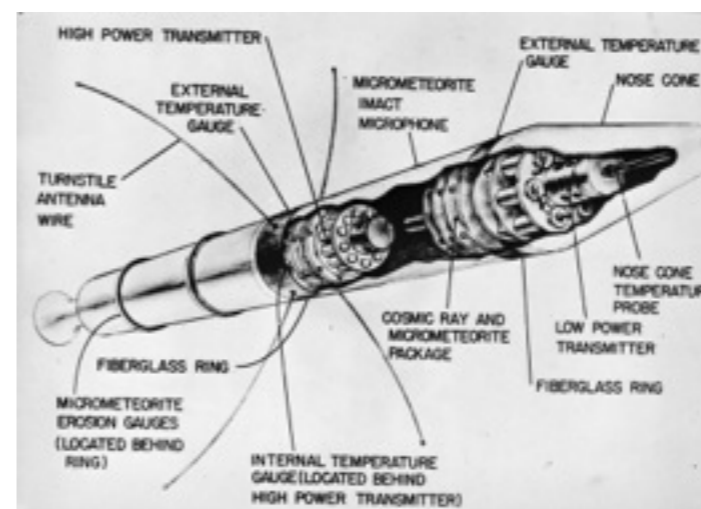
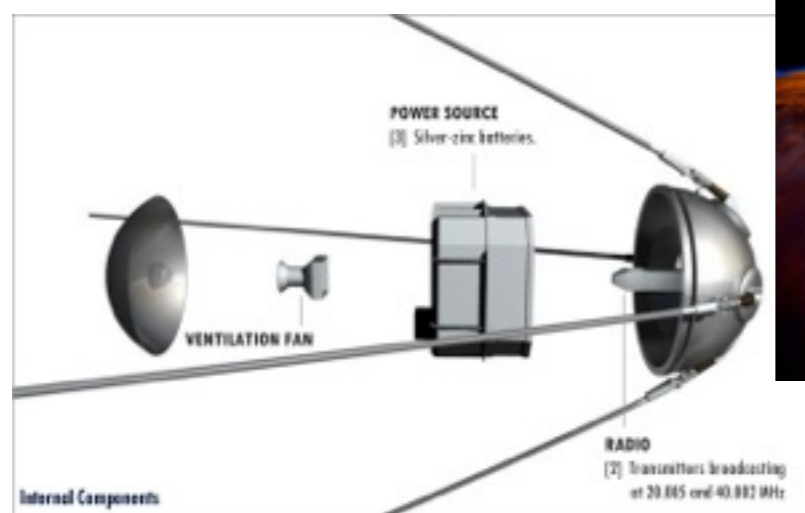
Spoutnik

to: 1957

Explorer I

October 1957

January 1958



JUPITER-C EXPLORER I

EXPLORER MAIN CHARACTERISTICS	
LENGTH	85 IN.
DIAMETER	5 IN.
WEIGHT	30.8 LB.
VELOCITY APPROX.	10,000 MPH
APOGEE ALTITUDE	1,700 MI.
PERIGEE ALTITUDE	325 MI.
PERIOD	104.9 MIN.
MAXIMUM ACCELERATION	33.3 G'S

WHY SPACE ?

Earth Observation



50 - 200 K

Planetary Exploration

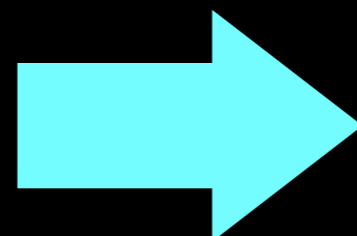


Scientific Missions

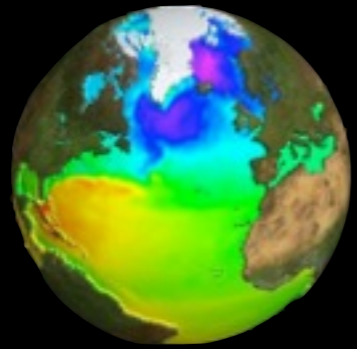
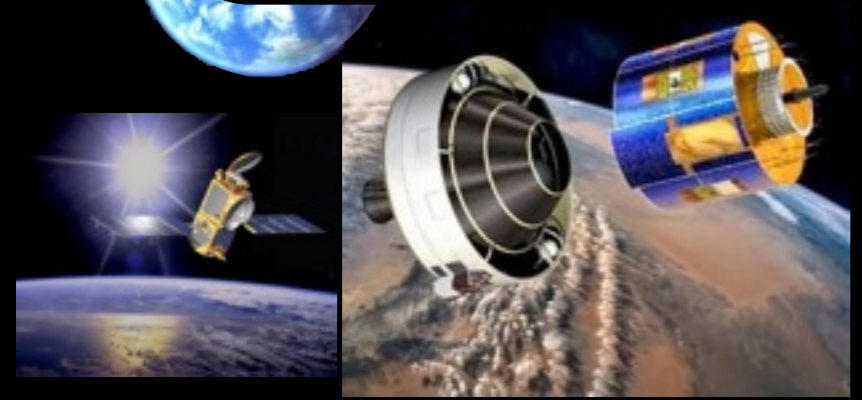


0.05 - 100 K

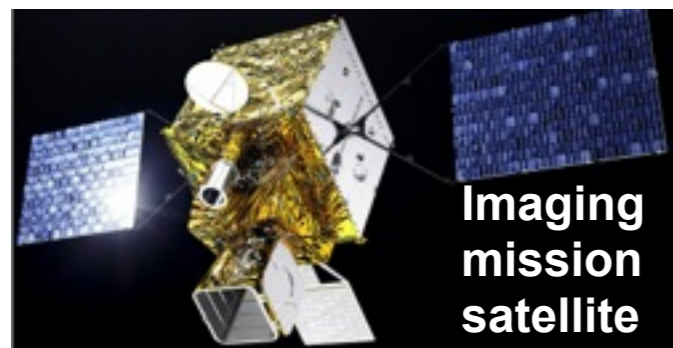
Cryogenics



- Detectors
- Optics



EARTH OBSERVATION: WEATHER SATELLITE MTG

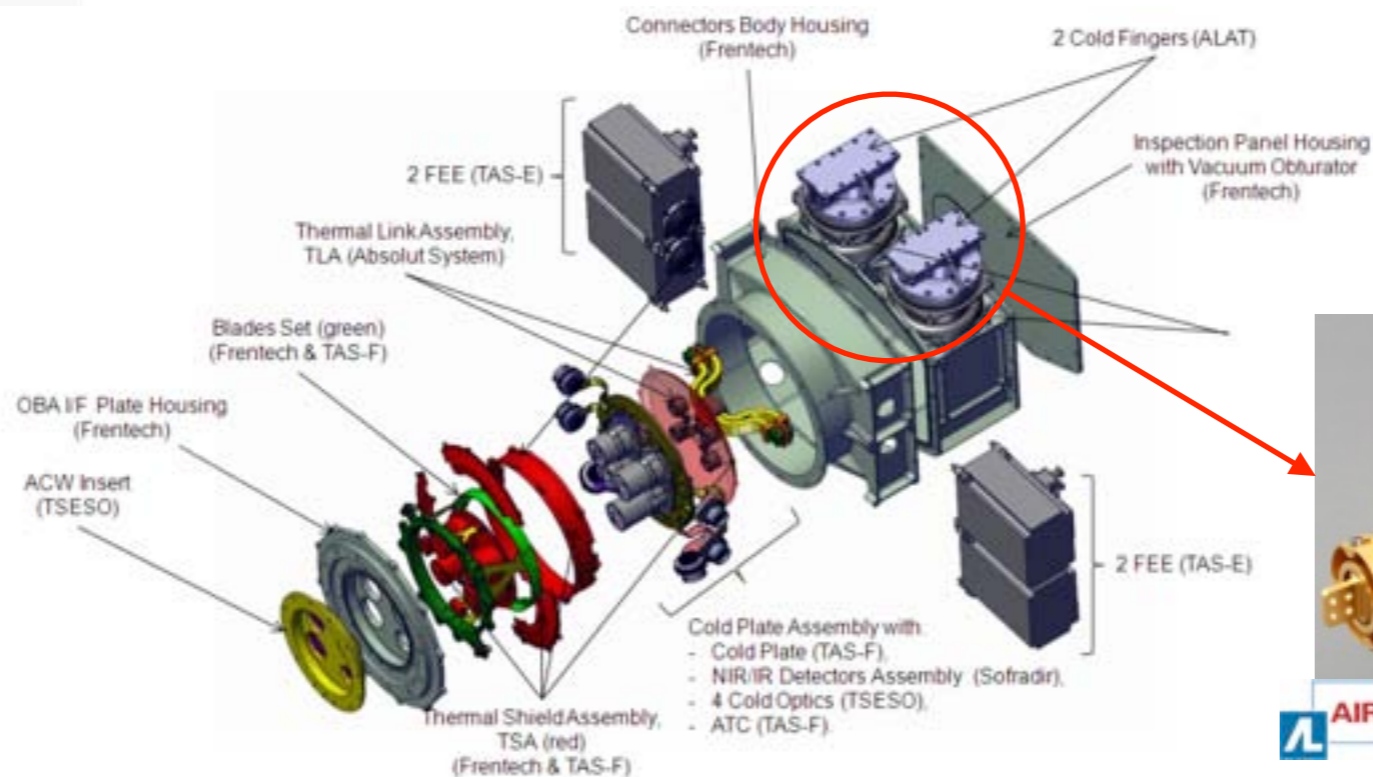
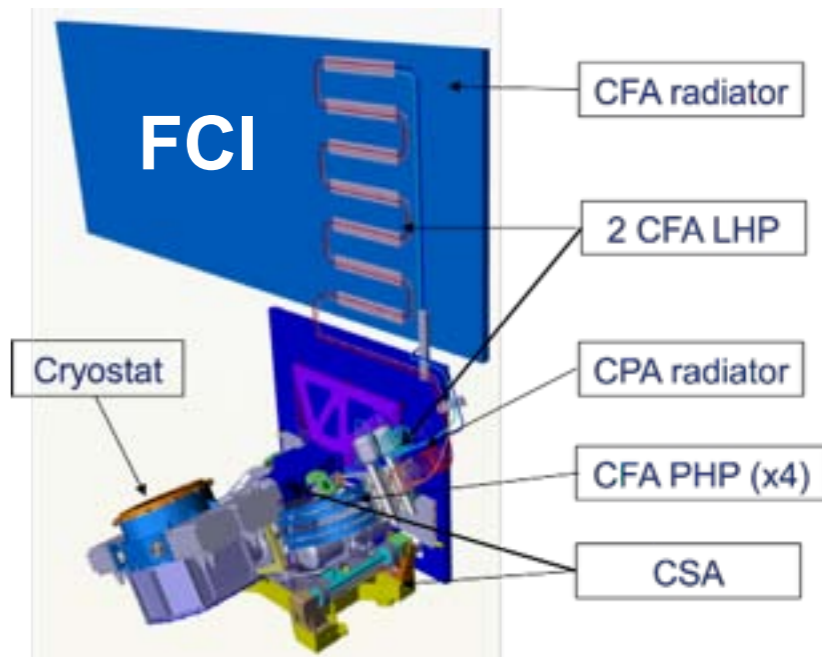
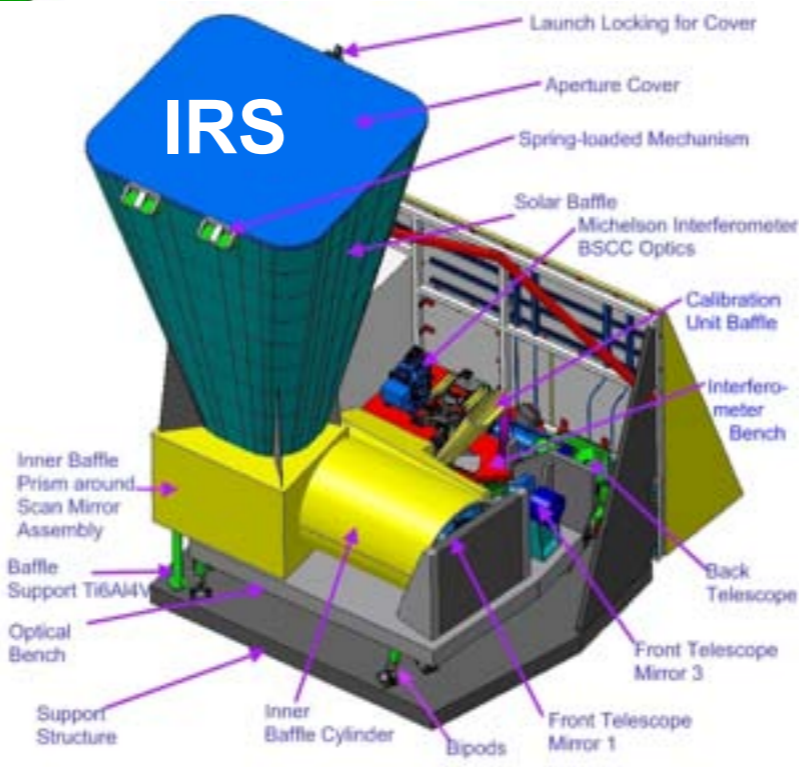
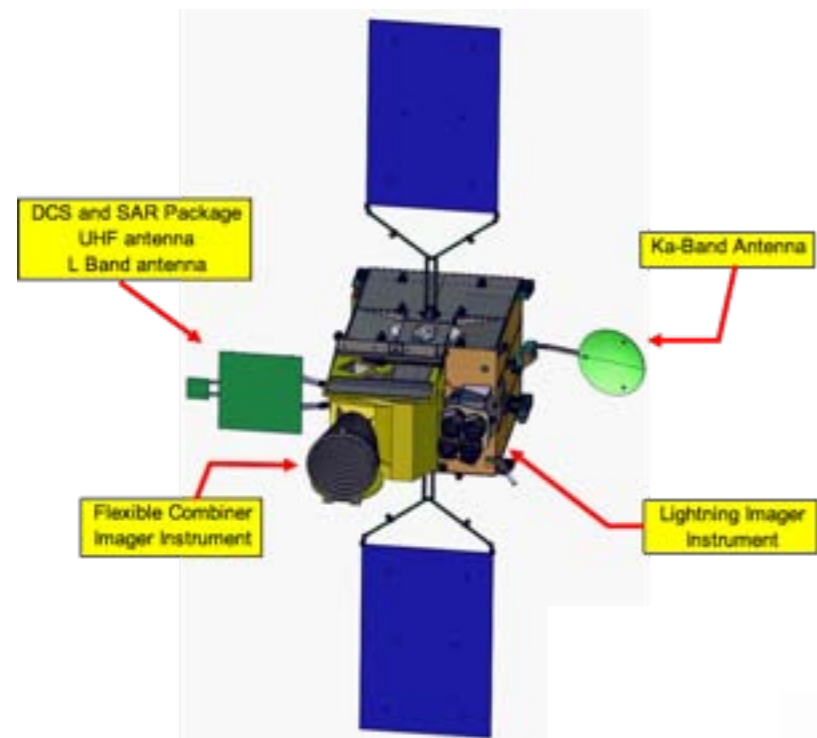


2 instruments

IRS: *Infrared Michelson Interferometer*

FCI: *High spatial and spectral resolution imagery*

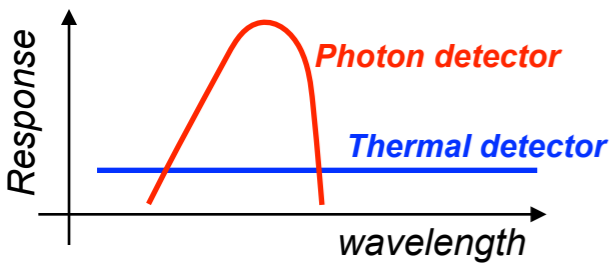
55 K and 60 K needed
(photoconductors HgCdTe)



AIR LIQUIDE

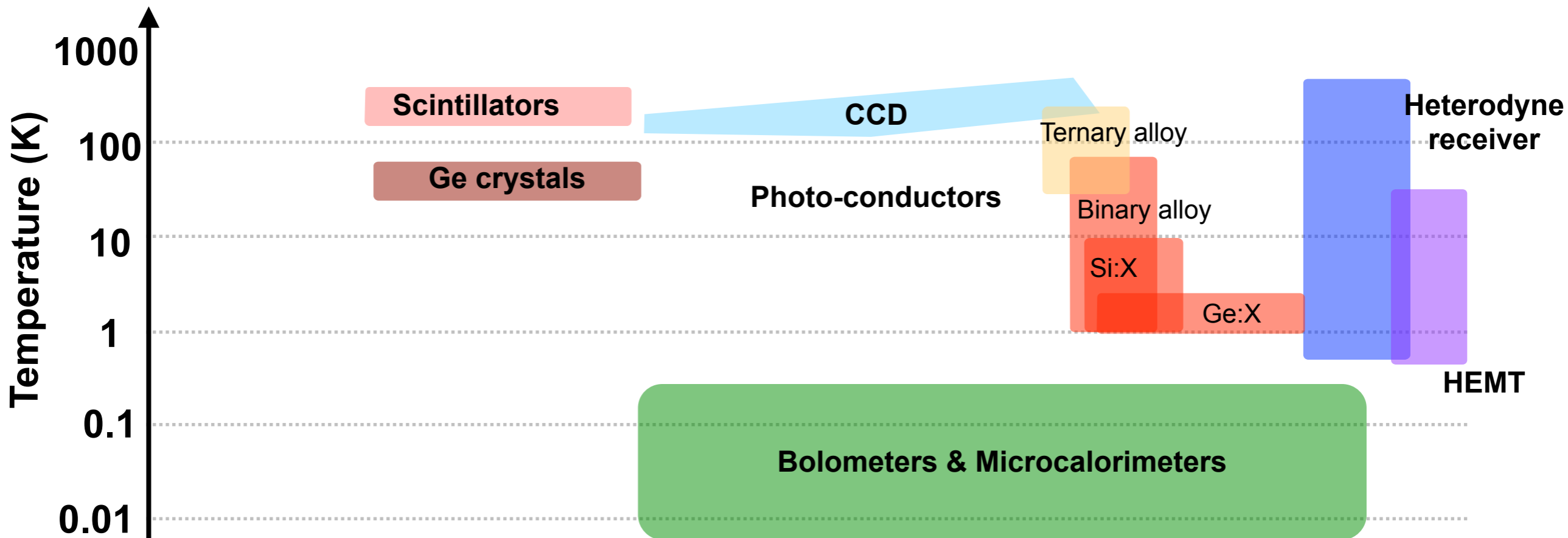
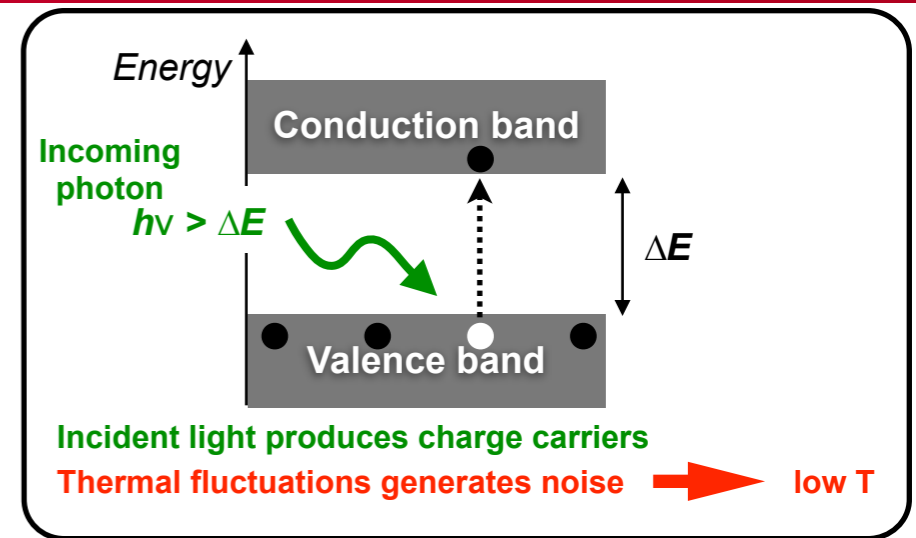
MAIN MOTIVATION: COOLING OF DETECTORS

	Photon detectors	Thermal detectors
Incoming radiation	Interaction with electrons	Temperature change



wavelength dependent
 $E \rightarrow \Delta I, \Delta V, \Delta R$

\approx wavelength independent
 $E \rightarrow \Delta T$

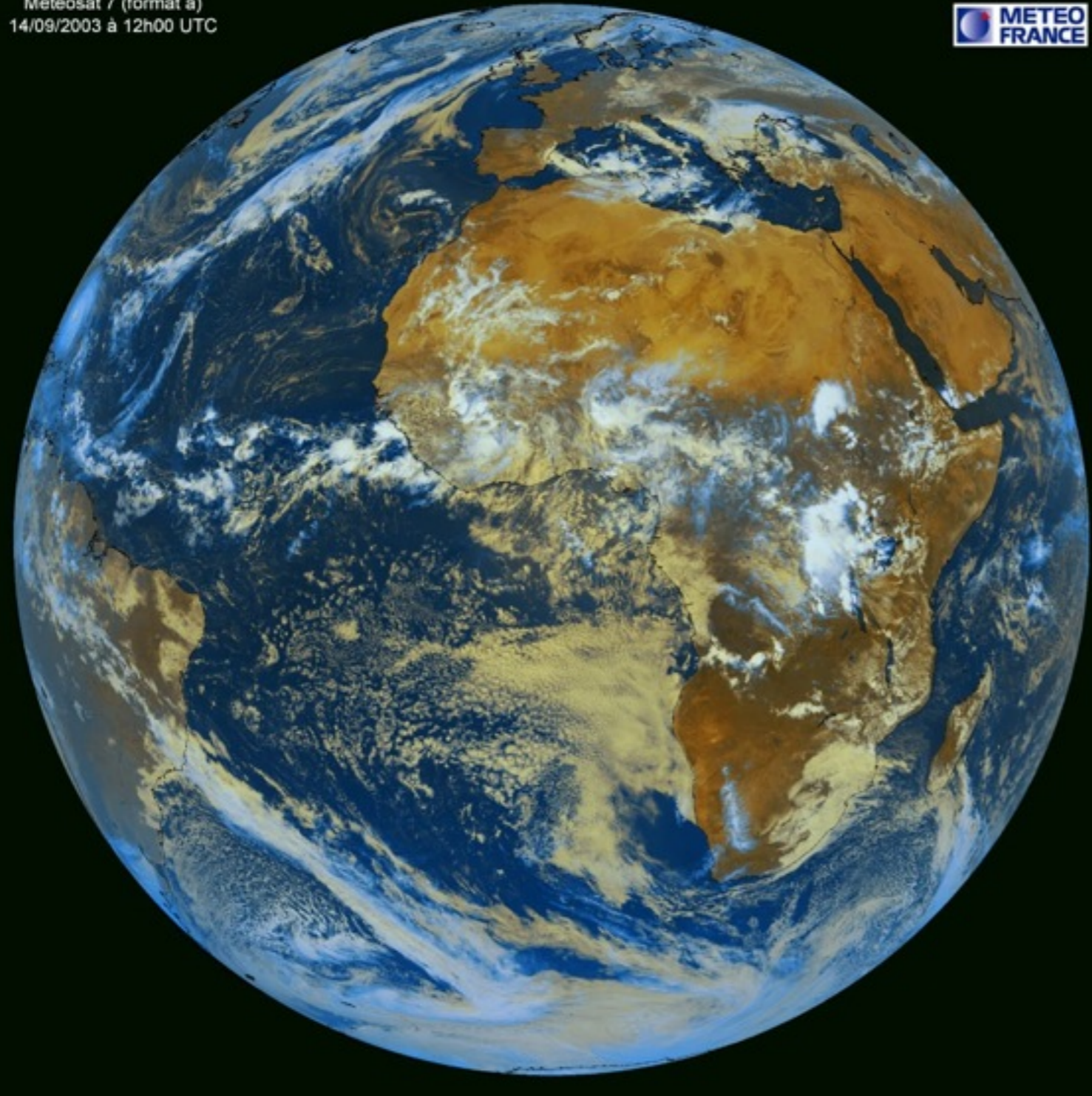


*Infrared instruments:
largest application of space cryogenics*



EARTH OBSERVATION: WEATHER SATELLITE

Météosat 7 (format a)
14/09/2003 à 12h00 UTC

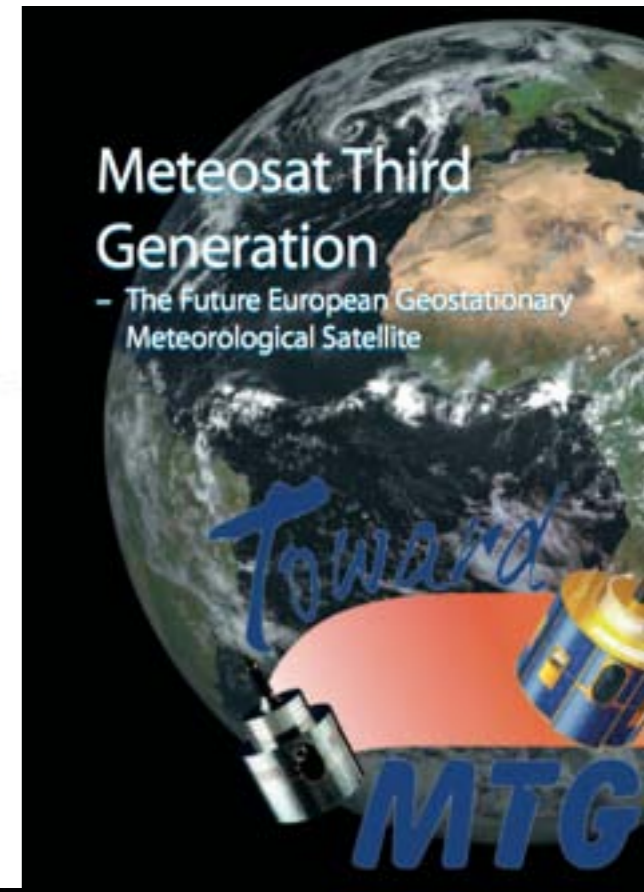


Meteosat 2G

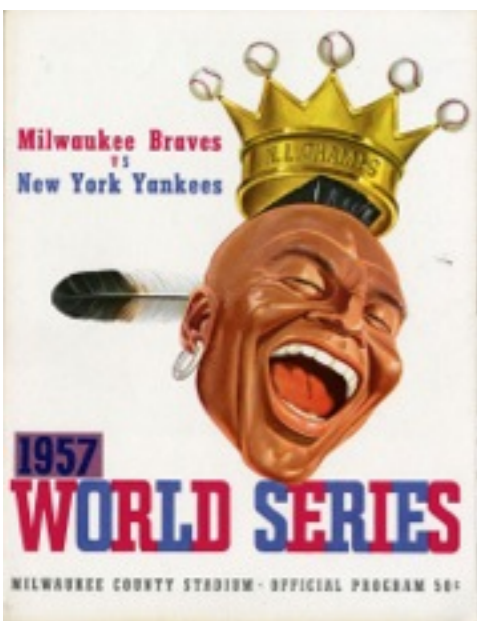


Meteosat Third Generation

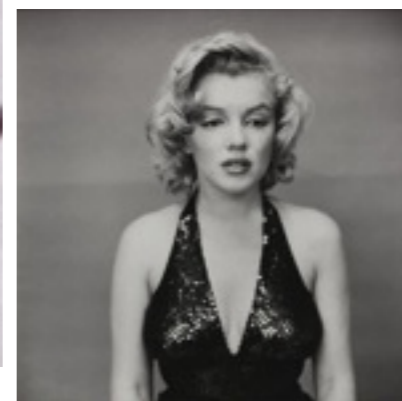
- The Future European Geostationary Meteorological Satellite



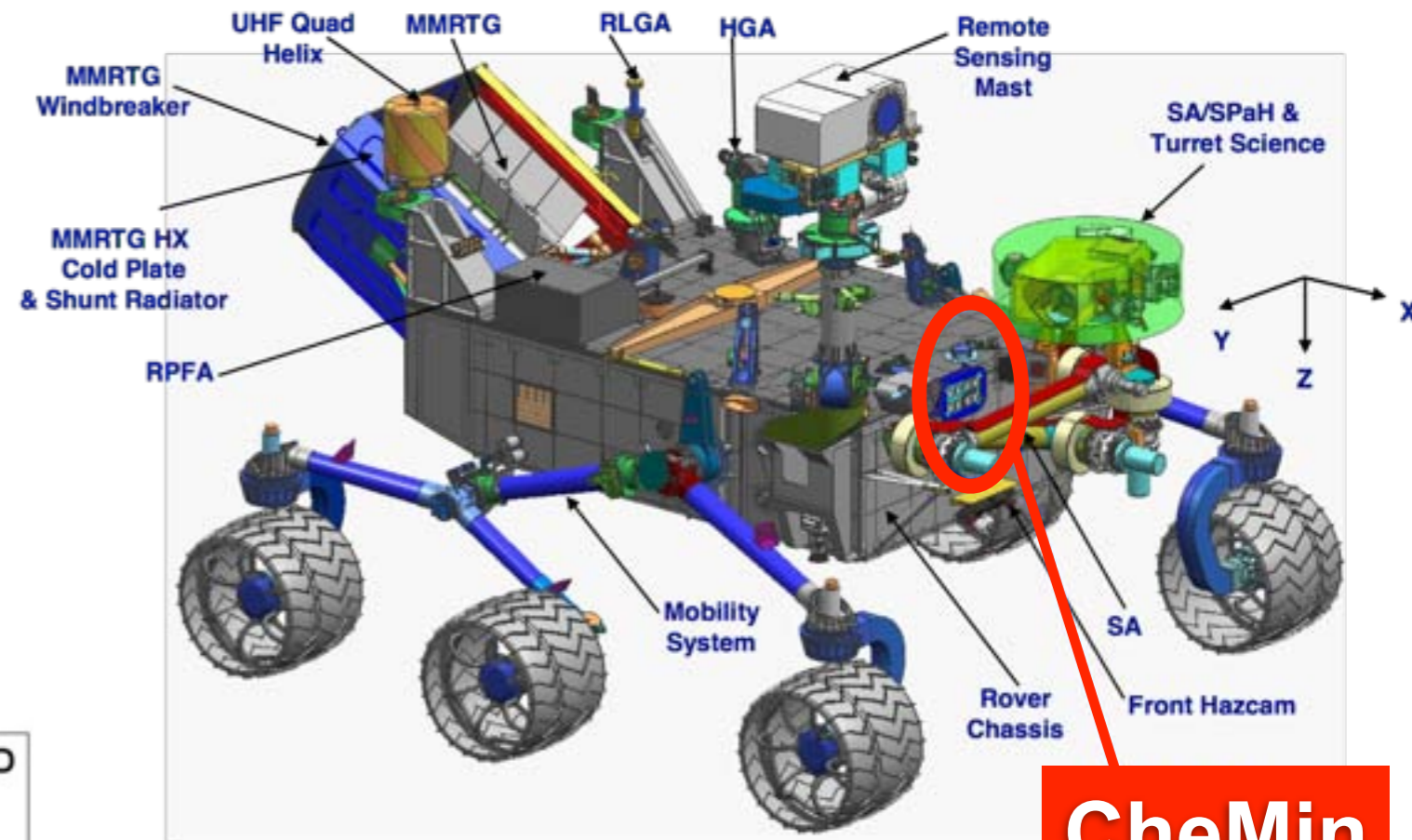
WHEN IT ALL STARTED



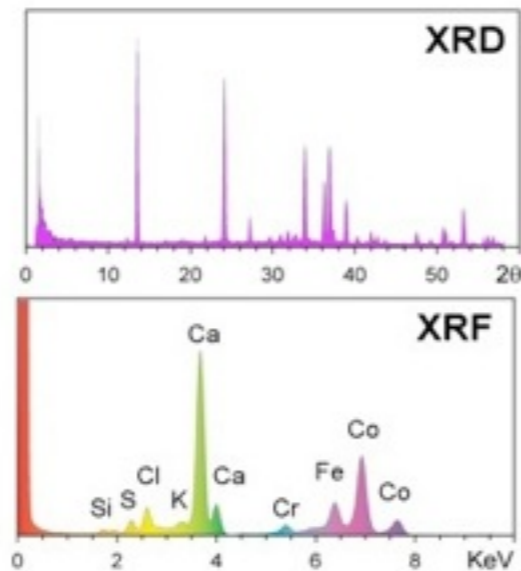
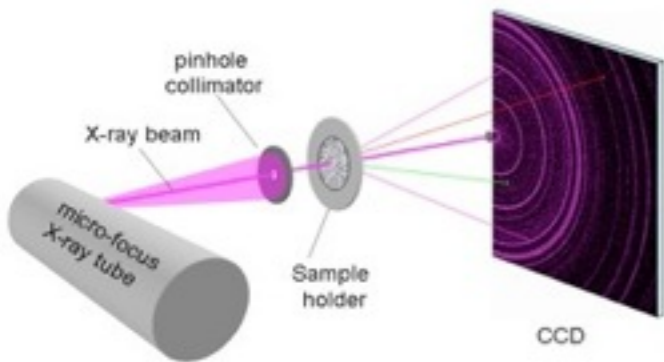
'57 CHEVROLET! SWEET, SMOOTH AND SASSY!



PLANETARY EXPLORATION



CheMin

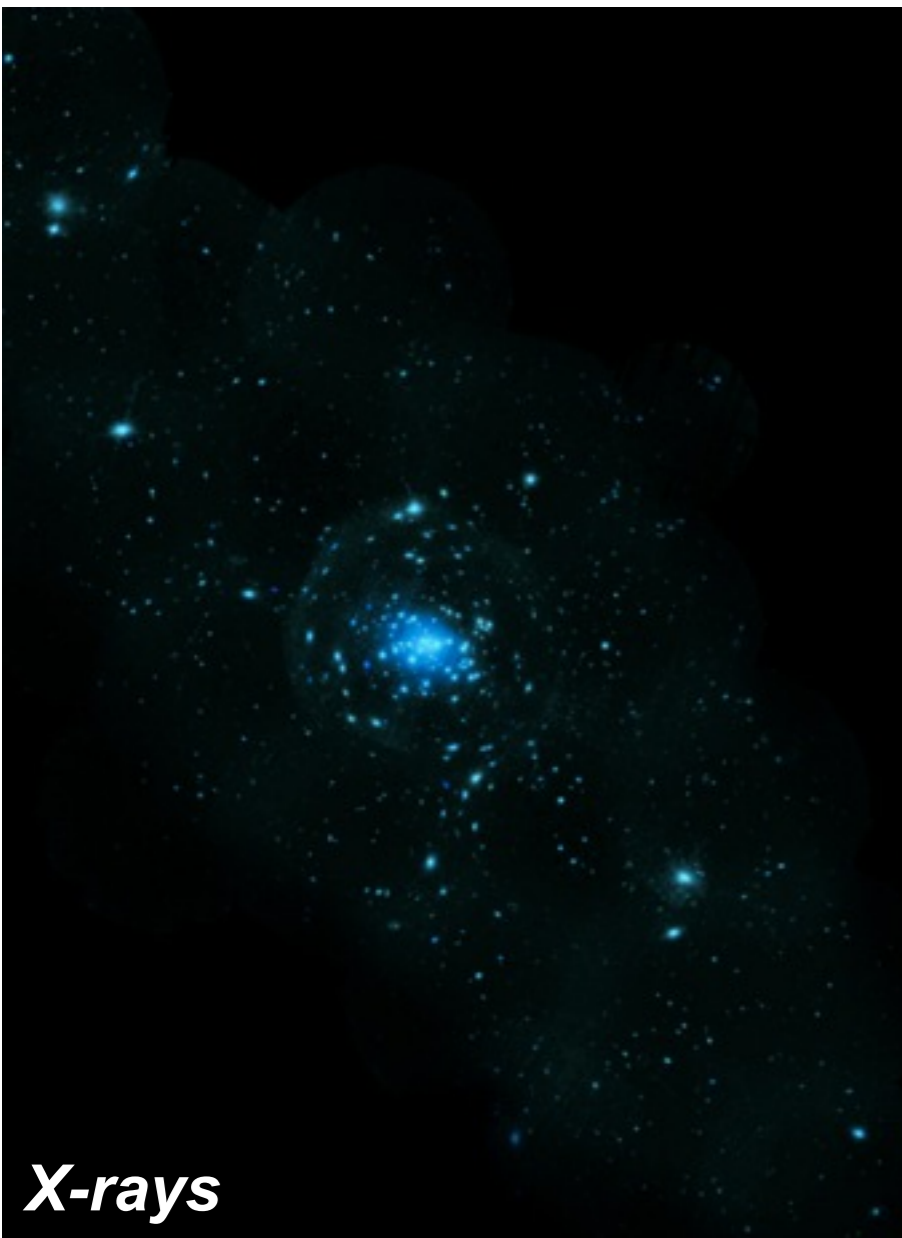


diffraction and fluorescence informations collected by a cooled CCD

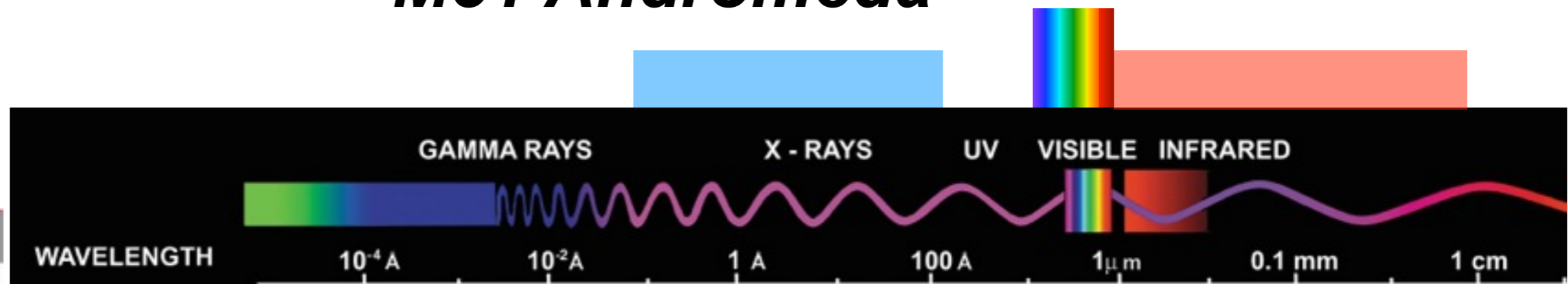
Chemistry and Mineralogy (CheMin): CheMin is a mineralogy instrument, onboard MSL, that identifies and quantifies the minerals present in rocks and soil delivered to it by the Sample Acquisition, Sample Processing and Handling (SA/SPaH) system. The Ricor K508 rotary cooler provides cooling to CCD at ~210K with a lifetime requirement of 1600hrs for surface operations.



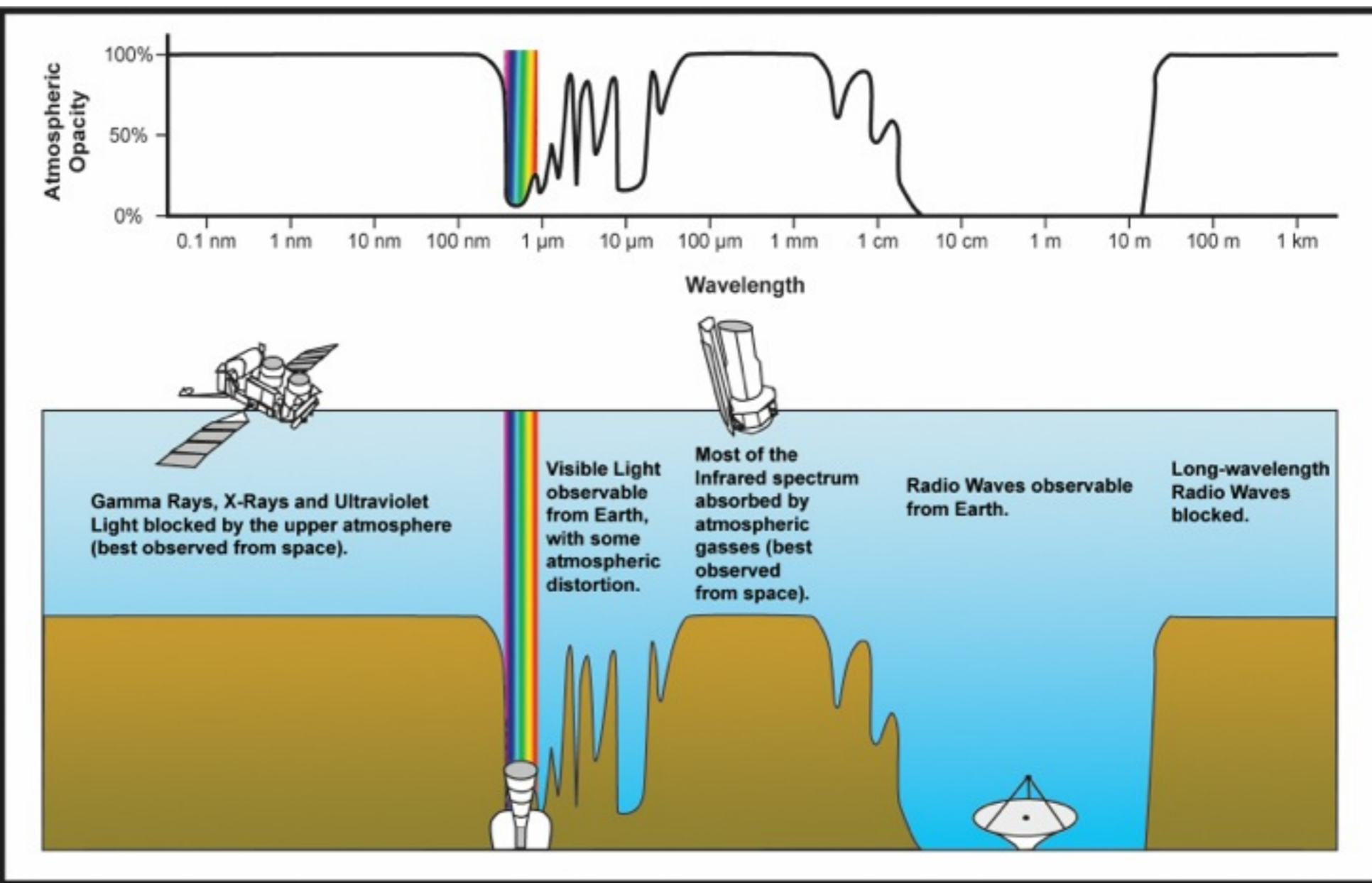
SCIENTIFIC MISSIONS



M31 Andromeda

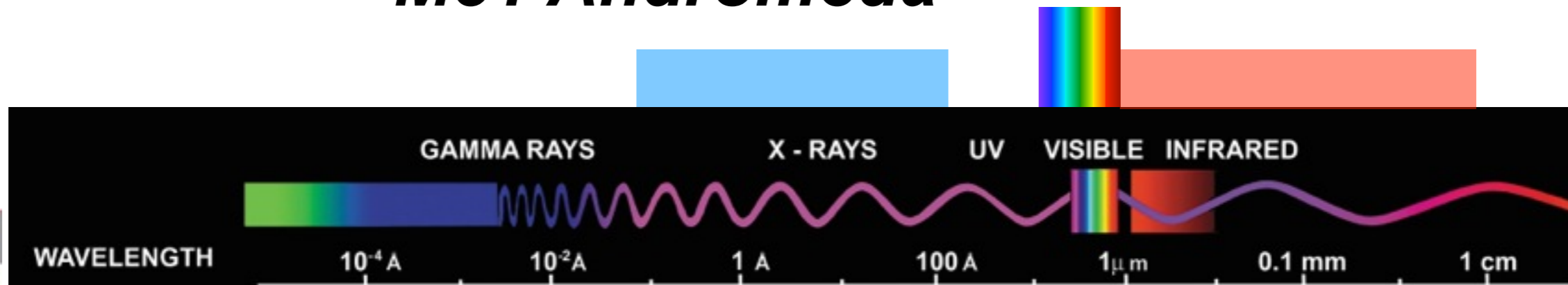


SCIENTIFIC MISSIONS



X-rays

M31 Andromeda

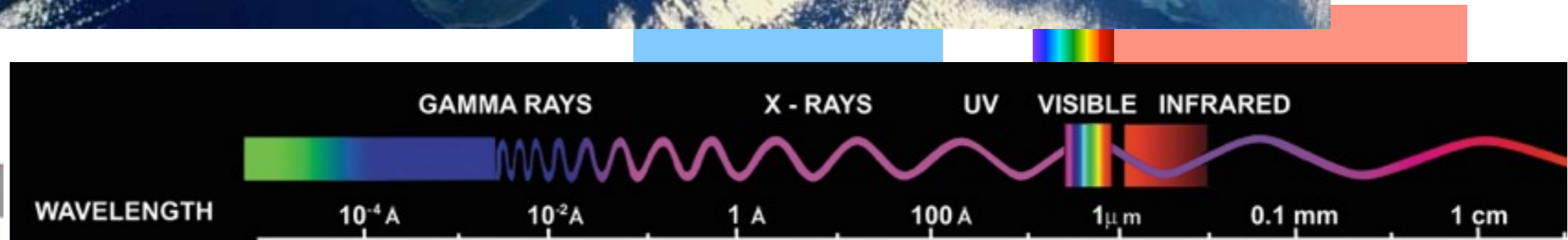


SCIENTIFIC MISSIONS

Access to extended spectrum

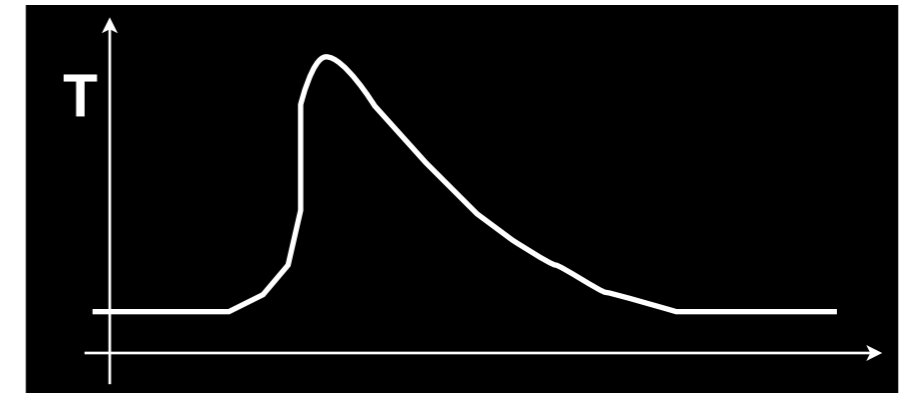
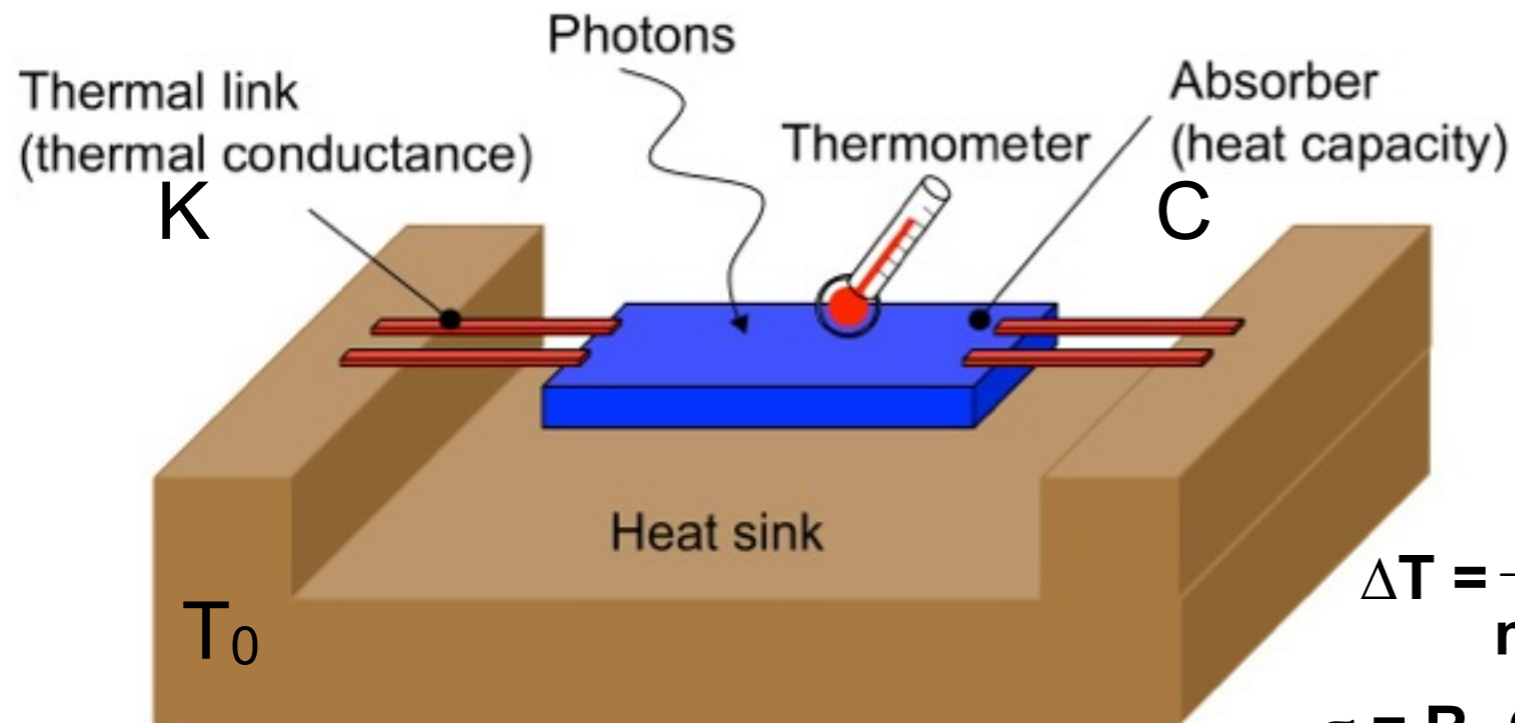


X-rays



How to measure those very faint signals ?

Thermal detector: Bolometer



$$\Delta T = \frac{E}{mC}$$

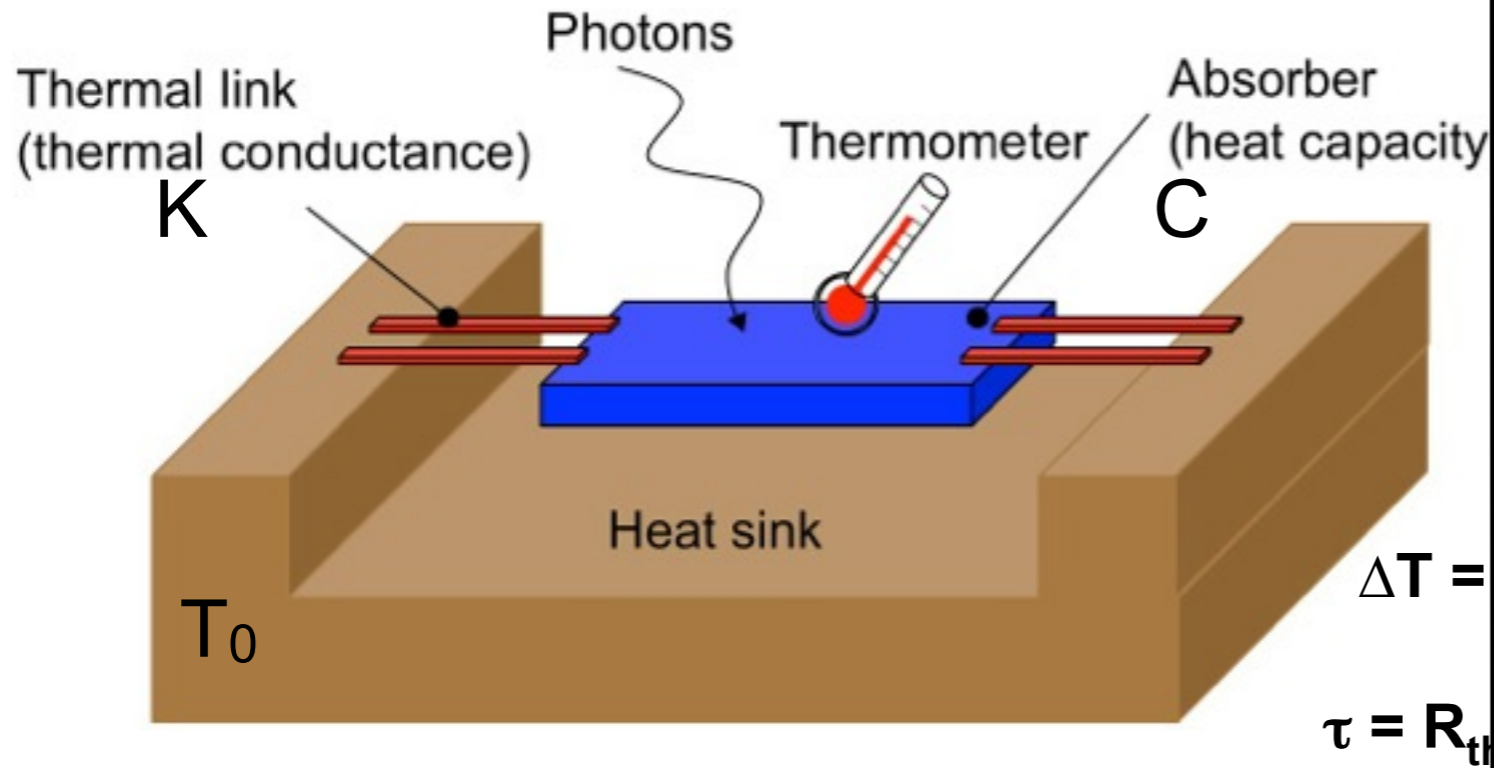
$$\tau = R_{th} C = \frac{C}{K}$$

Measurable ΔT ? \Rightarrow Minimize C \Rightarrow Low Temperature

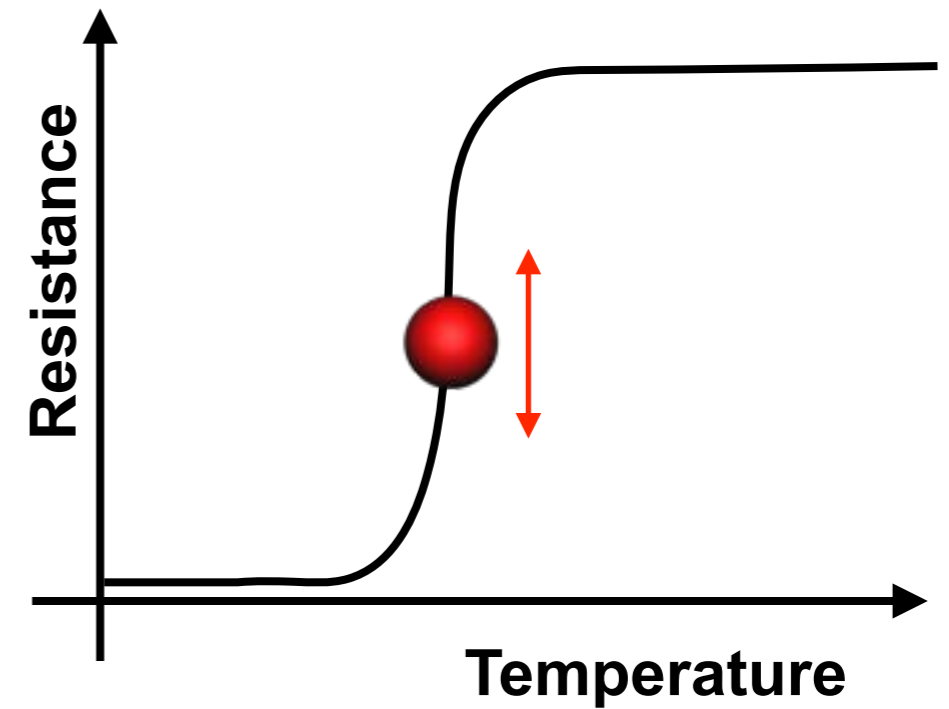
SCIENTIFIC MISSIONS - WHY CRYOGENIC ?

How to measure those very faint signals ?

Thermal detector: Bolometer



Transition Edge Sensor (TES)



small change in T

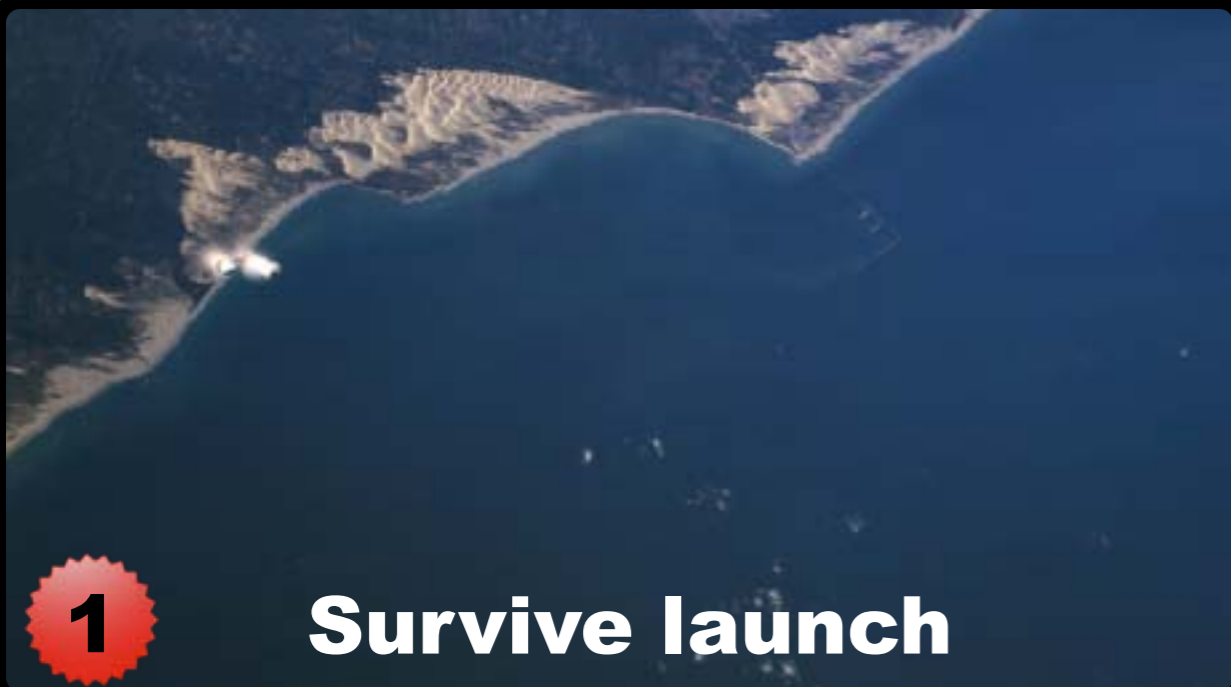


big change in R


Measurable ΔT ? Minimize

SPACE ENVIRONMENT: CONSTRAINTS

1 Survive launch

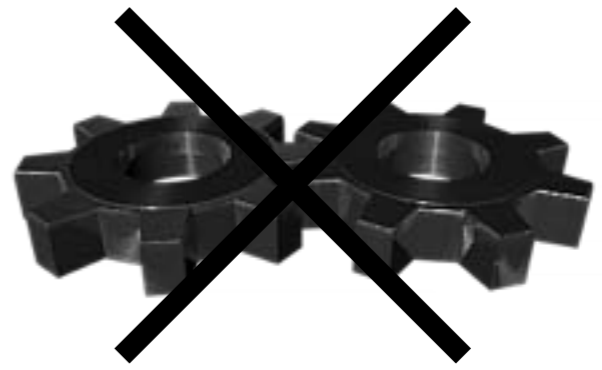


4 Reliability 1500 000 km 

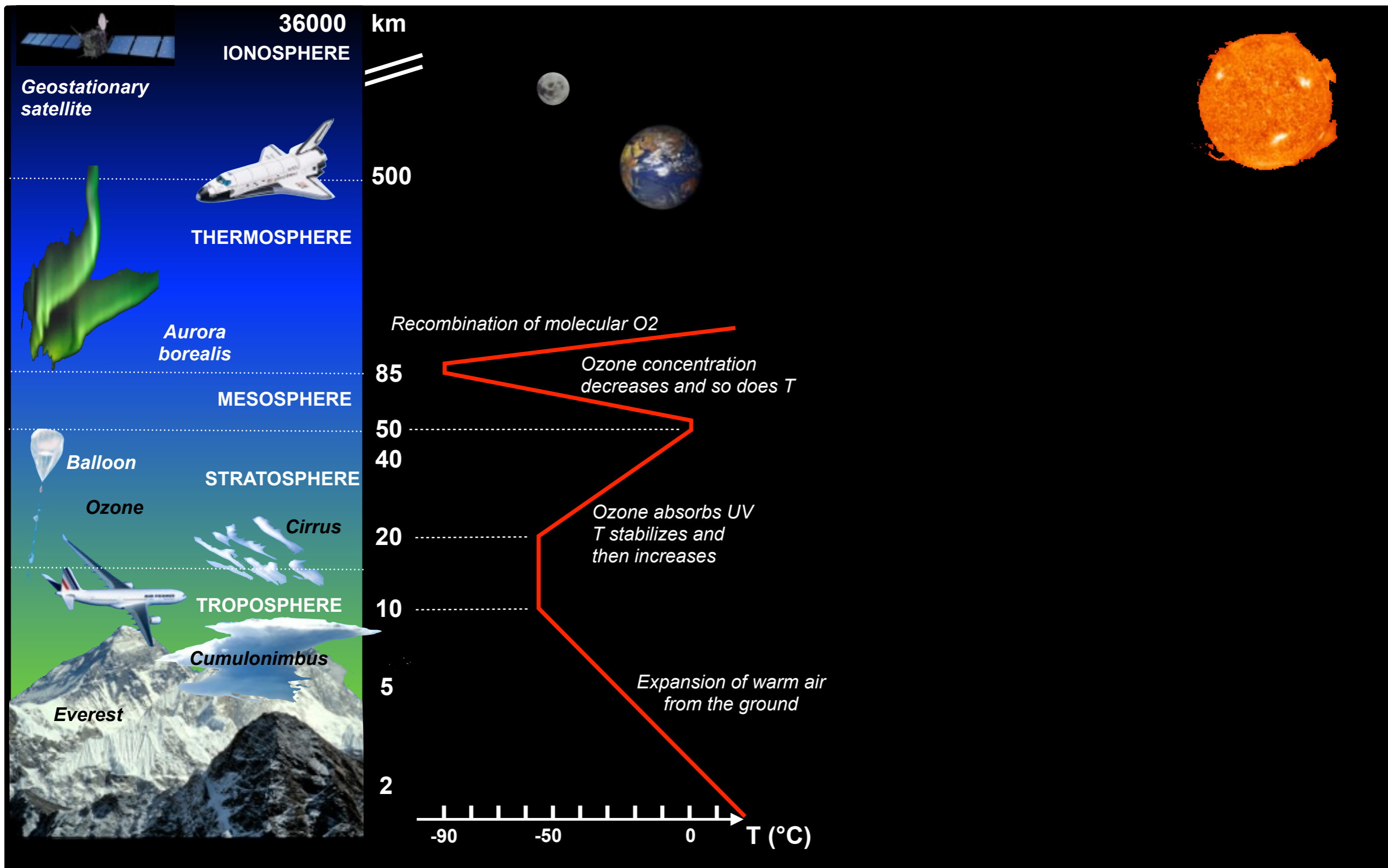
5 Lifetime (> 5 ans) 

2 Mass, power to be minimized 

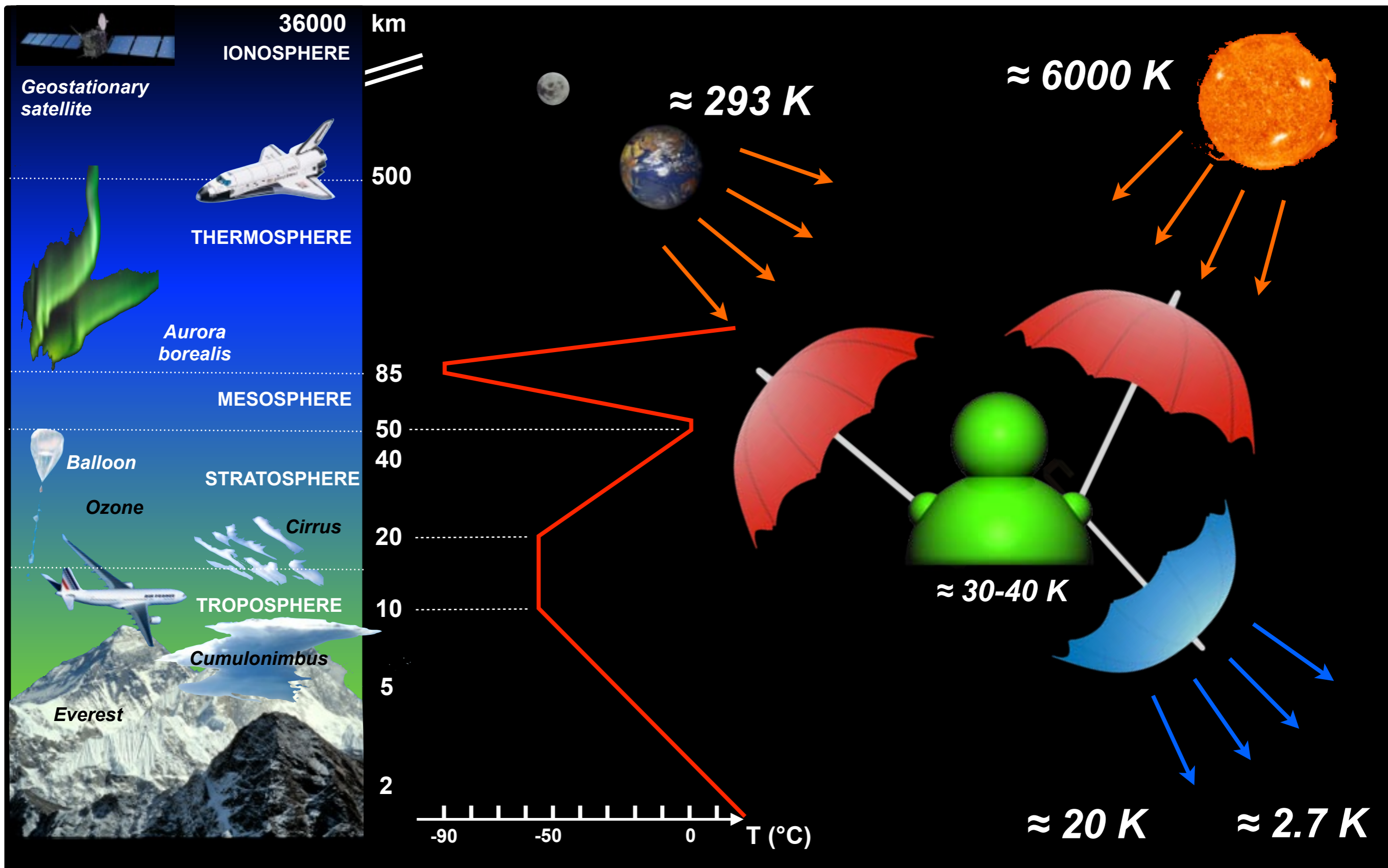
3 Operation in free fall 


No moving parts or absence of any friction

IS IT COLD OUT THERE ?



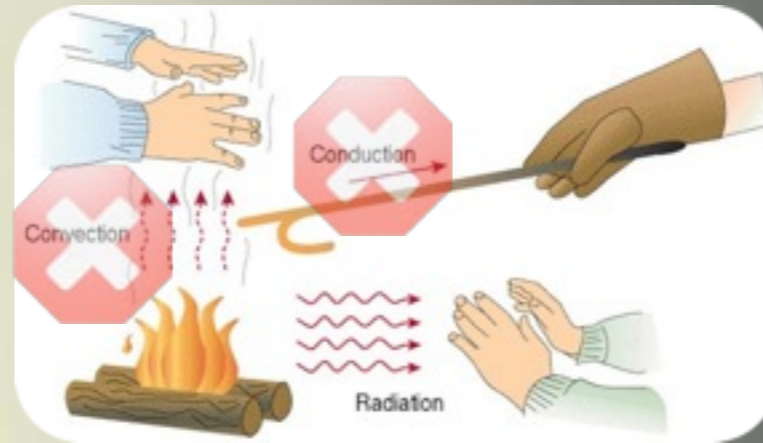
IS IT COLD OUT THERE ?



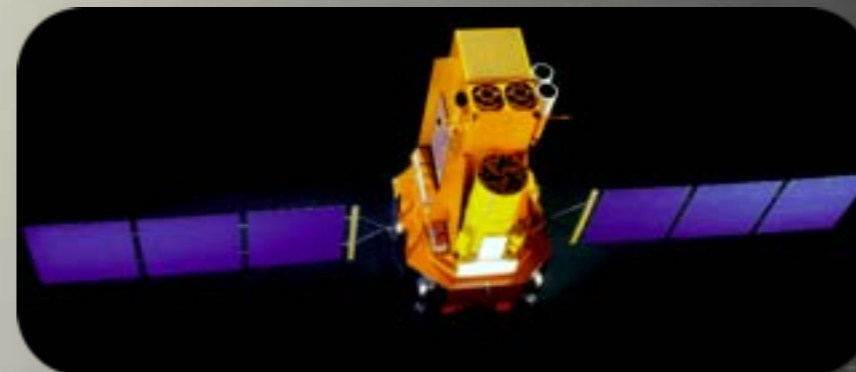
SPACE: NATURAL RESOURCES



≈ High Vacuum



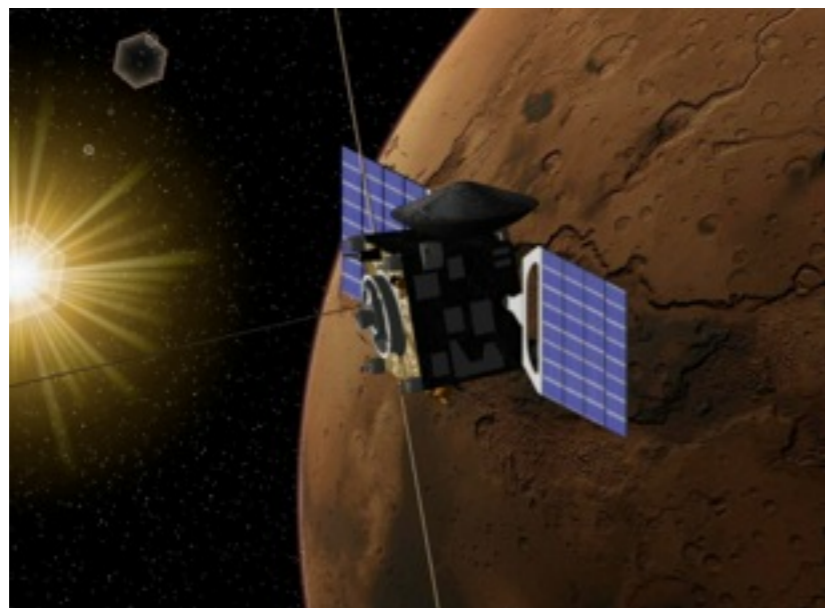
Radiative heat transfert
(deep space @ 2.7 K)



Solar photons

SPACE COOLING CHAINS

0.01 K 0.1 K 1 K 10 K 100 K 300 K



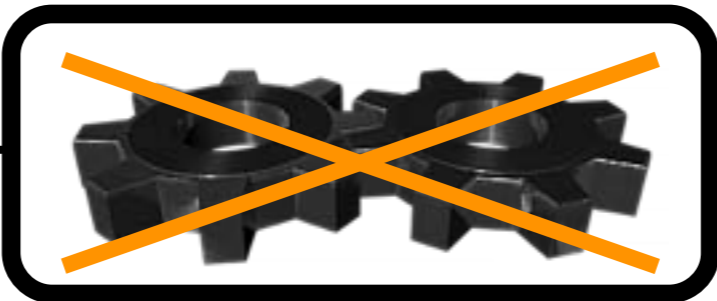
Passive Radiators

Stored cryogen (Cryostat)

Active "mechanical" coolers

"Ultra" low T coolers
(3 technologies: sorption, magnetic, dilution)

Reliability



No moving parts or absence of friction

SPACE CRYOGENICS - CURRENT GENERAL TENDENCY



Extended mission
Cryogenics Coolers

Higher and farther
Engines & μ Gravity
Zero boil off



Mass
Thermal links
Stability & T control

Colder
50 mK



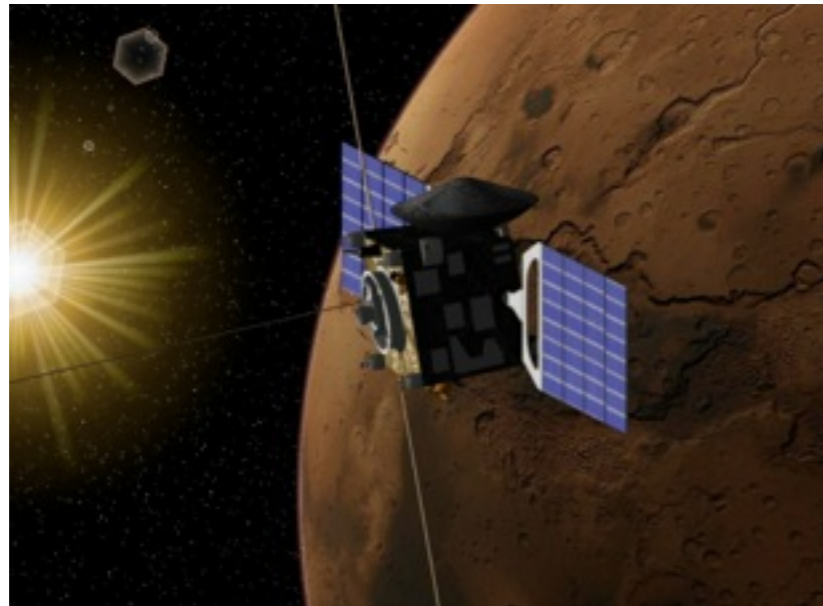
I CANNOT BE EXHAUSTIVE - SORRY



SPACE CRYOGENICS

SPACE COOLING CHAINS

0.01 K 0.1 K 1 K 10 K 100 K 300 K



Passive Radiators

Stored cryogen (Cryostat)

Active "mechanical" coolers

"Ultra" low T coolers
 (3 technologies: sorption, magnetic, dilution)



- * Efficient
- * Simple
- * Reliable
- * Vibration free



- * Limited performance @ low T
- * Orbit / orientation dependent

RADIATIVE COOLING

Direct coupling to deep space via thermal self emission

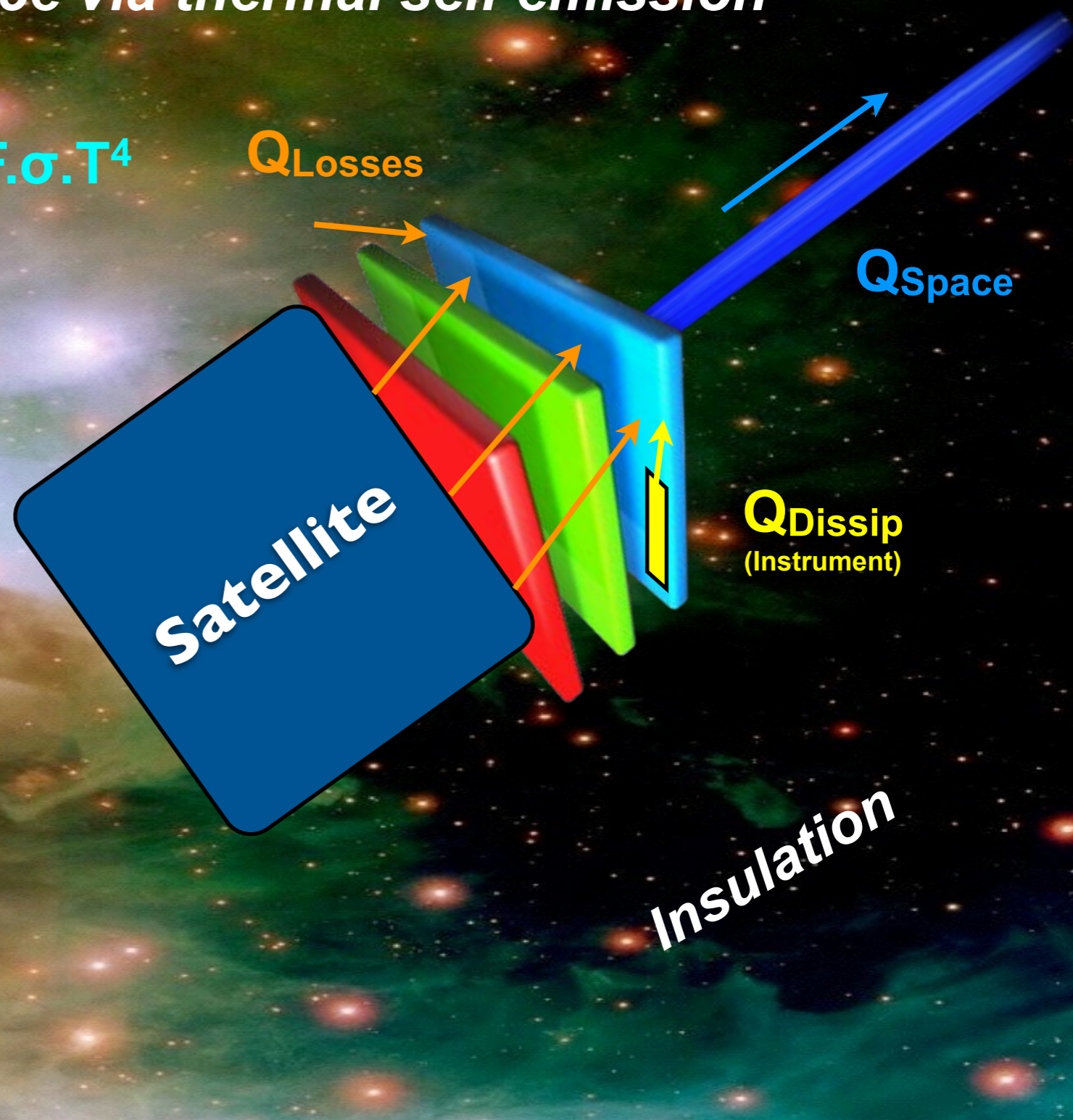
$$Q_{\text{Space}} = Q_{\text{Dissip}} + Q_{\text{Losses}}$$

$$Q_{\text{Space}} = \varepsilon \cdot A \cdot F \cdot \sigma \cdot (T^4 - T_s^4) \approx \varepsilon \cdot A \cdot F \cdot \sigma \cdot T^4$$

$$\sigma = 5.67 \cdot 10^{-8} \text{ W} \cdot \text{m}^{-2} \cdot \text{K}^{-4}$$

$F \approx 1$ (form factor)

$\varepsilon \approx 0.9$ to 1 (black paint, open honey comb)



RADIATIVE COOLING

Direct coupling to deep space via thermal self emission

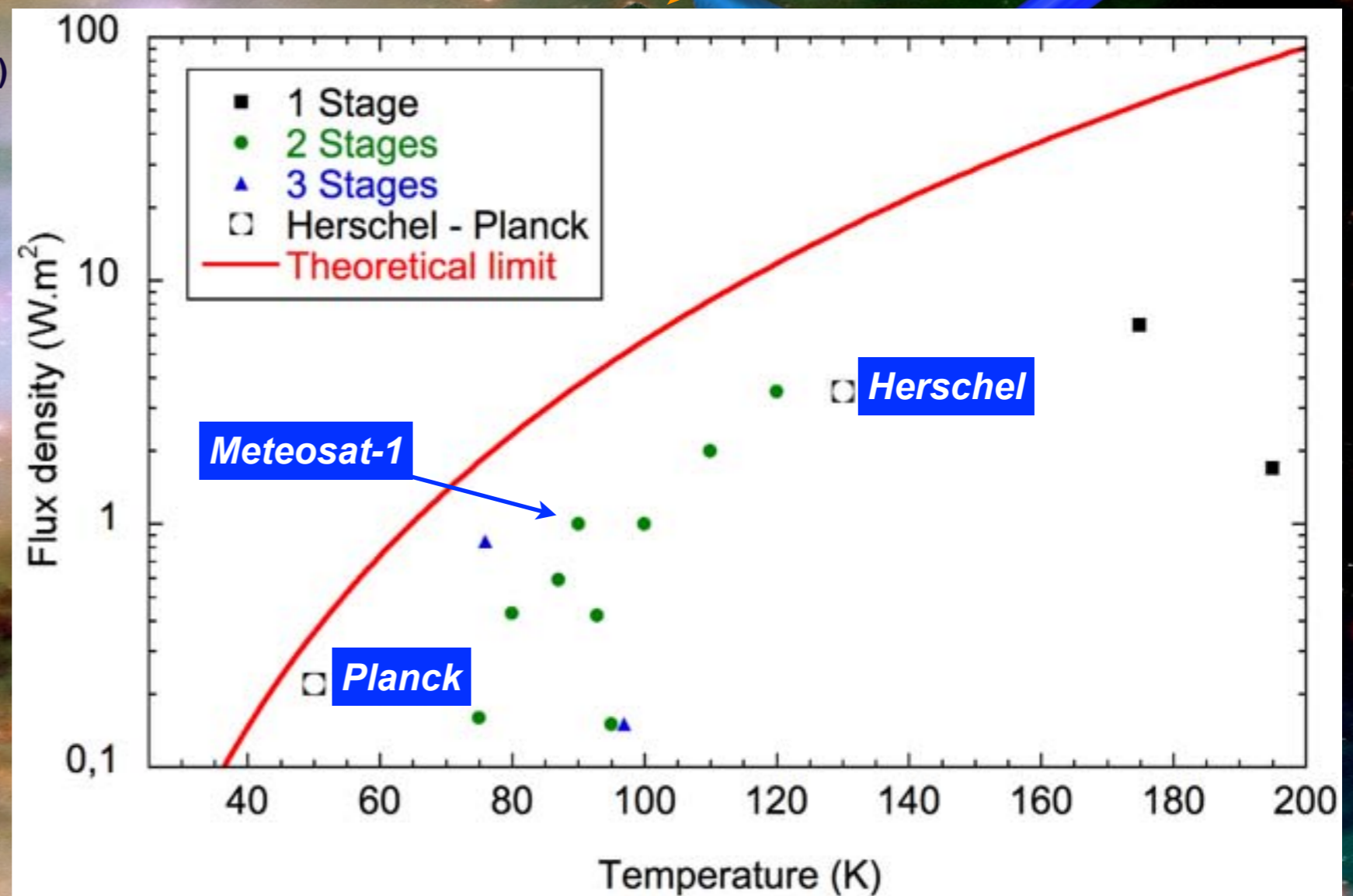
$$Q_{\text{Space}} = Q_{\text{Dissip}} + Q_{\text{Losses}}$$

$$Q_{\text{Space}} = \epsilon \cdot A \cdot F \cdot \sigma \cdot (T^4 - T_s^4) \approx \epsilon \cdot A \cdot F \cdot \sigma \cdot T^4 \quad Q_{\text{Losses}}$$

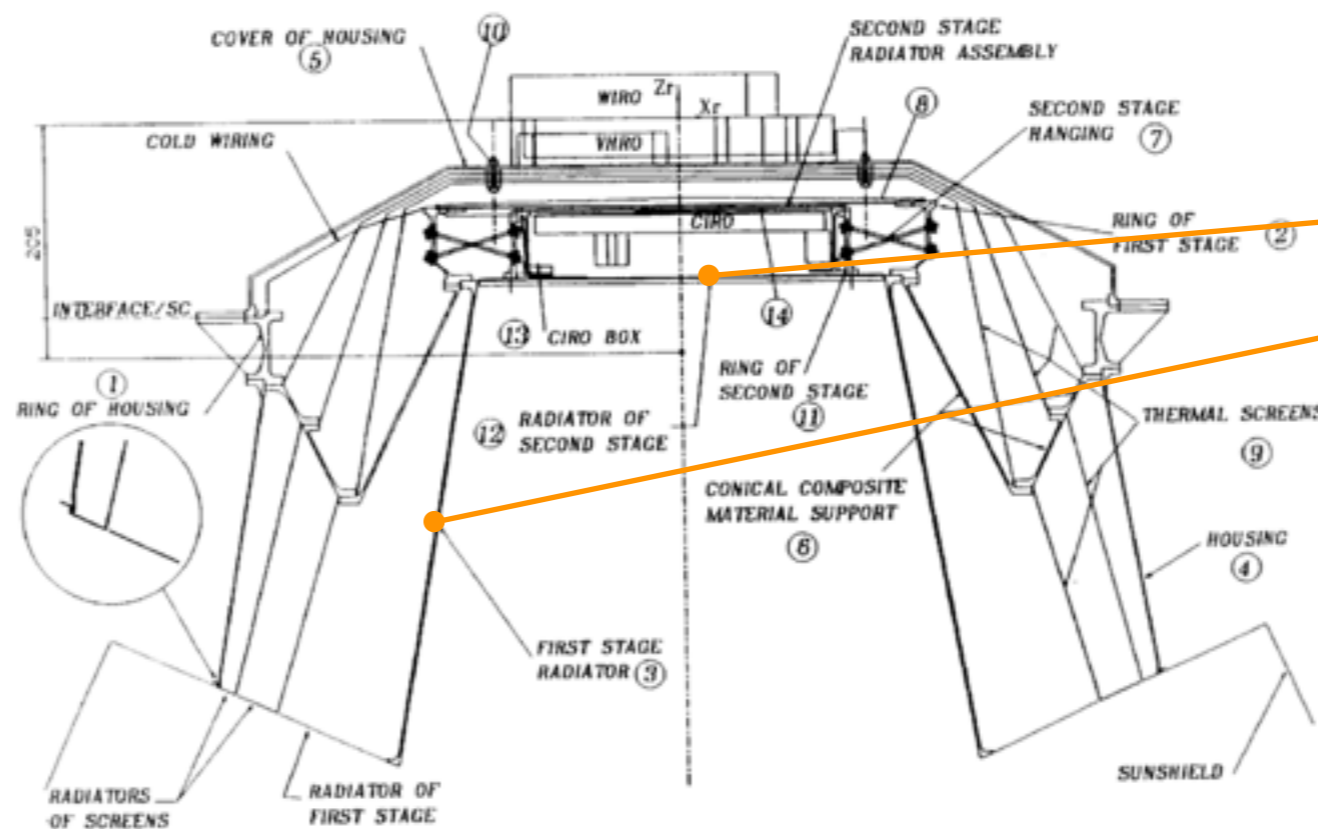
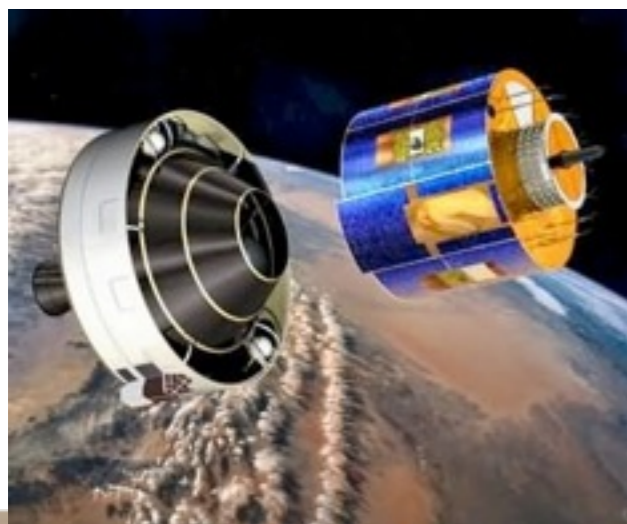
$$\sigma = 5.67 \cdot 10^{-8} \text{ W.m}^{-2}.\text{K}^{-4}$$

$F \approx 1$ (form factor)

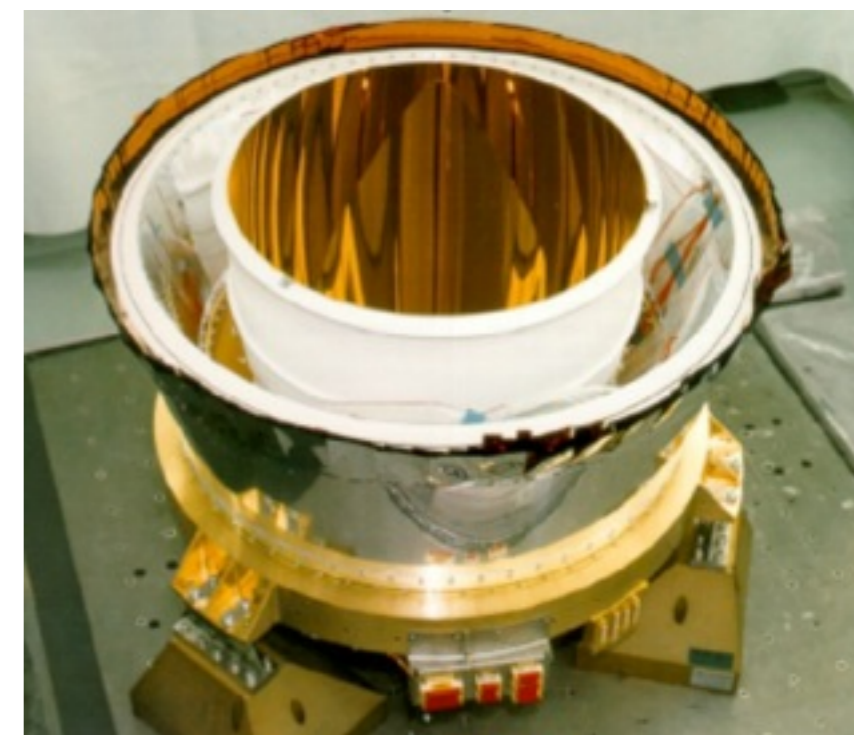
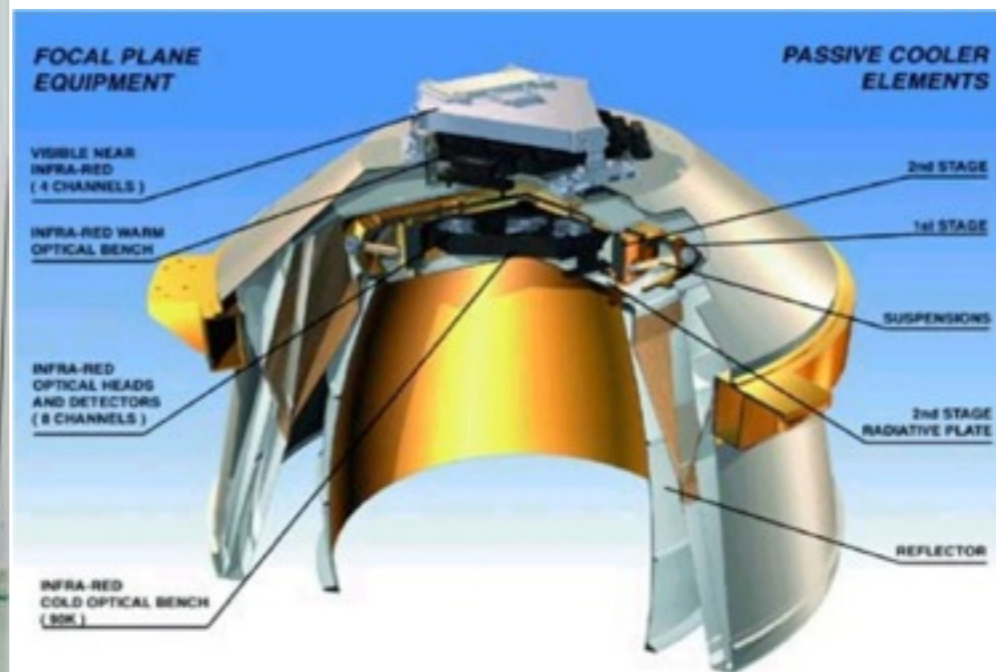
$\epsilon \approx 0.9$ to 1 (black paint, open honey comb)



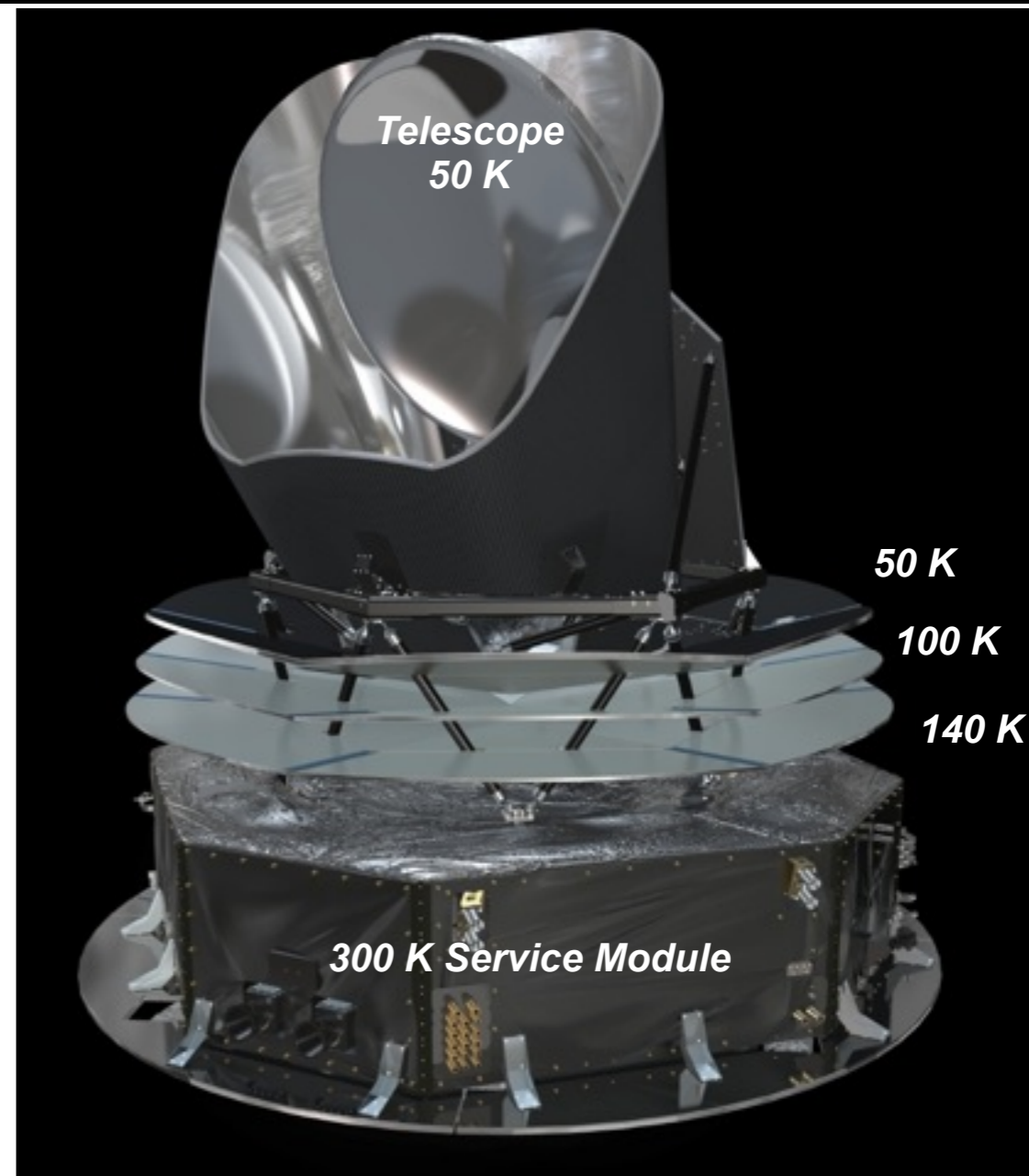
EXAMPLE METEOSAT 2 GENERATION



**Two stages
85 K - 95 K
≈ 120 K**

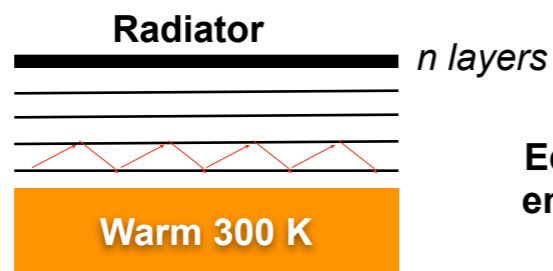


MLI / V-GROOVES - HERSCHEL / PLANCK



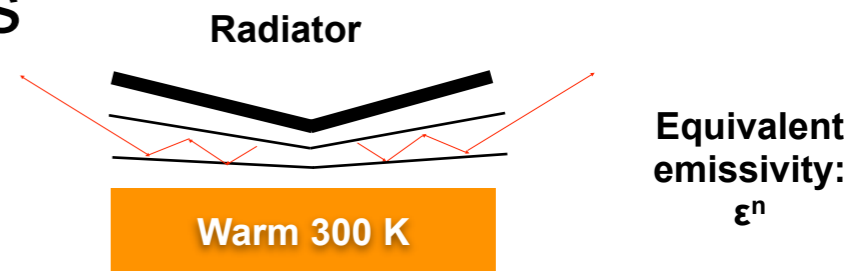
“Classical”

MLI: radiation trapped between layers

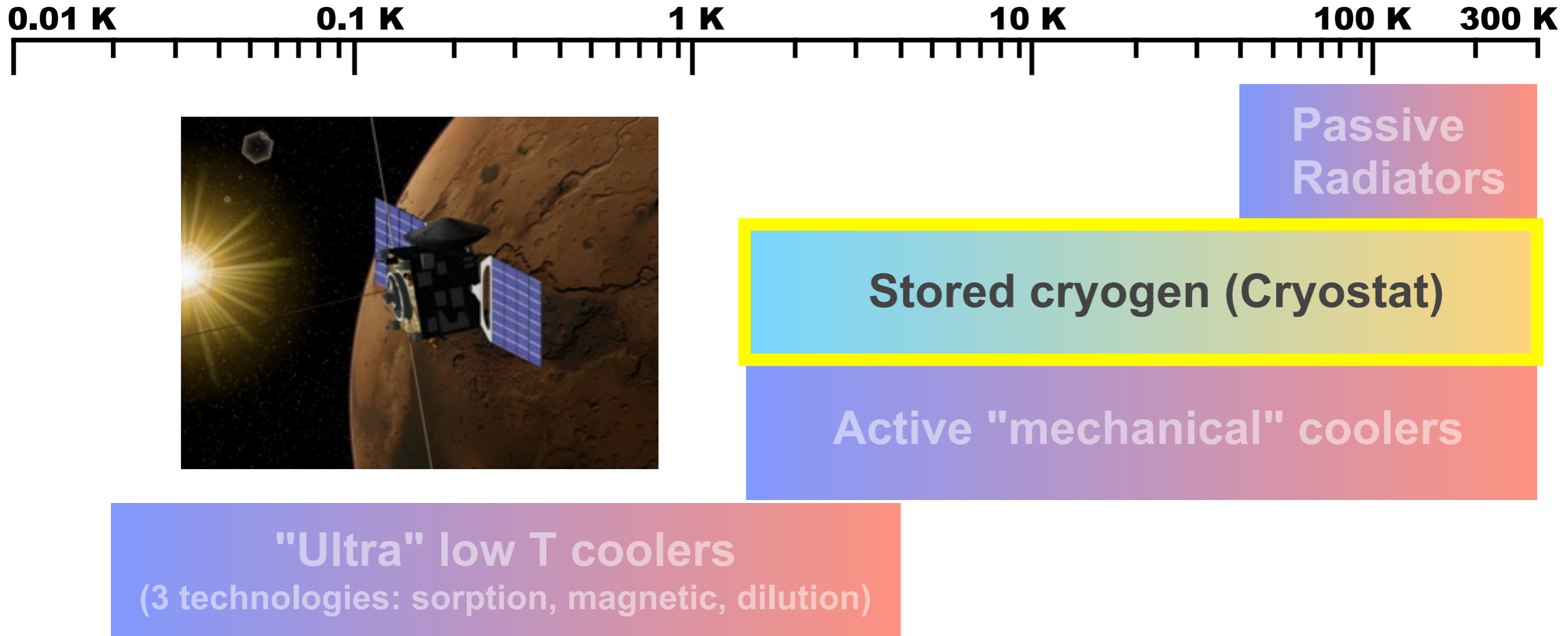


“V-Grooves”

Radiation rejected to space



SPACE COOLING CHAINS



- * Efficient
- * Simple
- * Reliable
- * Cold vapor



- * Limited mission duration
- * only selected T available
- * Volume & Mass
- * On ground management

LOW TEMPERATURE MISSIONS WITH STORED CRYOGENS (4K OR LESS)

GRENOBLE

IRAS

ASTRO-E

SUZAKU

ISO

COBE

SFU-IRTS

SPITZER

HERSCHEL

Gravity Probe B

1983

1989

1995

2000

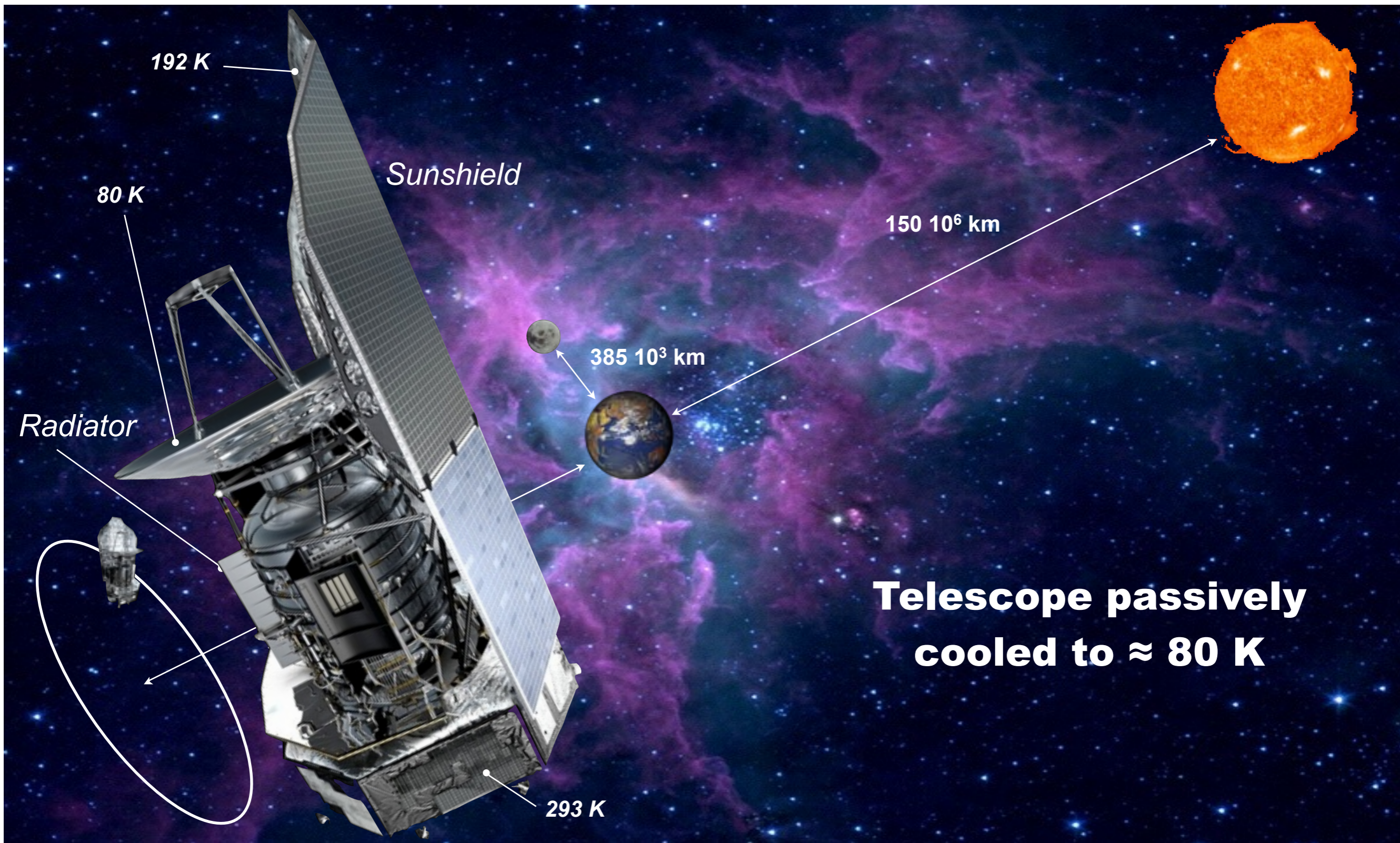
2003

2004

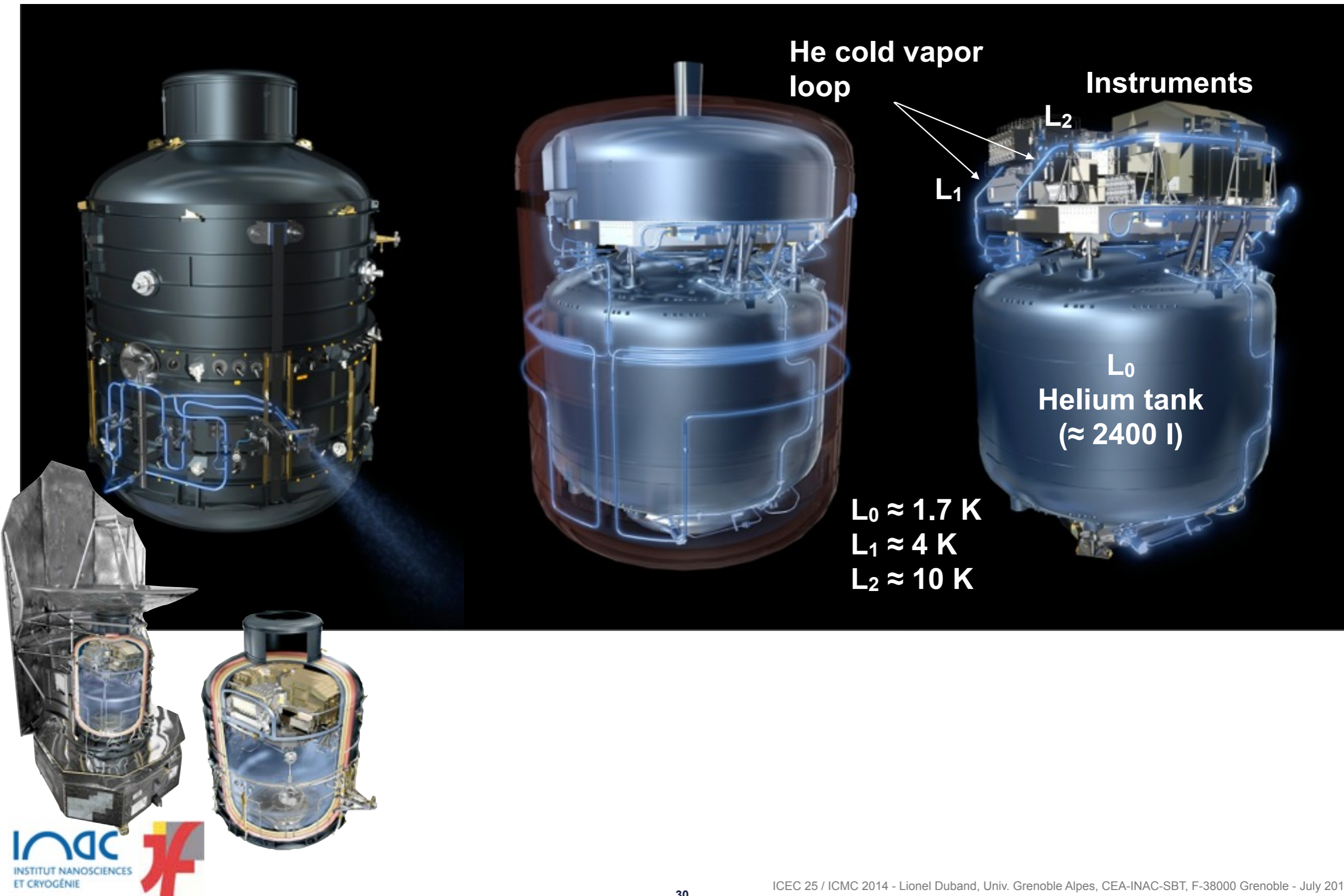
2005

2009

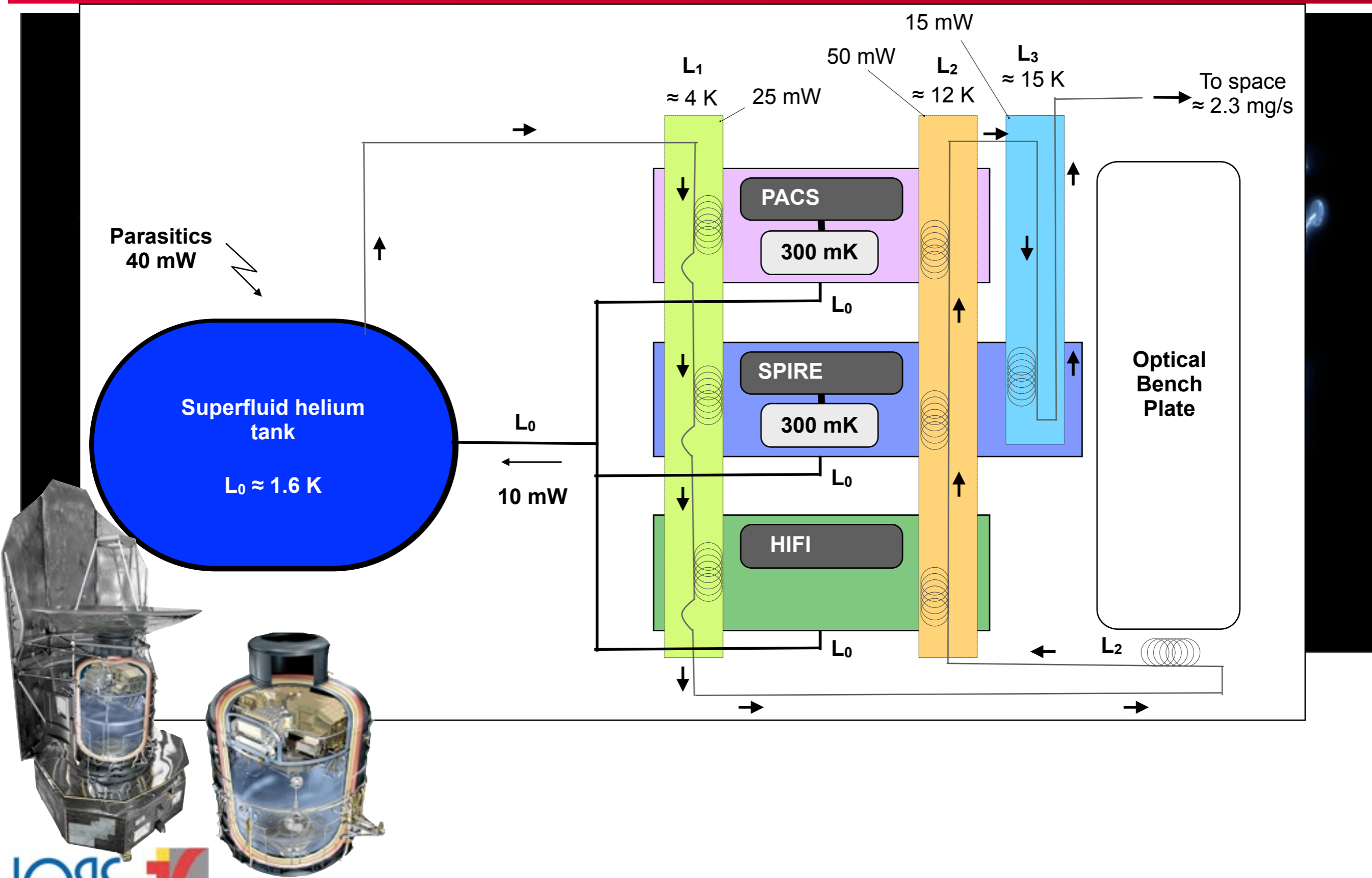
HERSCHEL CRYOGENIC CHAIN - 1



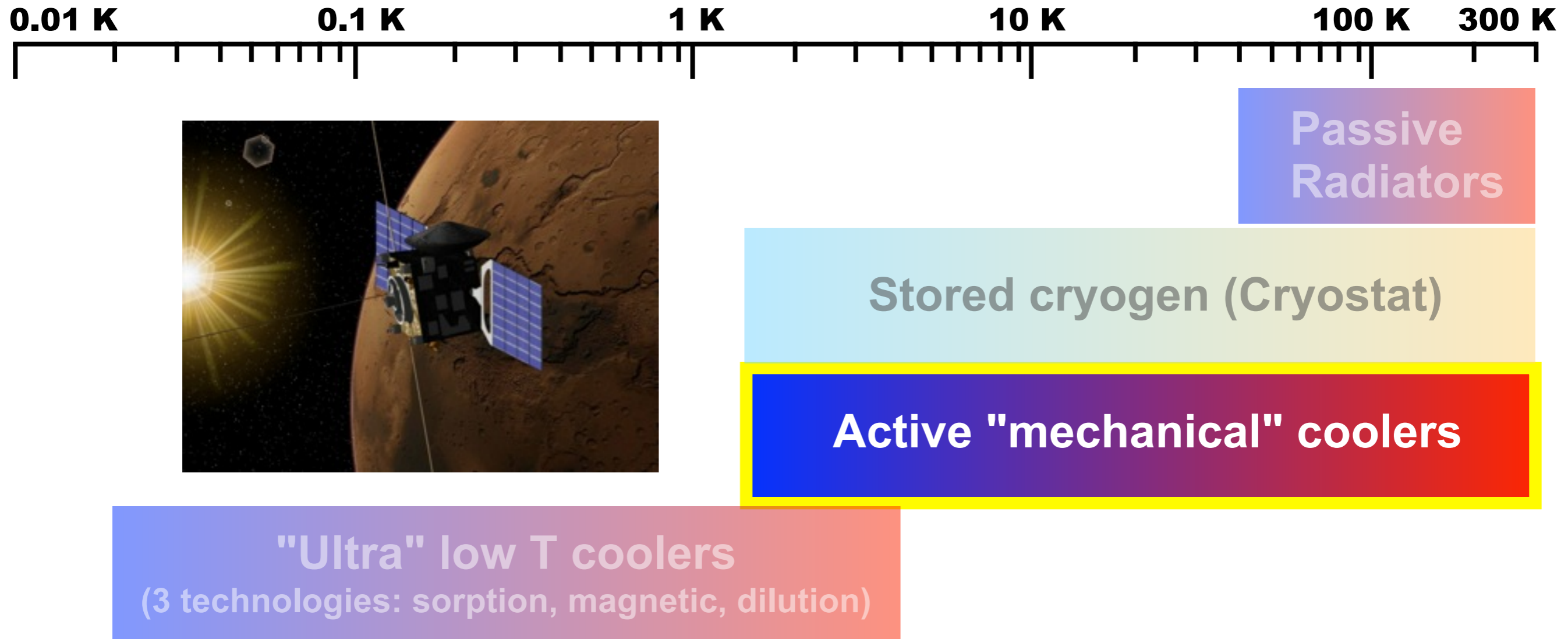
HERSCHEL CRYOGENIC CHAIN - 2



HERSCHEL CRYOGENIC CHAIN - 2



SPACE COOLING CHAINS



- * Lifetime
- * Warm launch
- * Ground tests



- * Peak power
- * Thermal interfaces
- * Vibration

SPACE CRYOGENICS: CURRENT TENDENCY

He cold vapor loop

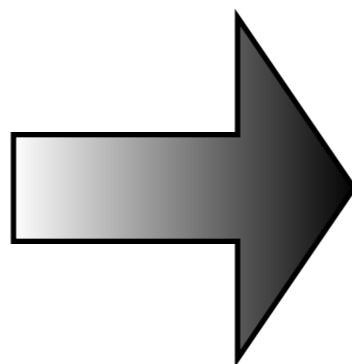
Instruments



Superfluid helium tank



- Thermal interface
- Peak Power
- Distribution (vapor)
- Vibration free
- Passive (once in space)

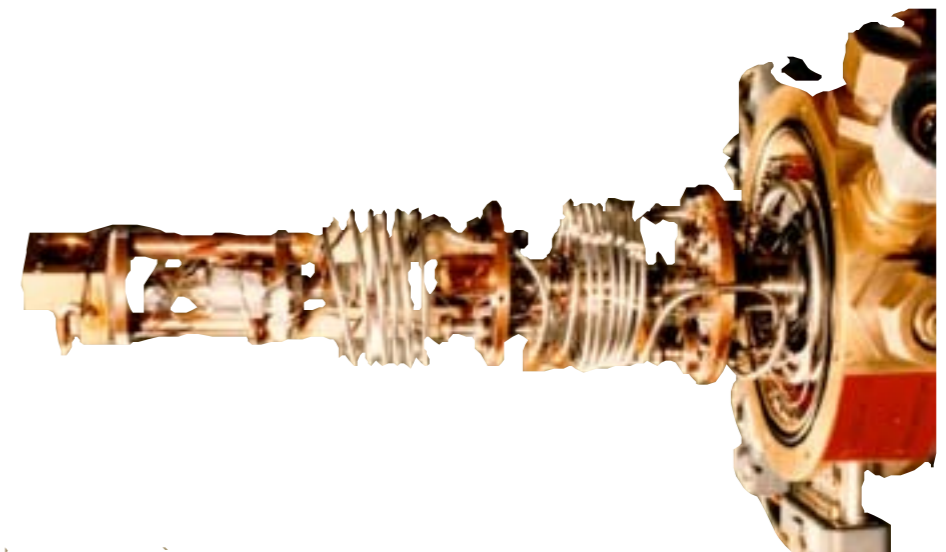
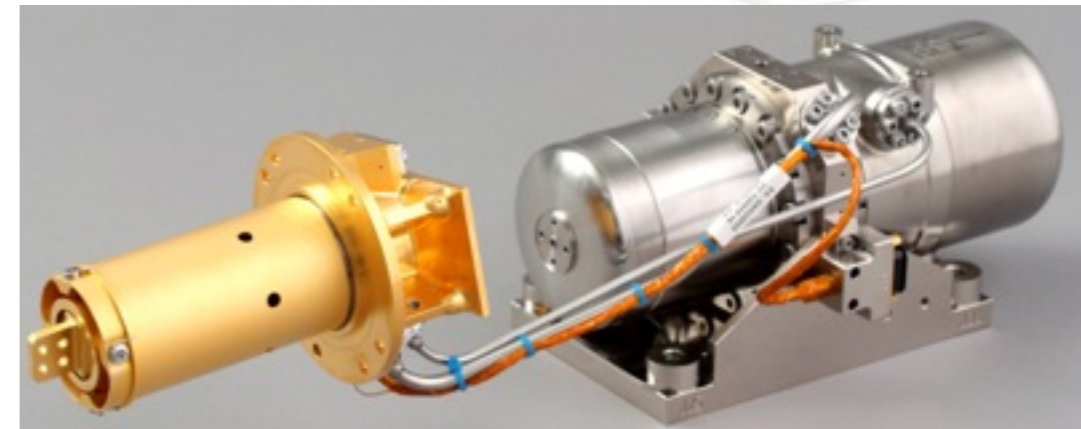
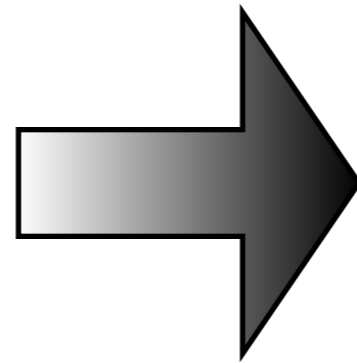


- Limited lifetime
- Volume
- Mass
- On ground management
 - Vacuum shell needed

SPACE CRYOGENICS: CURRENT TENDENCY

He cold vapor
loop

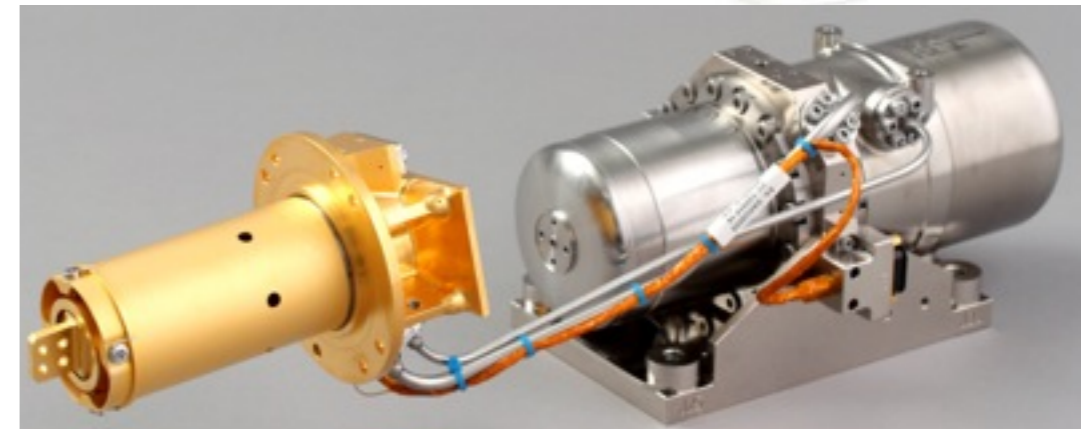
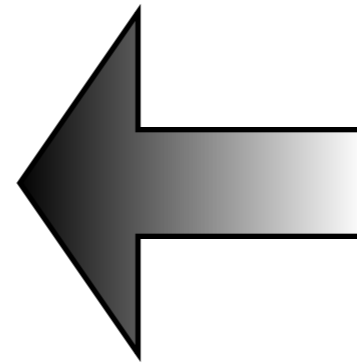
Instruments



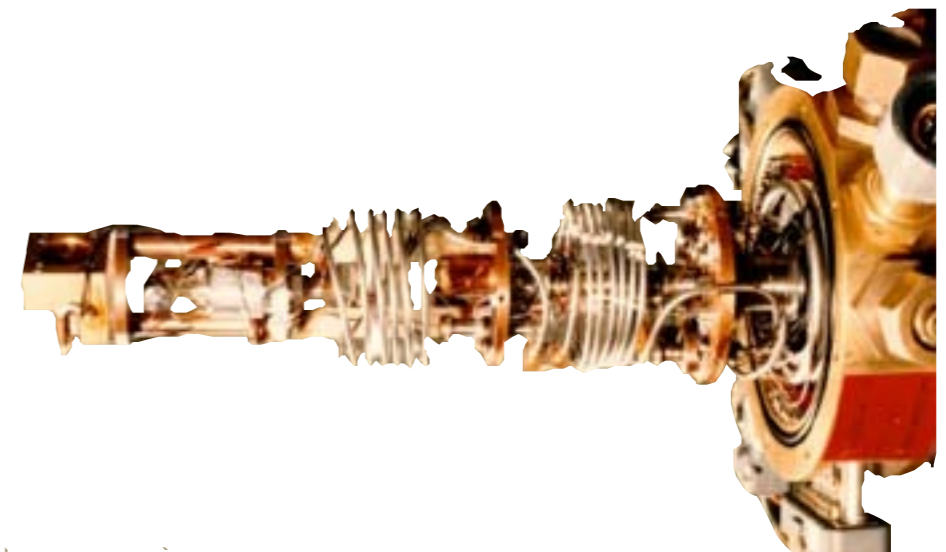
SPACE CRYOGENICS: CURRENT TENDENCY



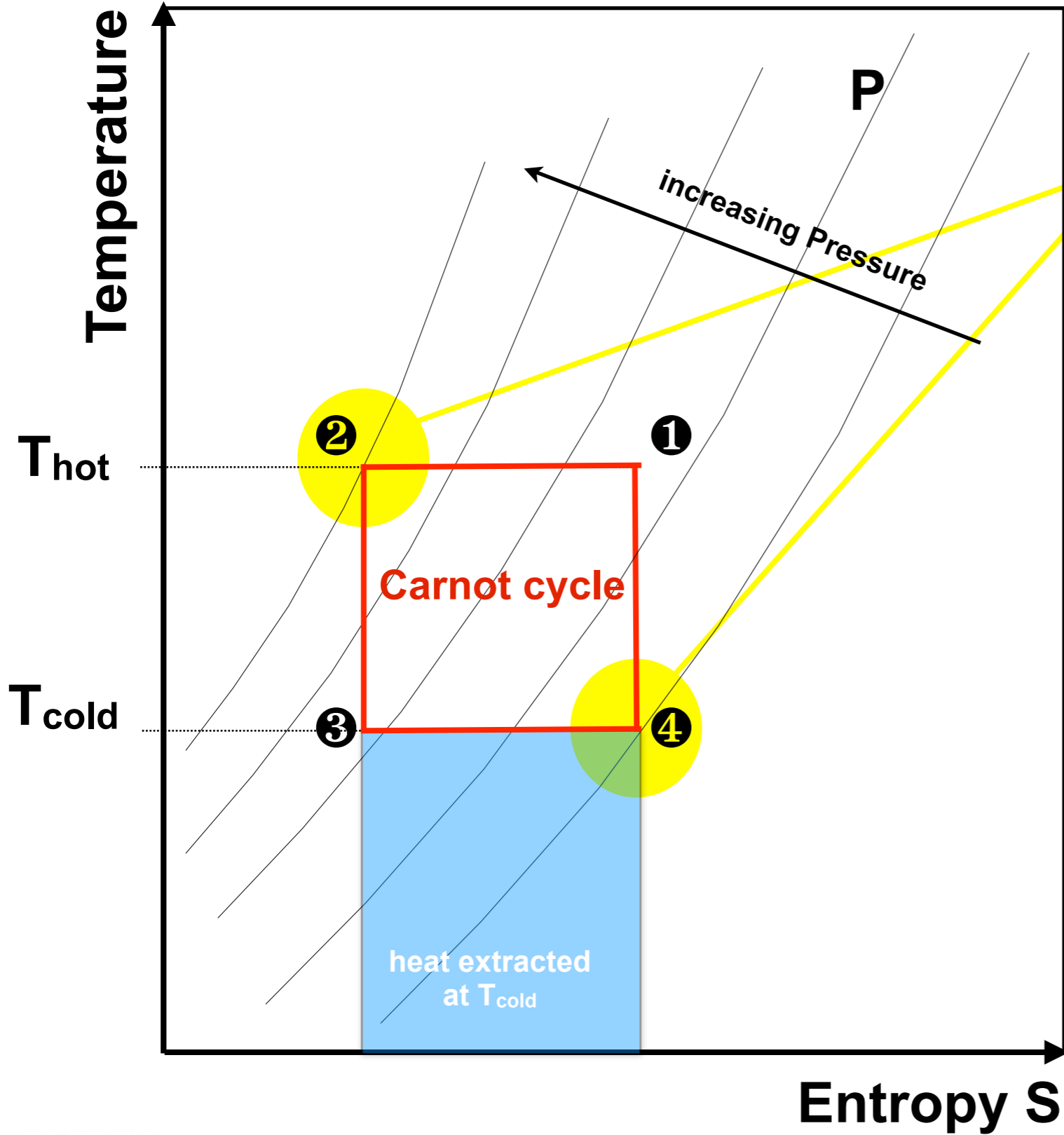
- **Lifetime**
- **Volume**
- **Ground tests**
- **Warm launch**



- **Heat distribution**
- **Interface**
- **Vibration**
- **Thermal peak power**

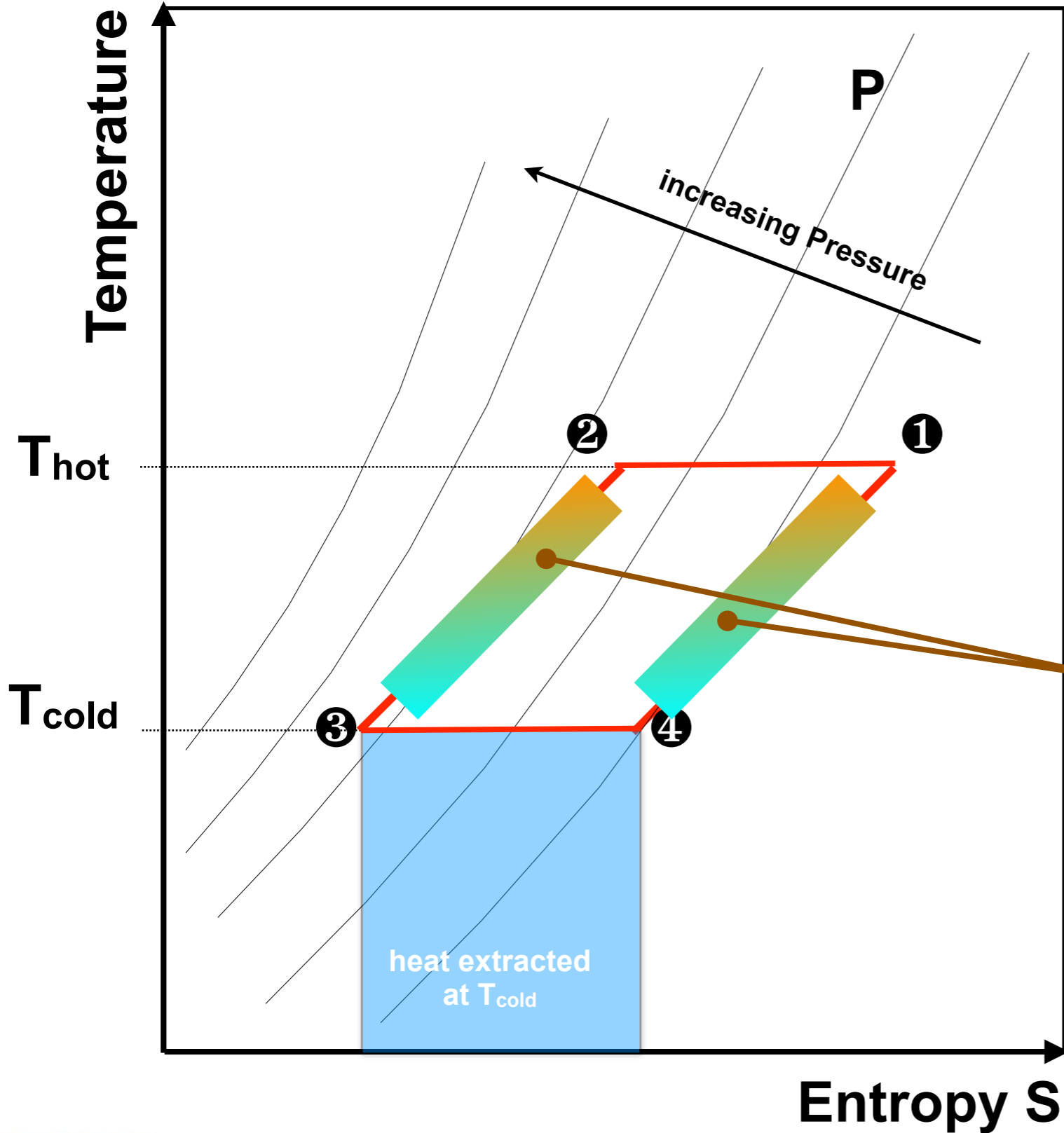


ONLY THE BEST: MYTHICAL CARNOT CYCLE



Example
300 K - 60 K cooler
Carnot cycle
pressure ratio ②/④ \approx 120 !!

ONLY THE BEST: MYTHICAL CARNOT CYCLE



Example
300 K - 60 K cooler

Carnot cycle
pressure ratio $2/4 \approx 120 !!$

Isentropic evolution

Isochoric
(Stirling)

Isobaric
(Ericsson, GM)

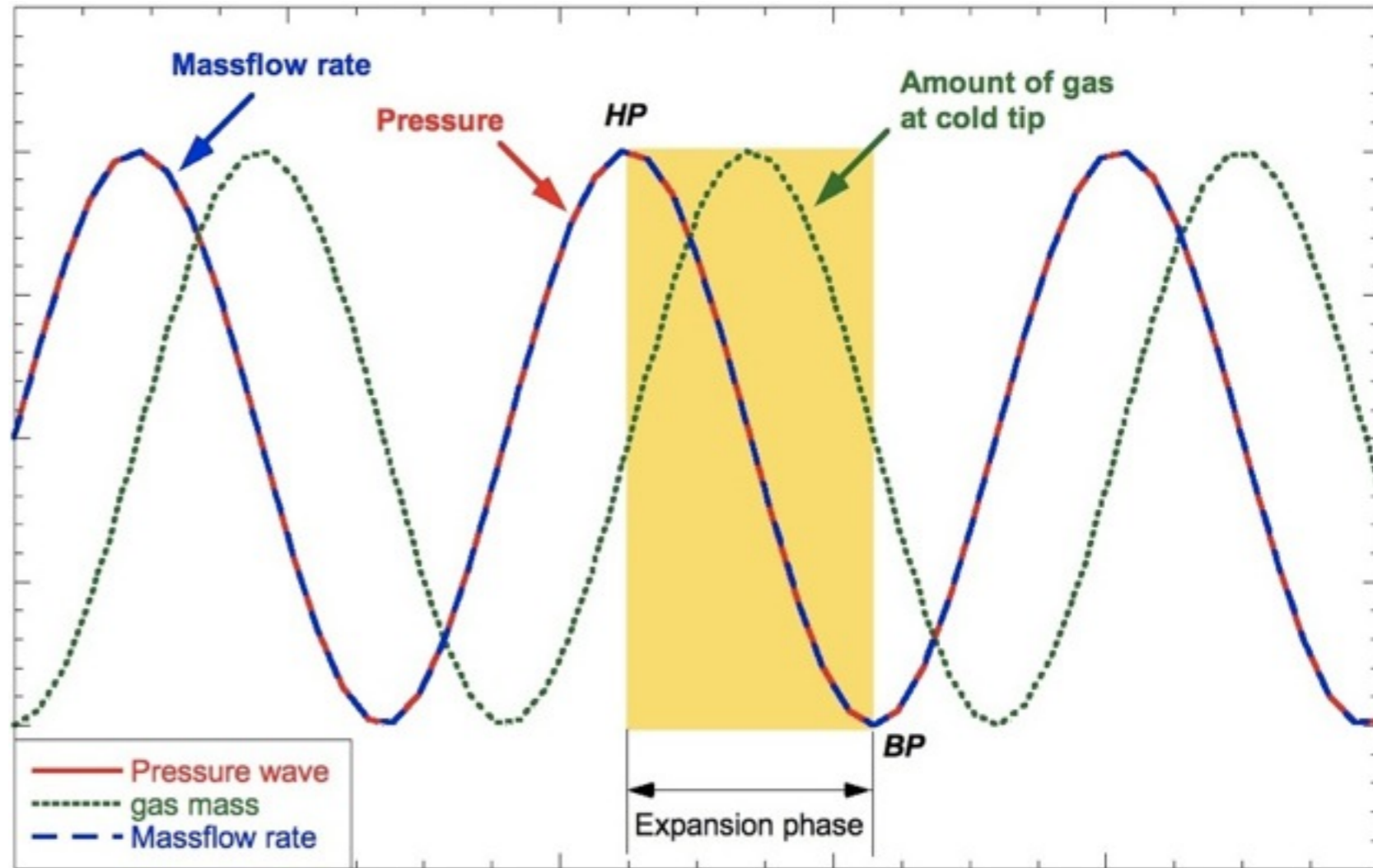
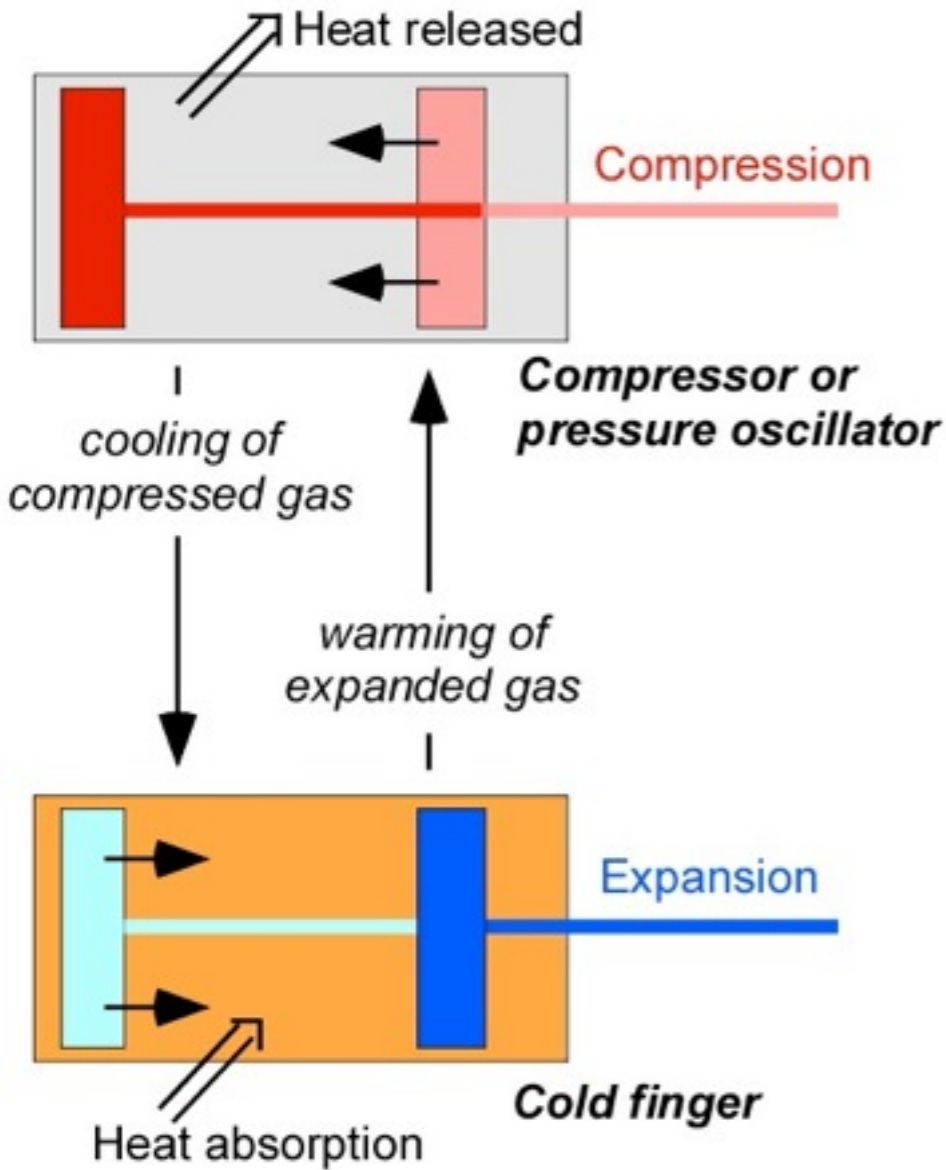
In theory, COP \approx Carnot

heat transfer at variable T



*Regenerator (thermal sponge)
or counter flow heat exchanger*

(MOST) MECHANICAL CRYOCOOLERS



Expansion of maximum amount of gas at the right place and right time



Mass flow rate & pressure oscillation in phase

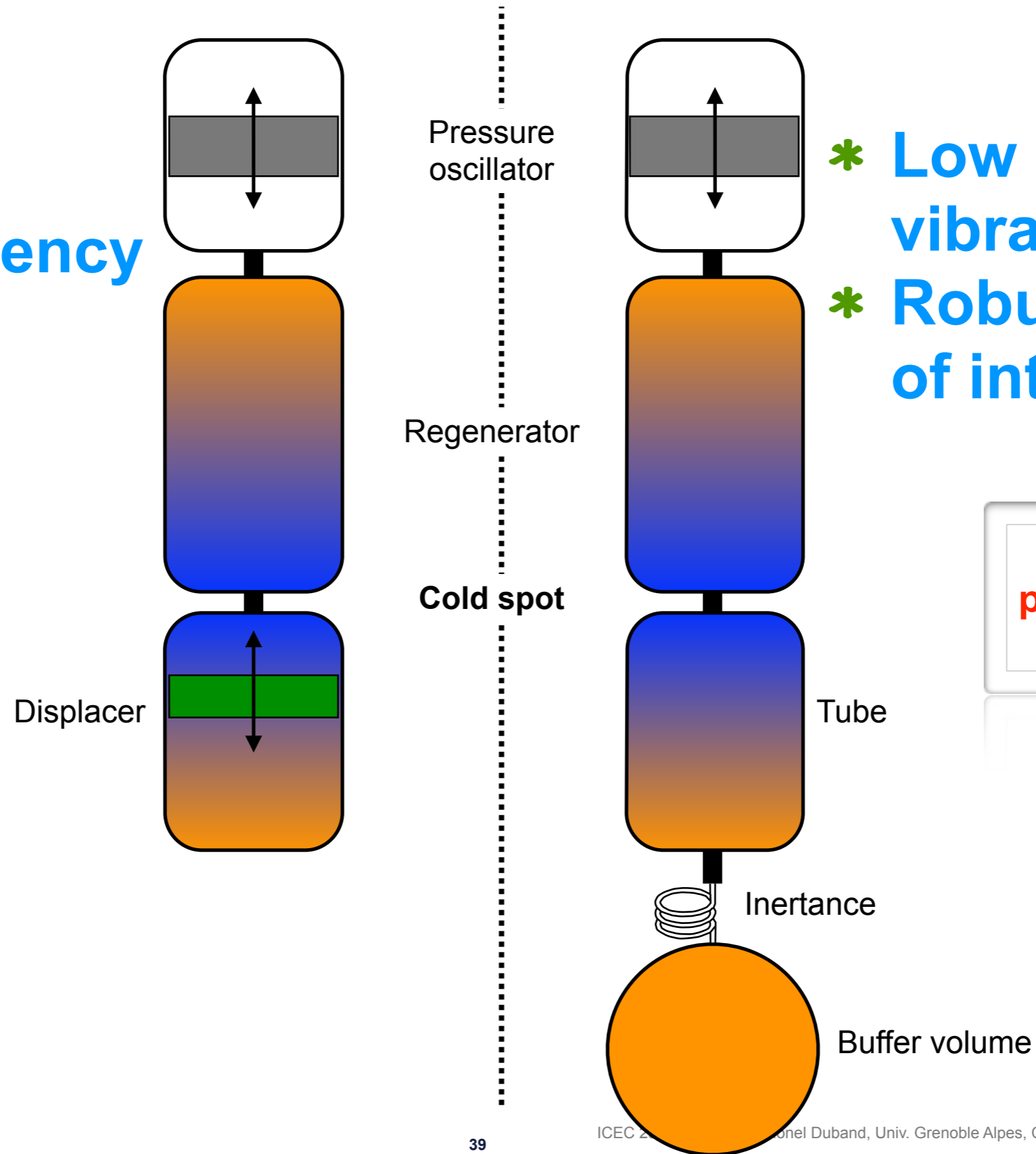
STIRLING VERSUS PULSE TUBE

* **High efficiency**

* **Low induced vibration**
 * **Robust / Ease of integration**

Phase shift mechanically controlled

Phase shift pneumatically controlled



controlled

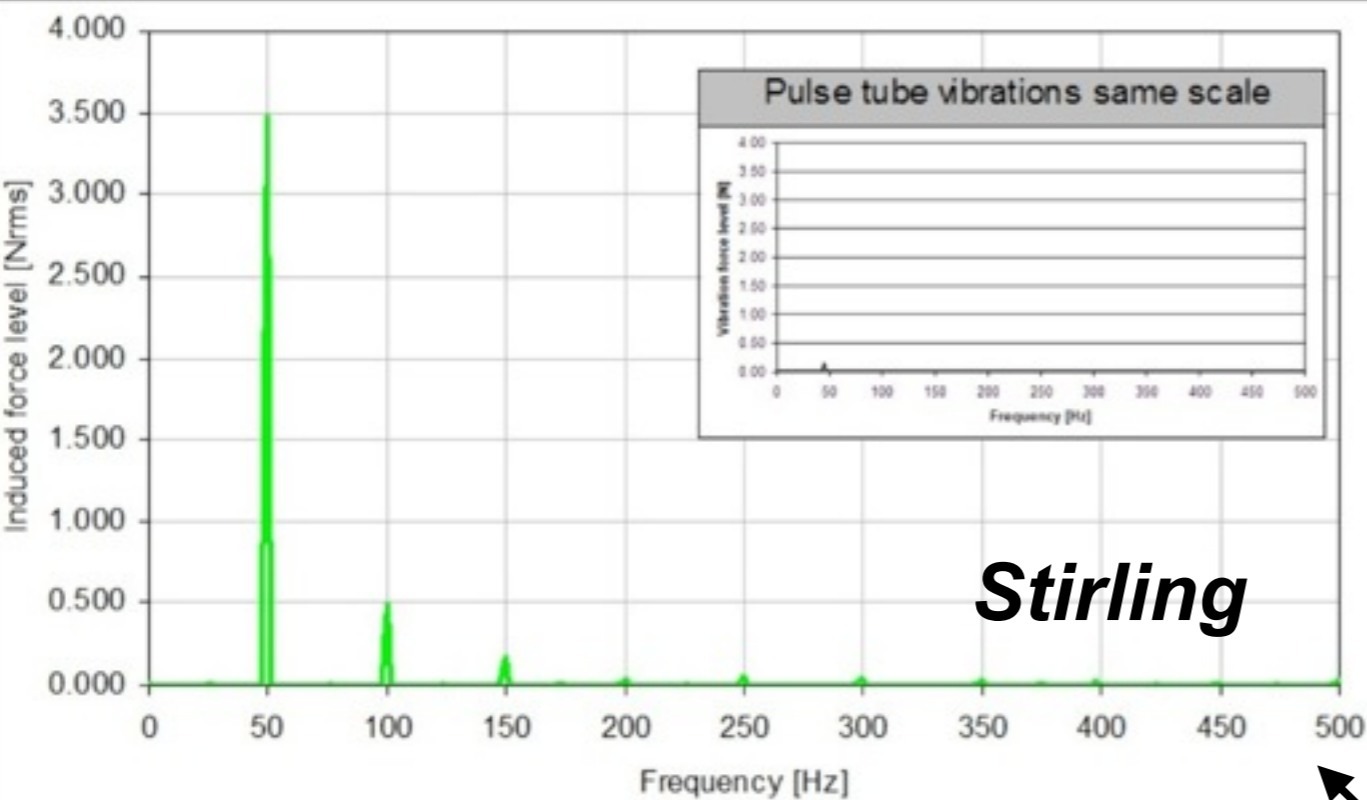
controlled

STIRLING VERSUS PULSE TUBE

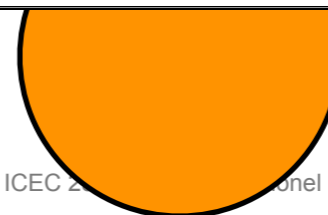
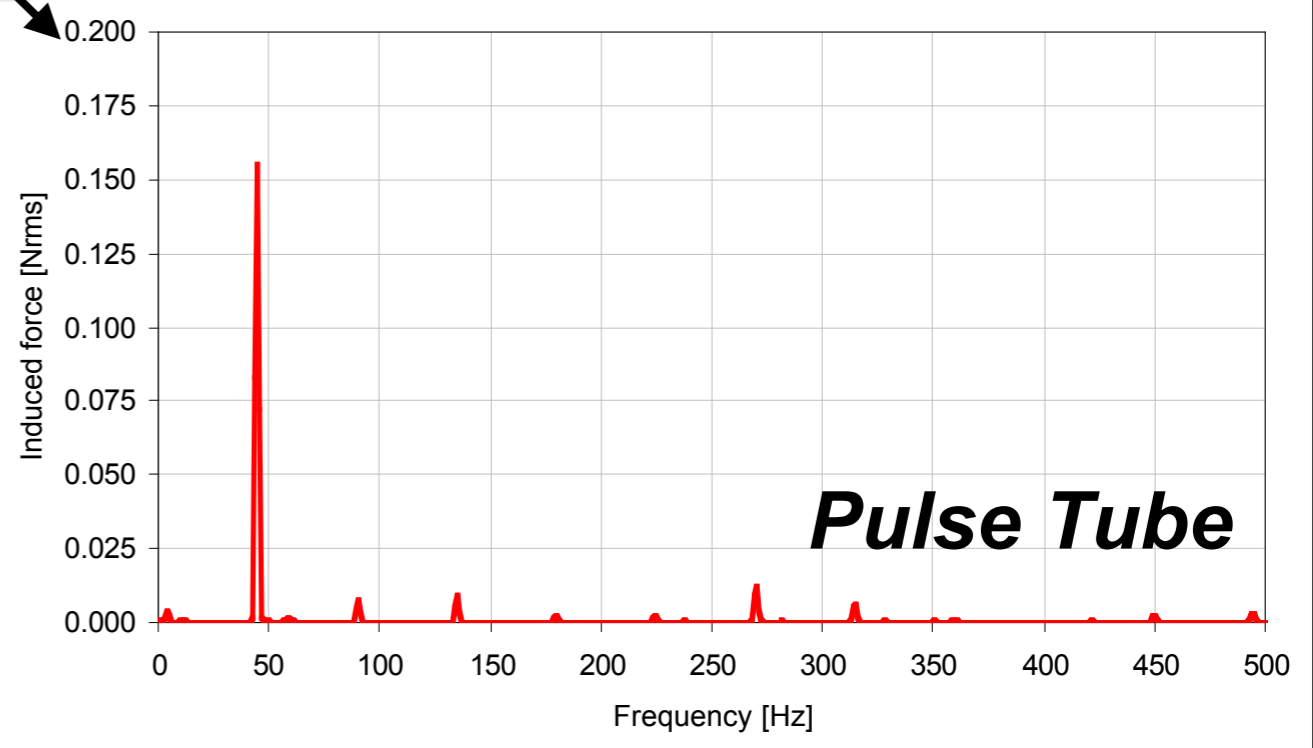
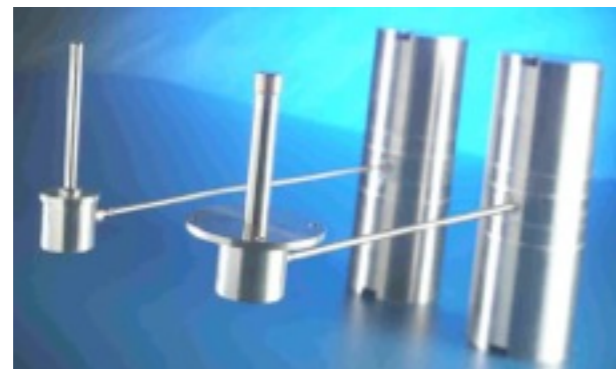
THALES



LPT 9510 Pulse Tube



LSF 9589 Stirling



Buffer volume

COOLER DNA (PRESSURE OSCILLATOR)

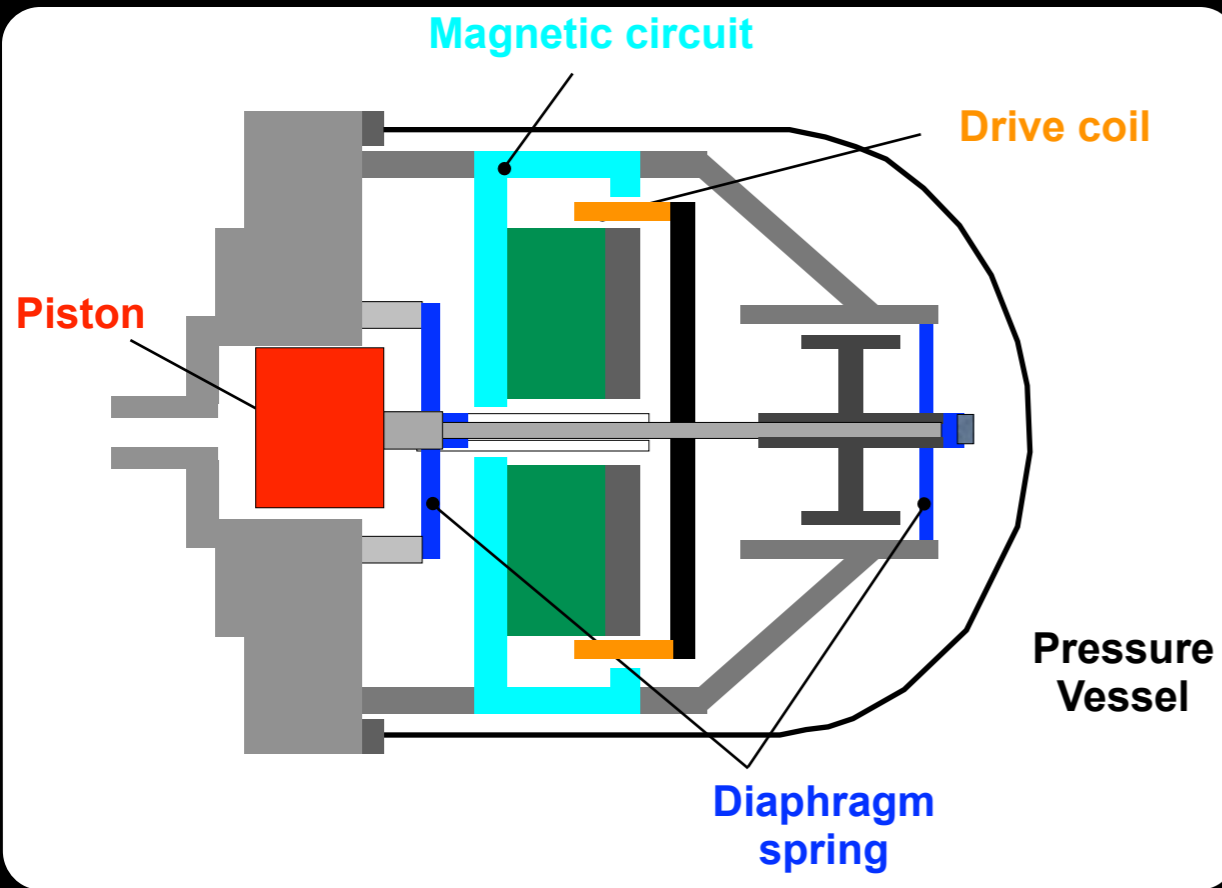
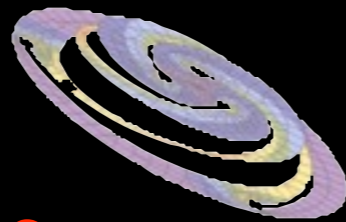
High Frequency ($\approx > 30$ Hz)

Motion without friction (\approx)

No Maintenance



Key Concept



So called "Oxford type"



Science & Technology Facilities Council
Rutherford Appleton Laboratory

The precursor 80K Single Stage Stirling



STIRLING AND PULSE TUBE COOLERS

Overall:

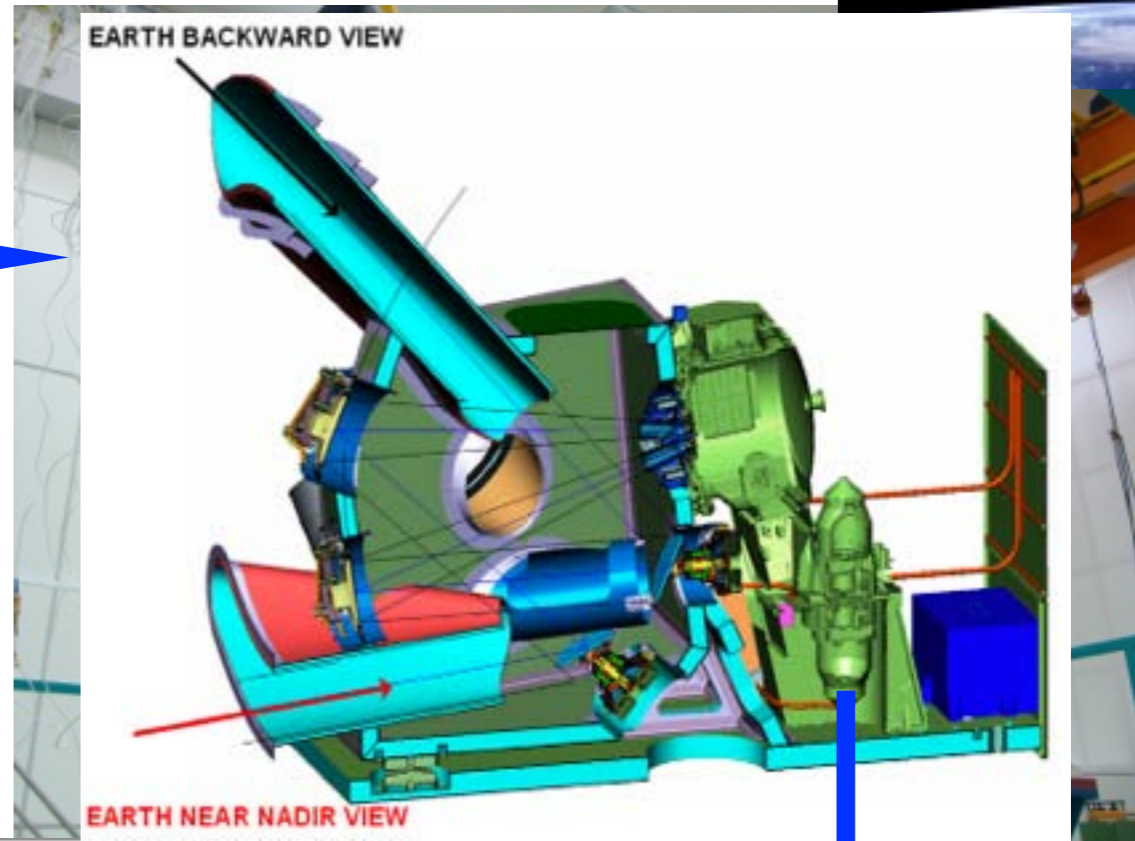
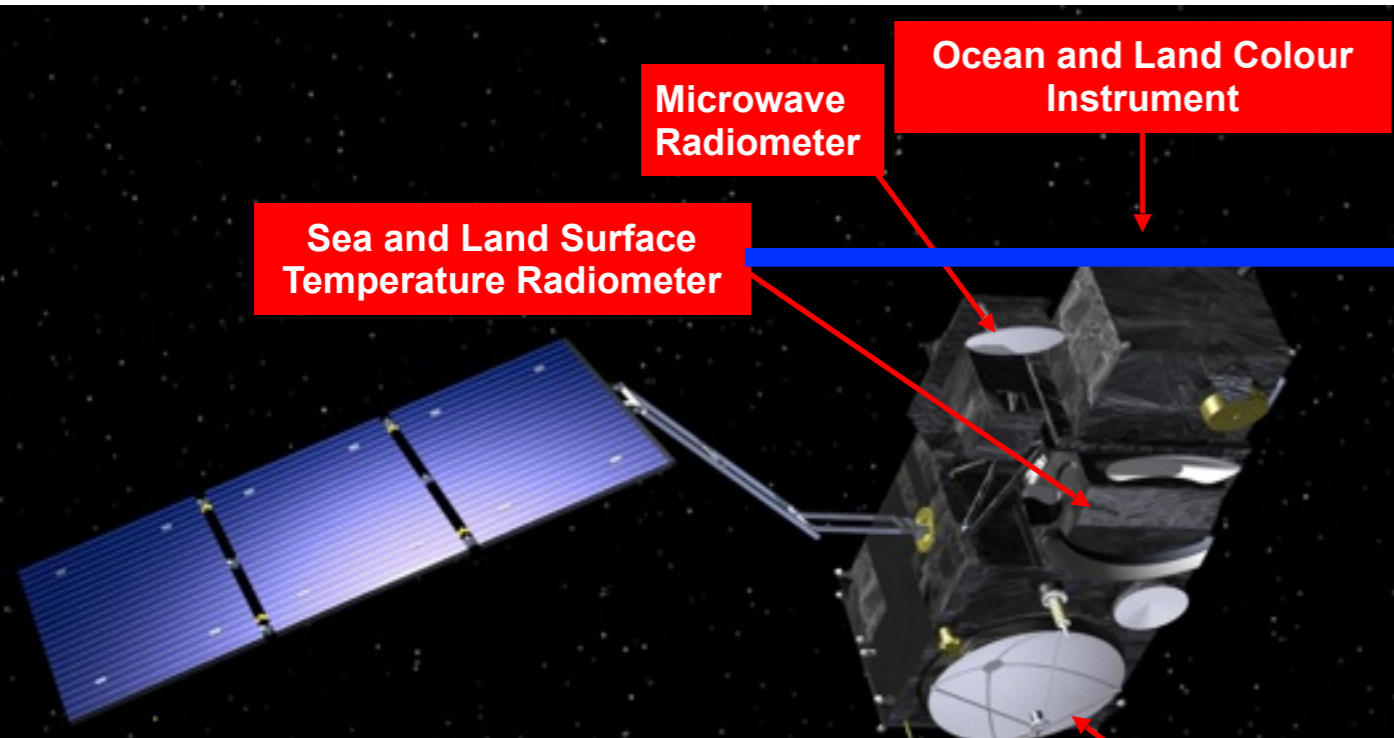
- Mature technology**
- High Technical Readiness Level (TRL)**
- several thousands of hours in operation in orbit**

Original BAe 50-80 K cooler: currently 23 in orbit, 177 years accumulated in orbit,

AIRS PT: over 12 years of operation

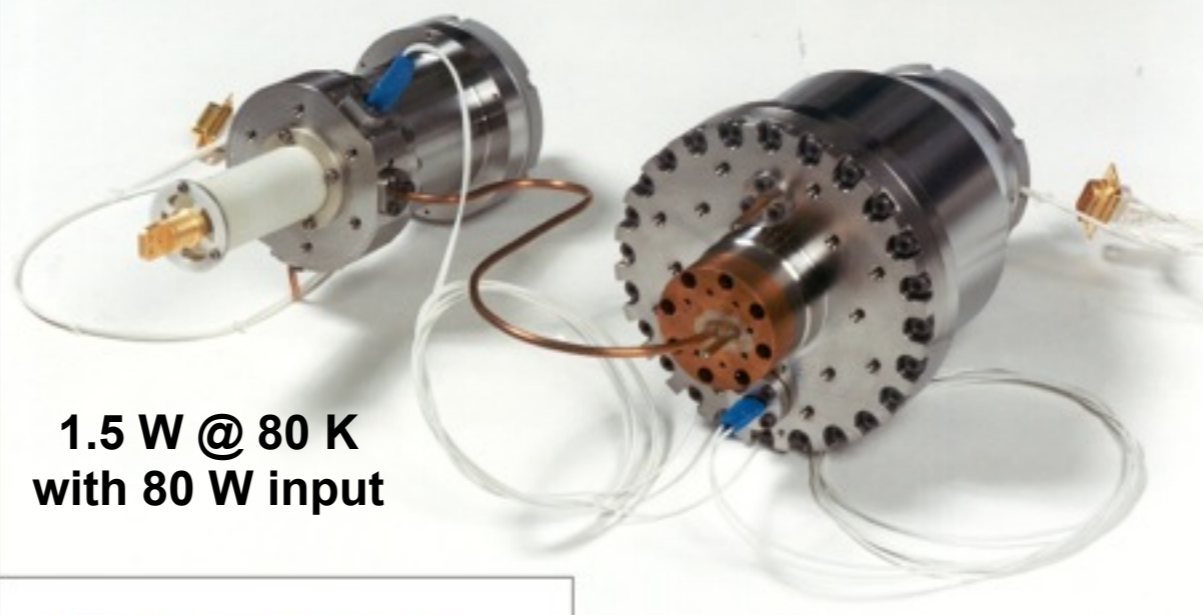
- On going: Multi stage, low T system, Miniature cooler, high frequency**

SENTINEL 3: OCEAN & GLOBAL LAND MONITORING

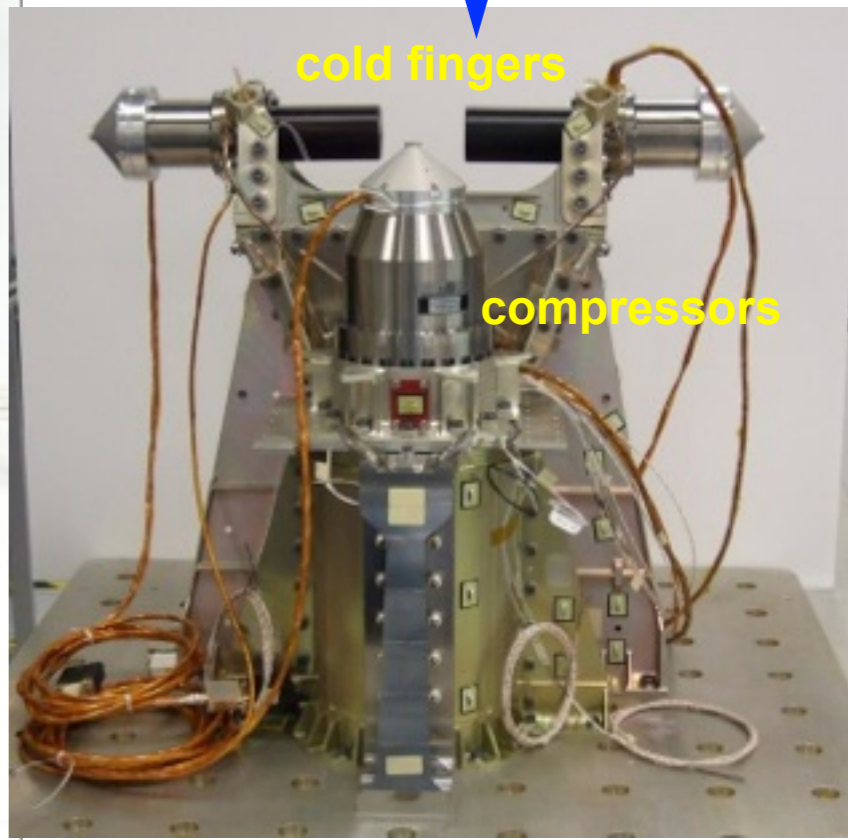


- Main satellite characteristics**
- 1250 kg maximal mass
 - 3.89 m x 2.202 m x 2.207 m
 - Average power consumption of 1100 W
 - **7.5 years lifetime (fuel for 5 add. years)**

Scheduled to launch mid 2015



**1.5 W @ 80 K
with 80 W input**

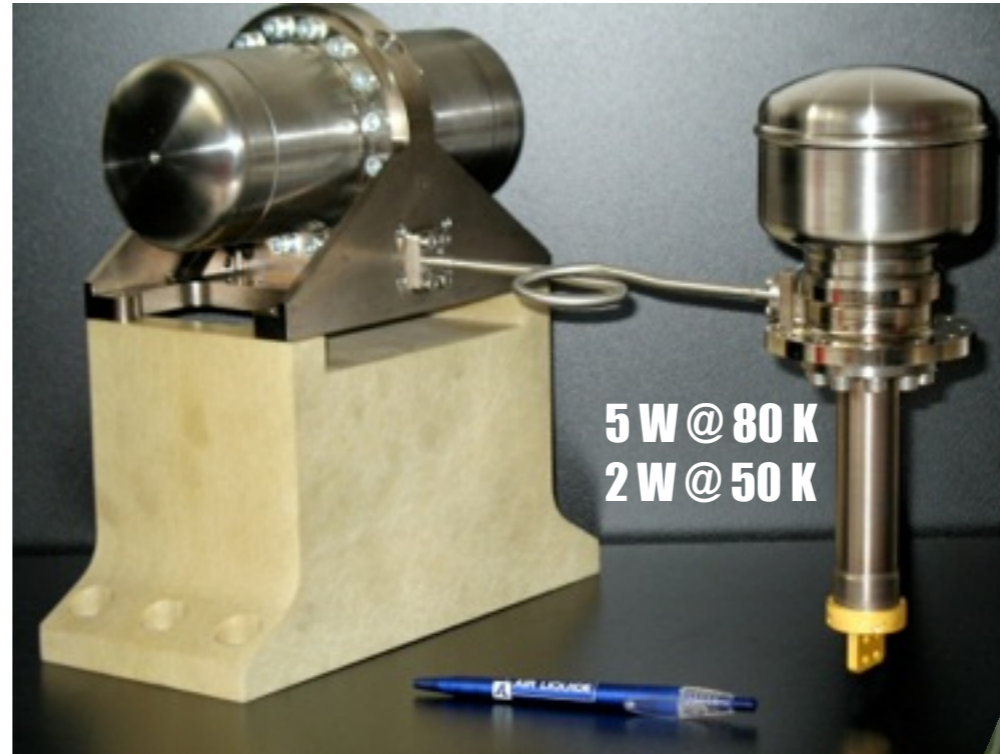
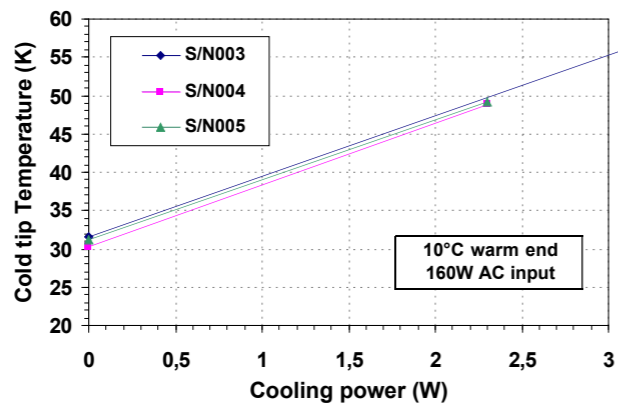


METEOSAT 3G

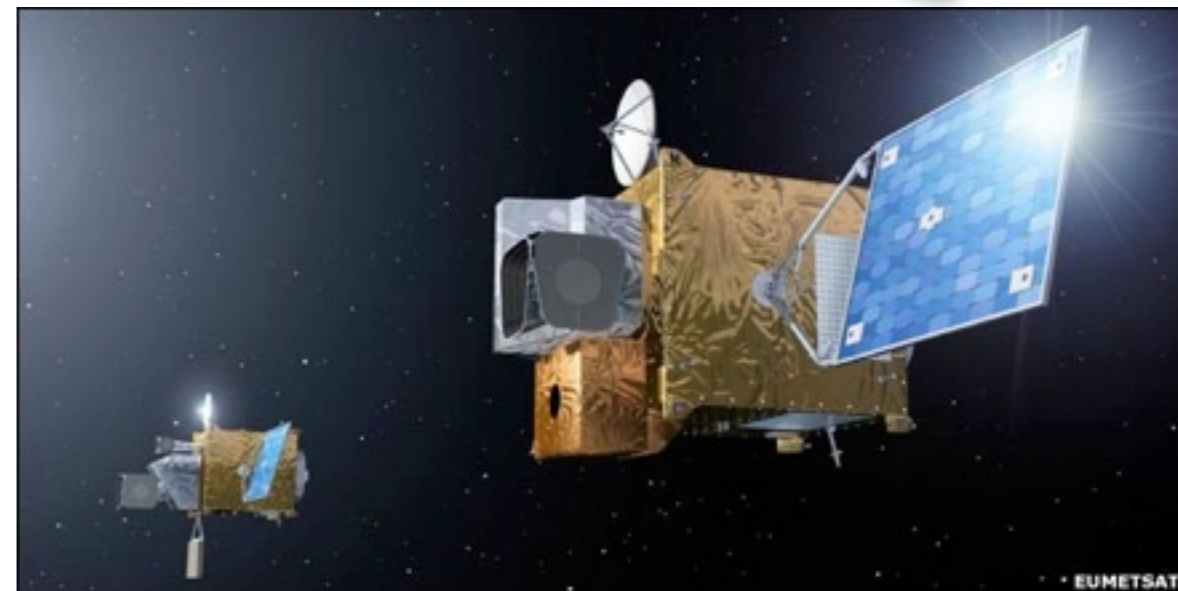
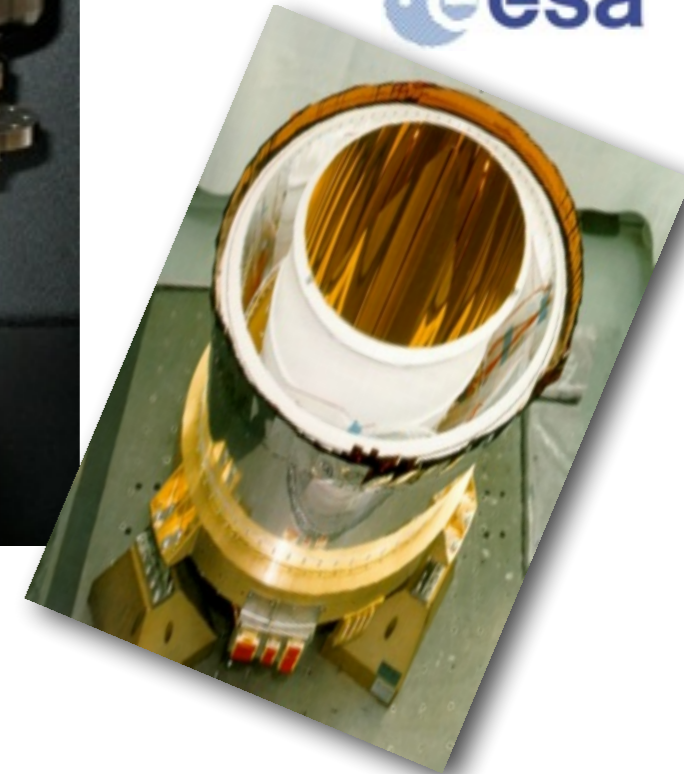
50 K Single Stage



2.3 W @ 50 K
160 W input
5.1 Kg



LPTC



ALTERNATIVE APPROACH: COOLER OF THE SHELF (COTS)

“High costs”: Full space qualification, Maximum performance.

- In line with ESA standards, adherence to ESA/NASA guidelines

“Low costs”: Reliability is important, High price level not affordable.

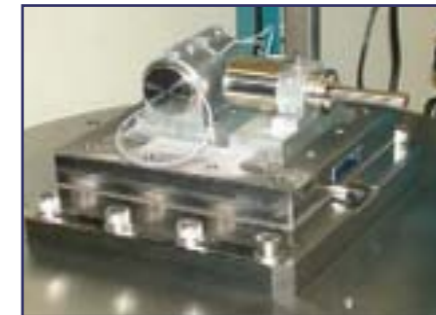
- Products based on normal definition and production standards
- Special but limited screening of parts / products
- Extra burn-in to avoid infant mortality

Cost effective solution: Reliability is key, No extreme performance.

- Design based on existing / proven definitions
- Extra but limited effort on parts and processes (based on risk assessment)
- Extensive screening of subassemblies / final products



LPT 9310 in test configuration



LPT 9510 on JPL test bench

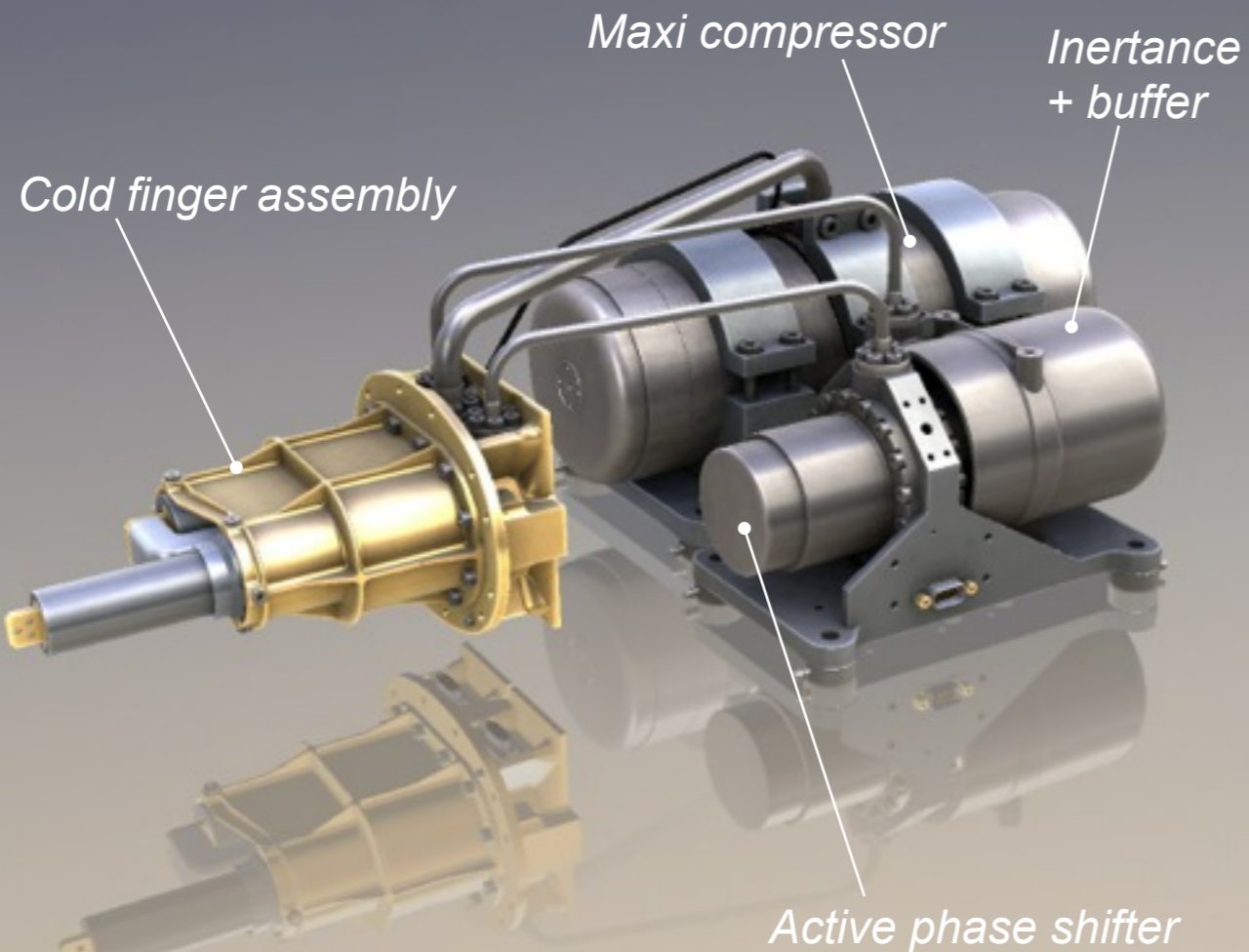


ORS I Satellite (launched 2011)



THALES Split Stirling LSF 9599

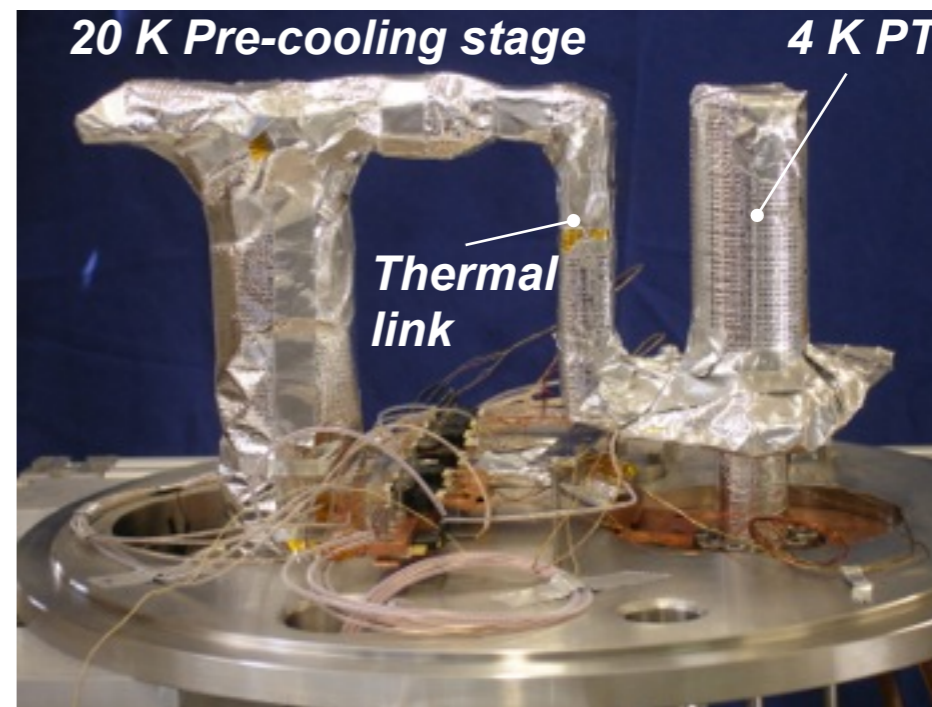
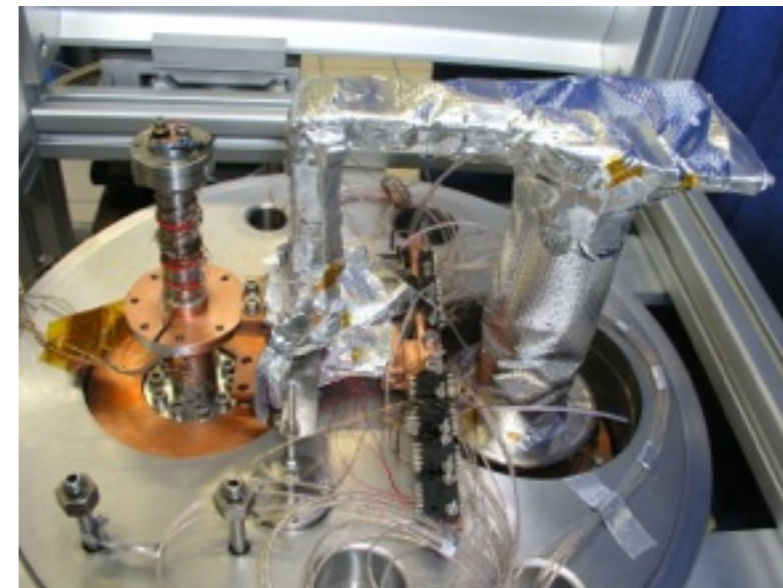
LOW TEMPERATURE MULTISTAGE PULSE TUBE



300 mW @ 15 K

THALES

4 K Pulse Tube



**3.86 K reached June 2014 !
15 mW @ 4.5 K**

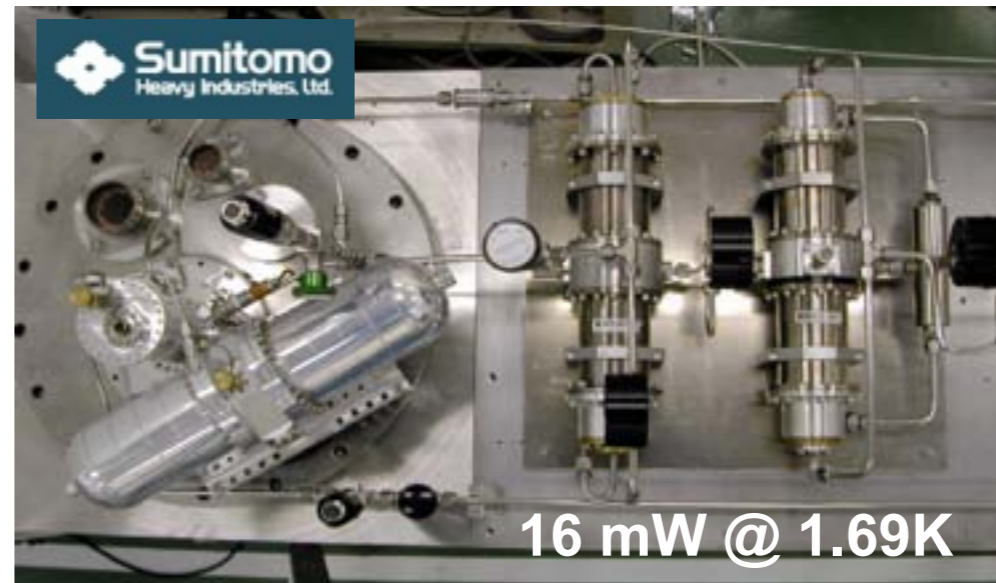
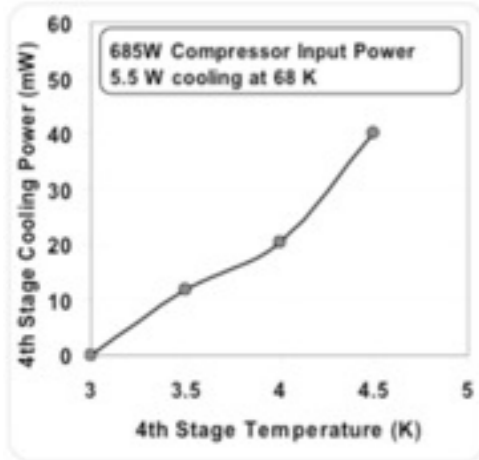
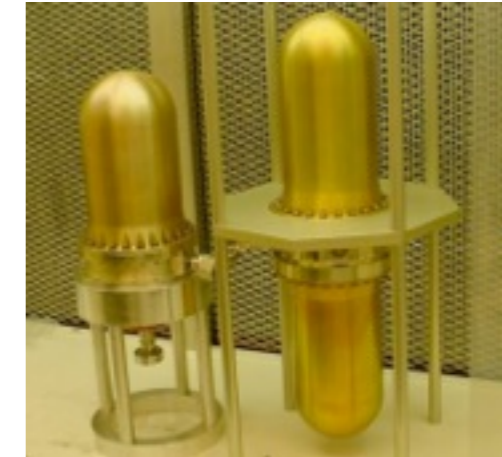
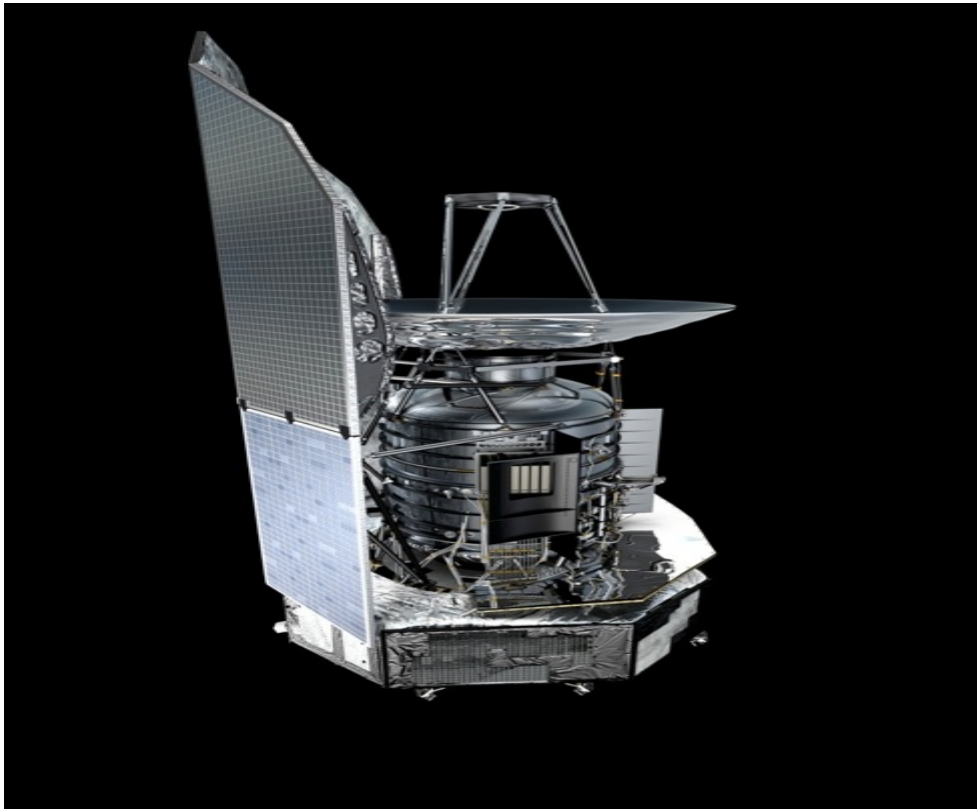
ATHENA Mission



Possible application

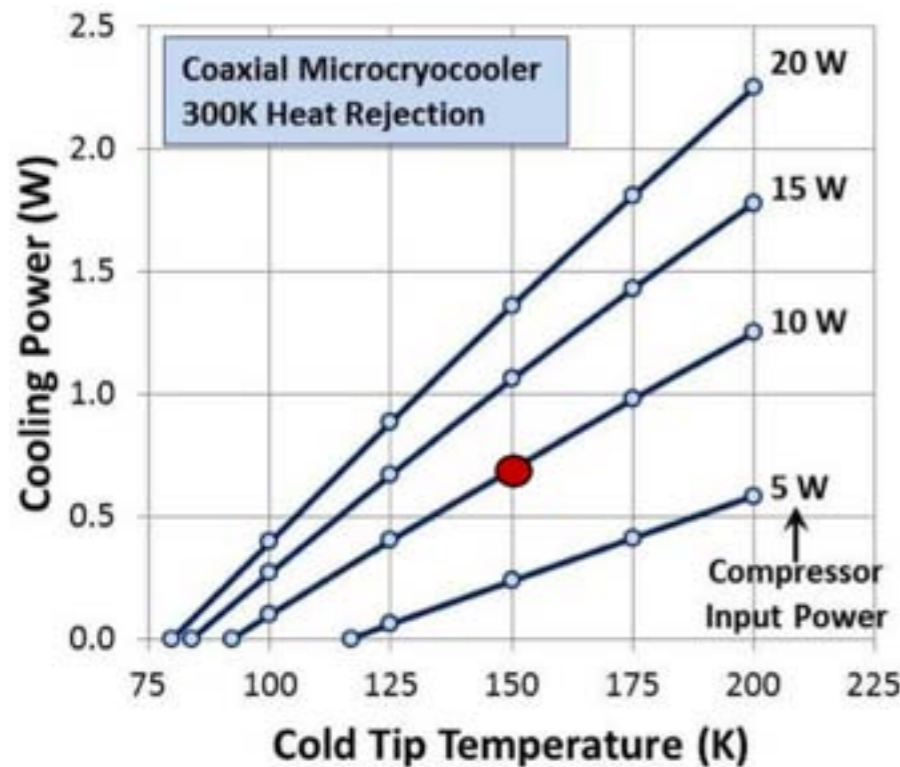


CAN'T BE EXHAUSTIVE - MANY SUPPLIERS !

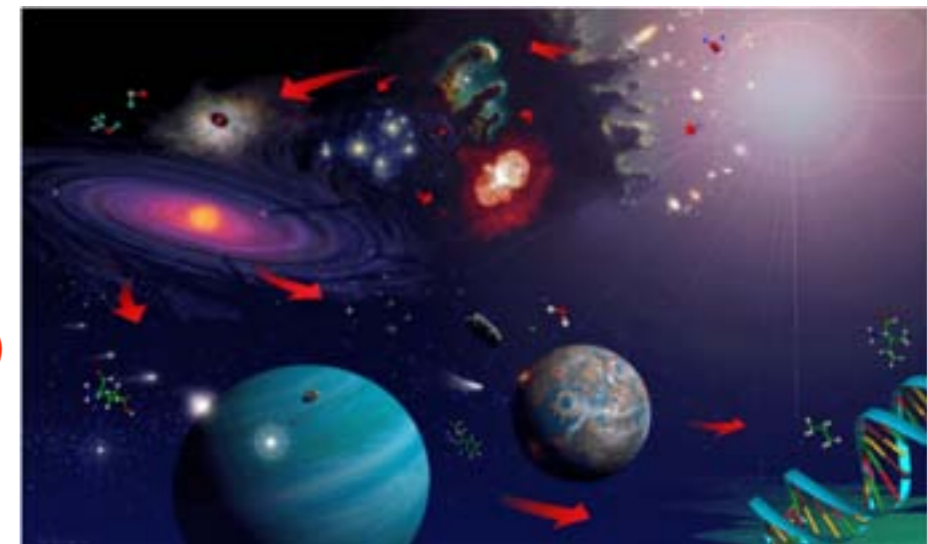


PULSE TUBE MICROCRYOCOOLER

Low SWaP (Size, Weight and Power)

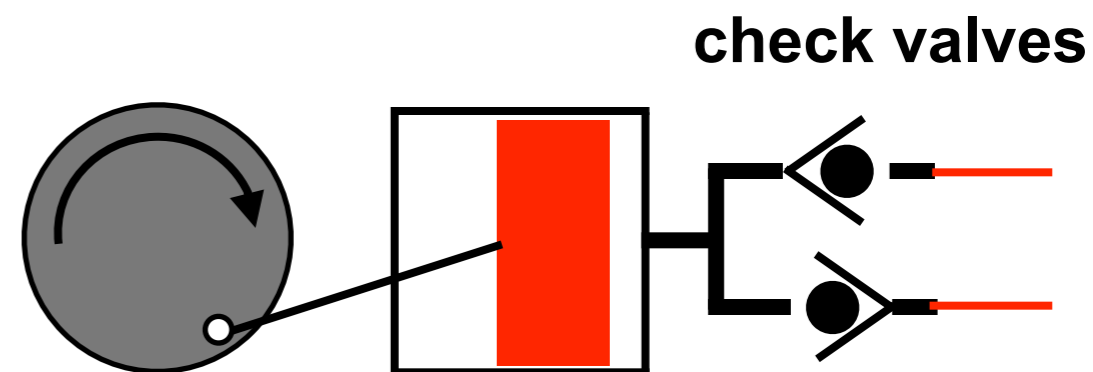
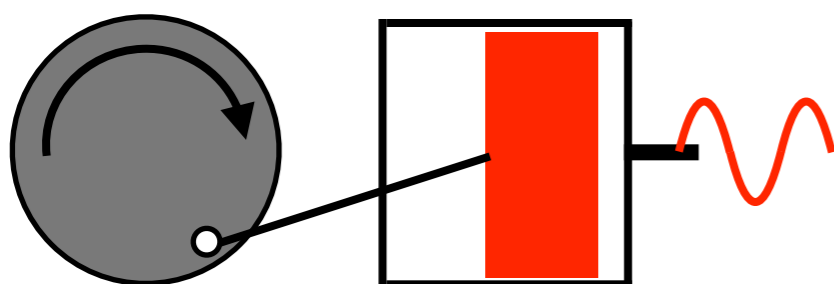
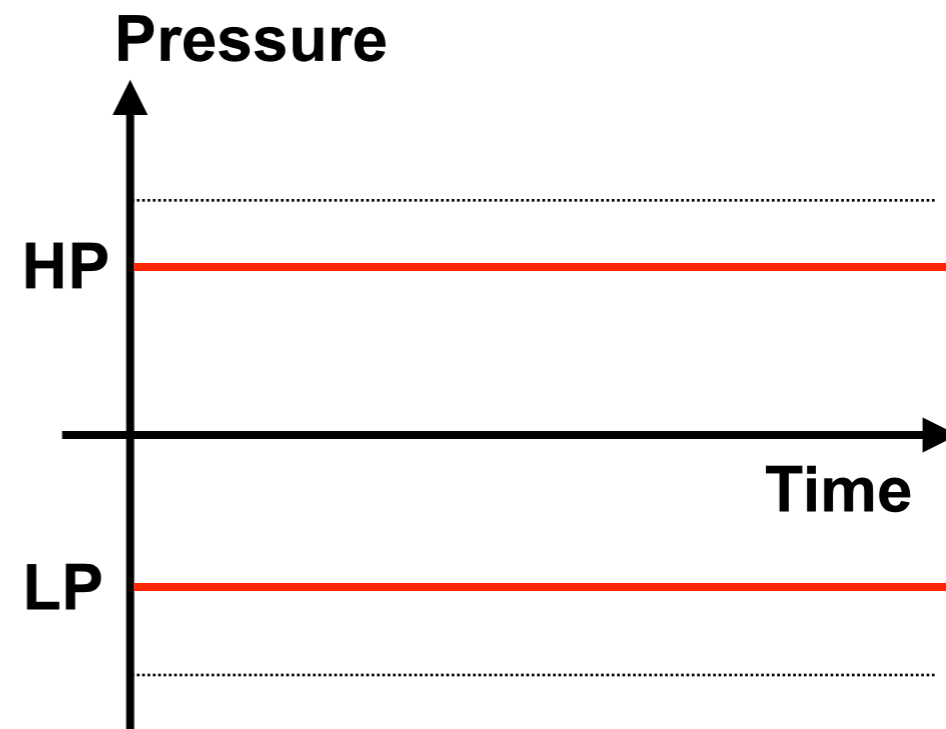
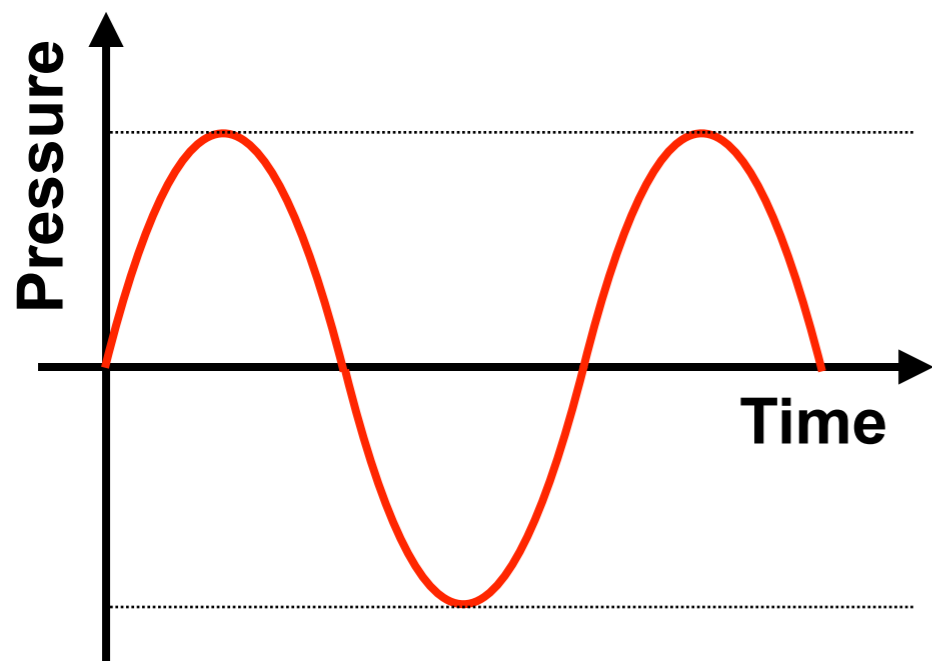


690 mW @ 150 K
10 W input
100 Hz
compressor Ø32 x 90 mm
Cold finger Ø42 x 110 mm
Mass: 328 gr (comp. 210, CF 118)



Could be used in the NASA MatISSE program
 (Maturation of Instruments for Solar System Exploration)

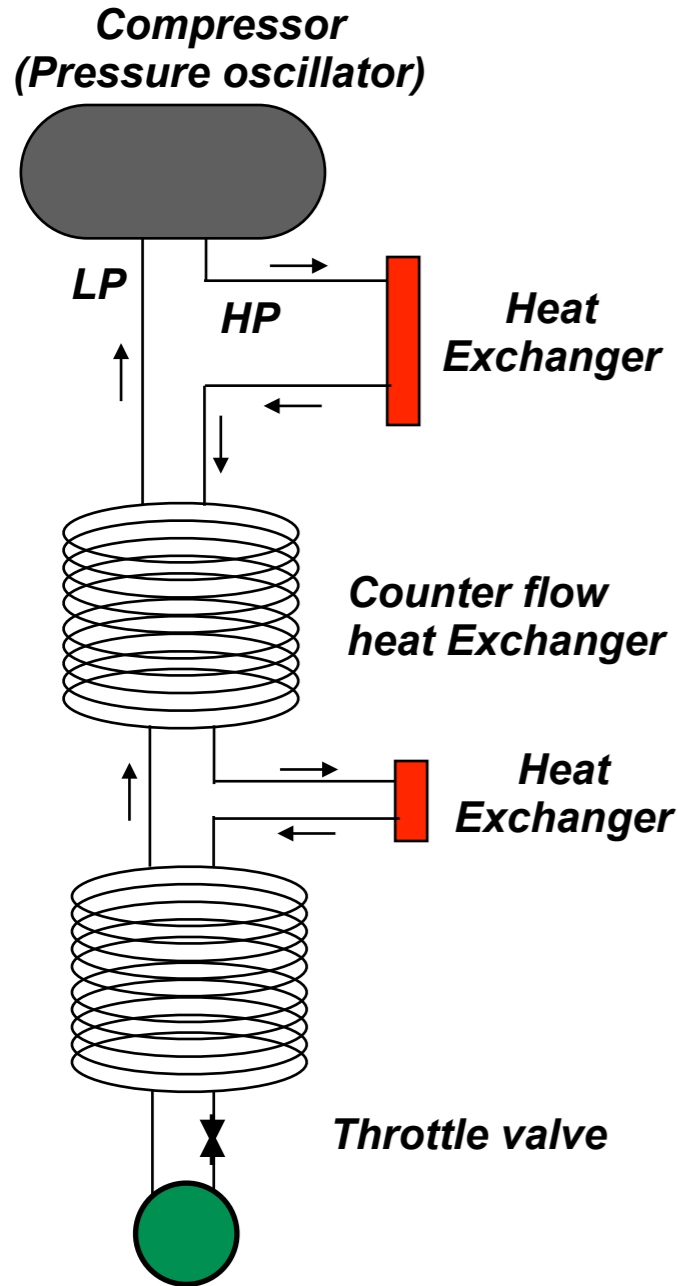
LOWER TEMPERATURE OR VIBRATION LESS: JOULE THOMSON SYSTEM



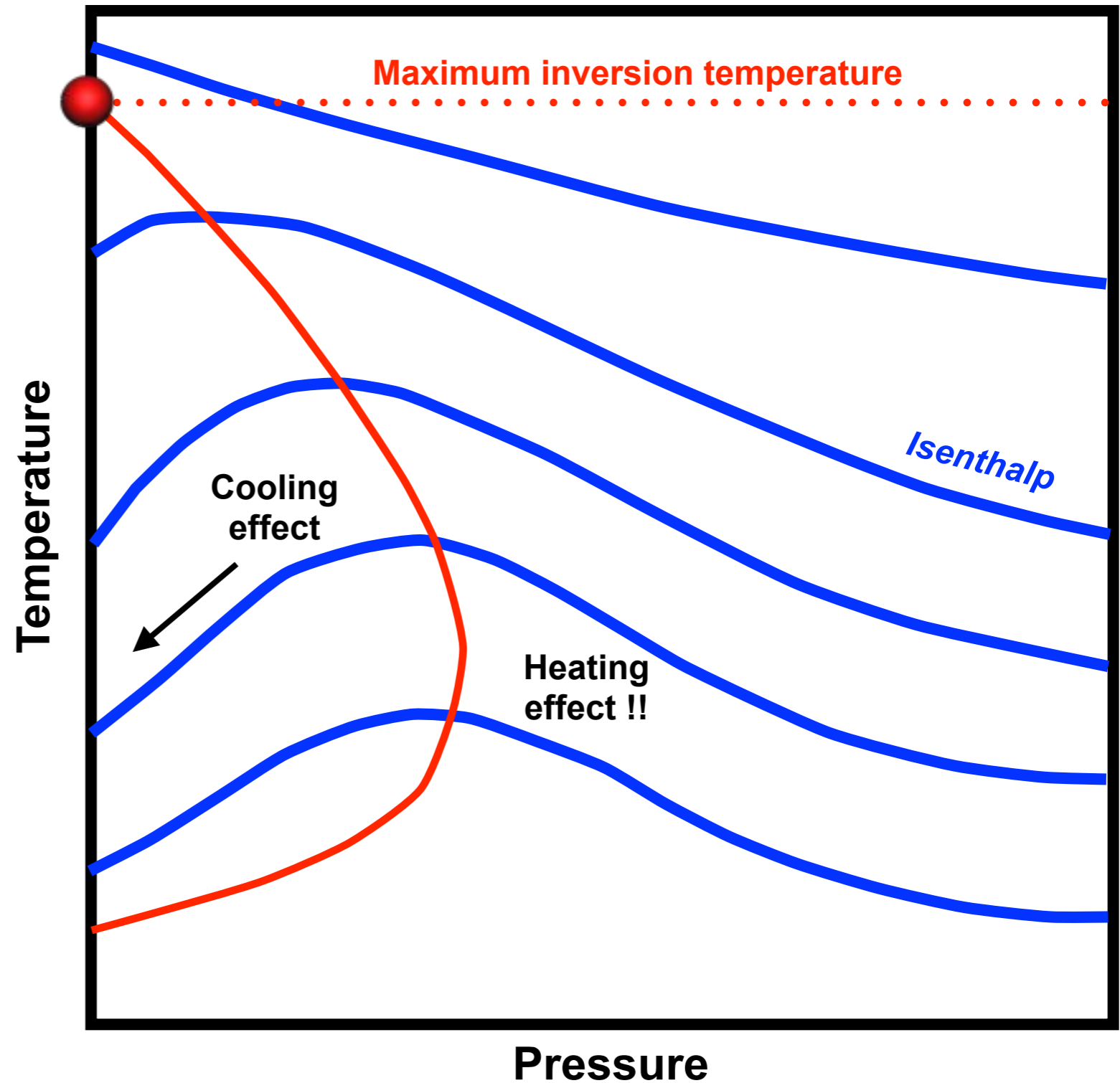
Direct benefit from heritage (flexure spring)

JOULE THOMSON LOOP

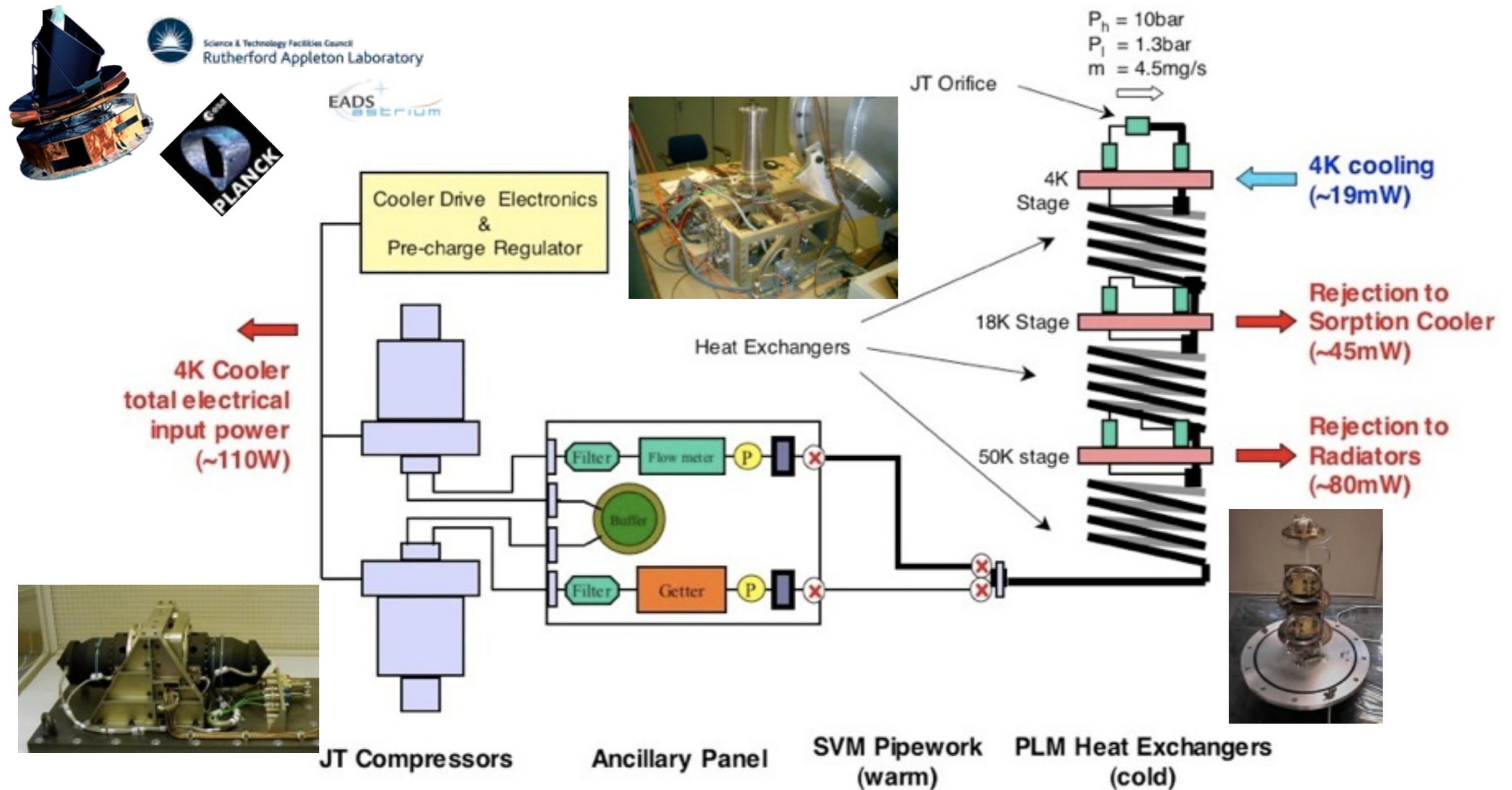
ISENTHALPIC EXPANSION



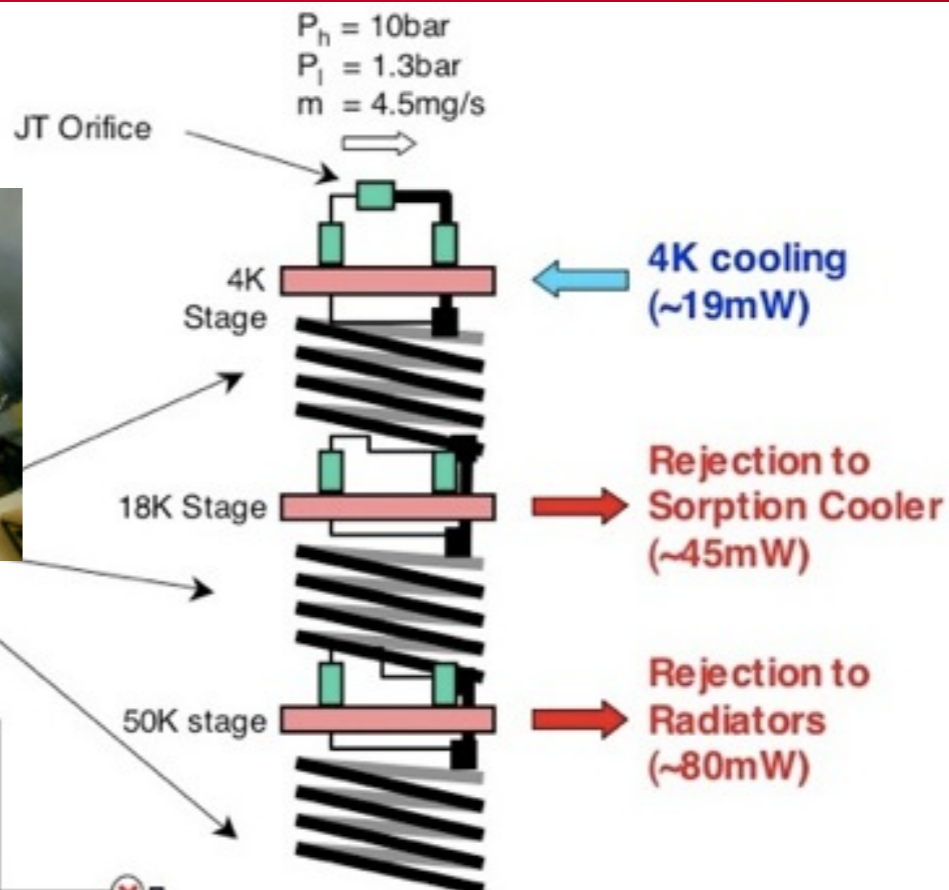
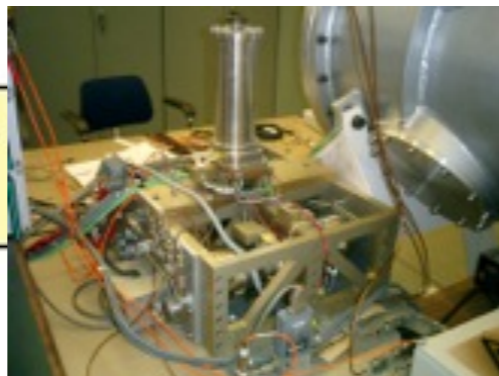
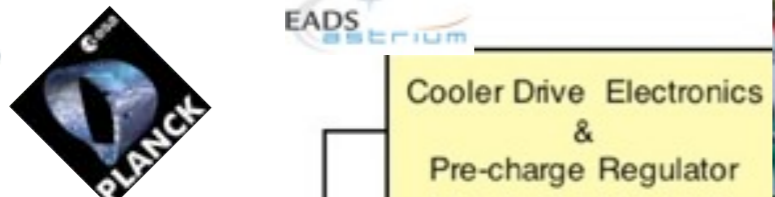
Helium:
 $\approx T \leq 40 \text{ K}$



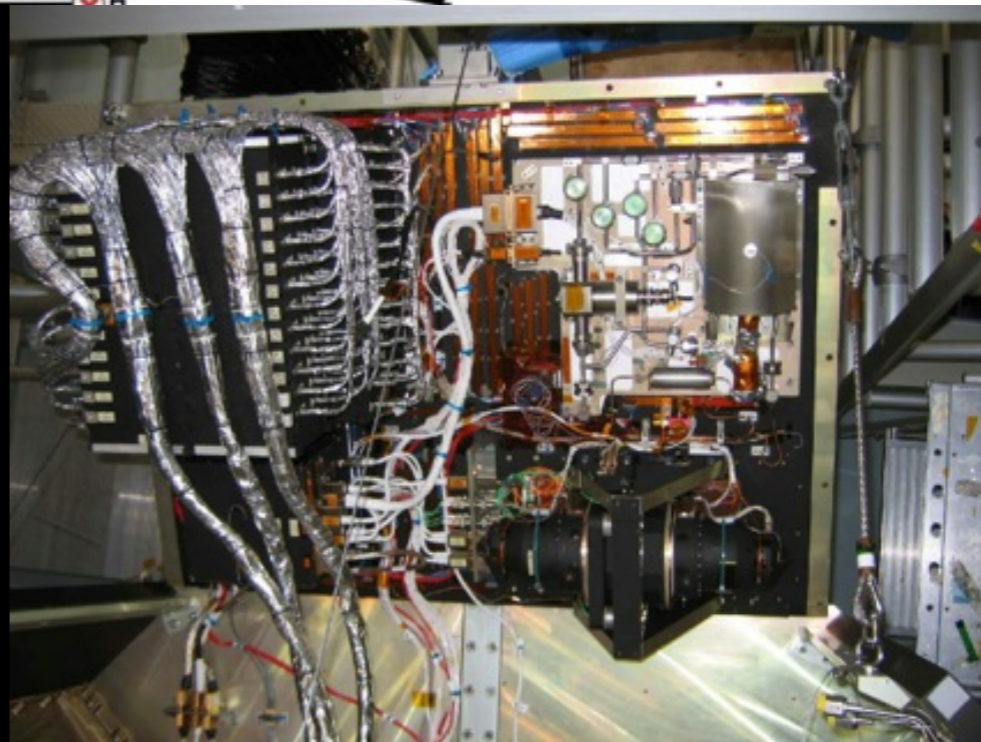
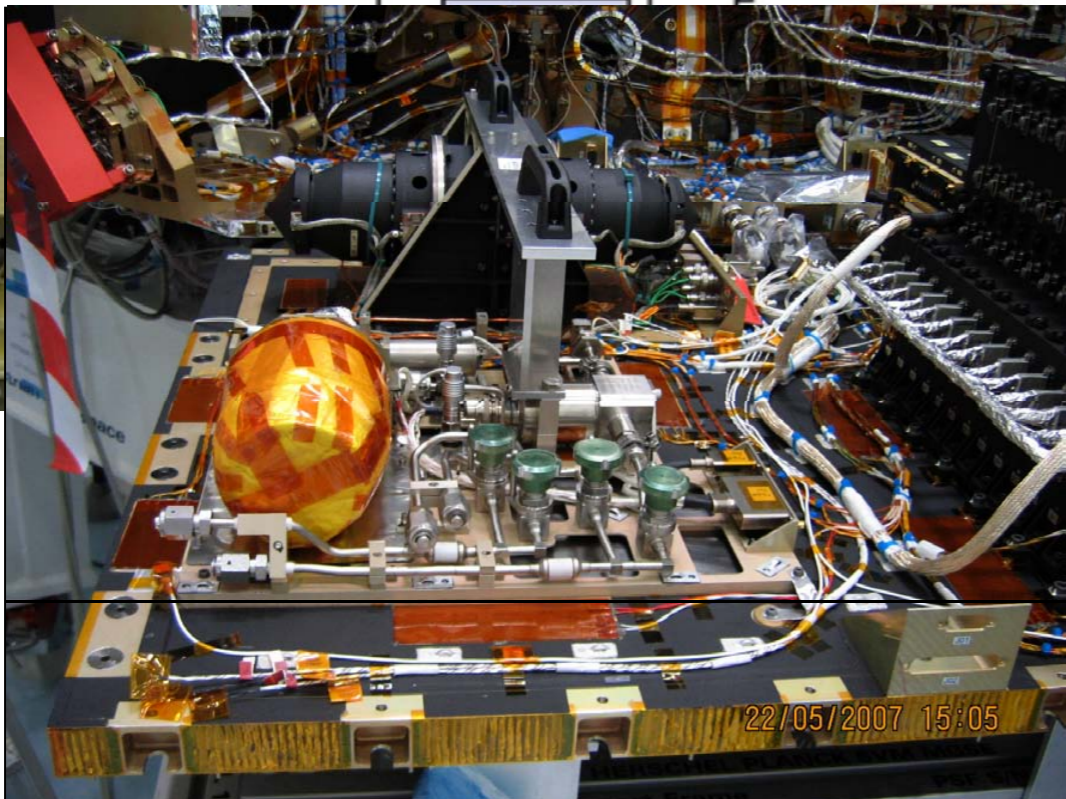
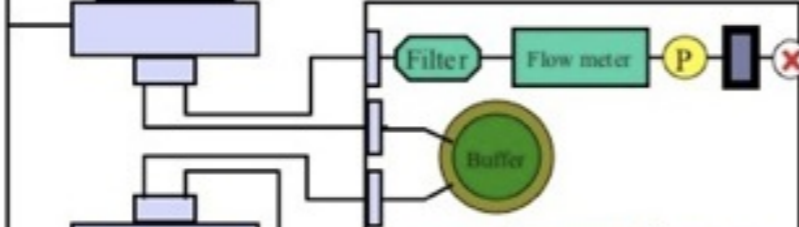
4K JT SYSTEM - PLANCK



4K JT SYSTEM - PLANCK

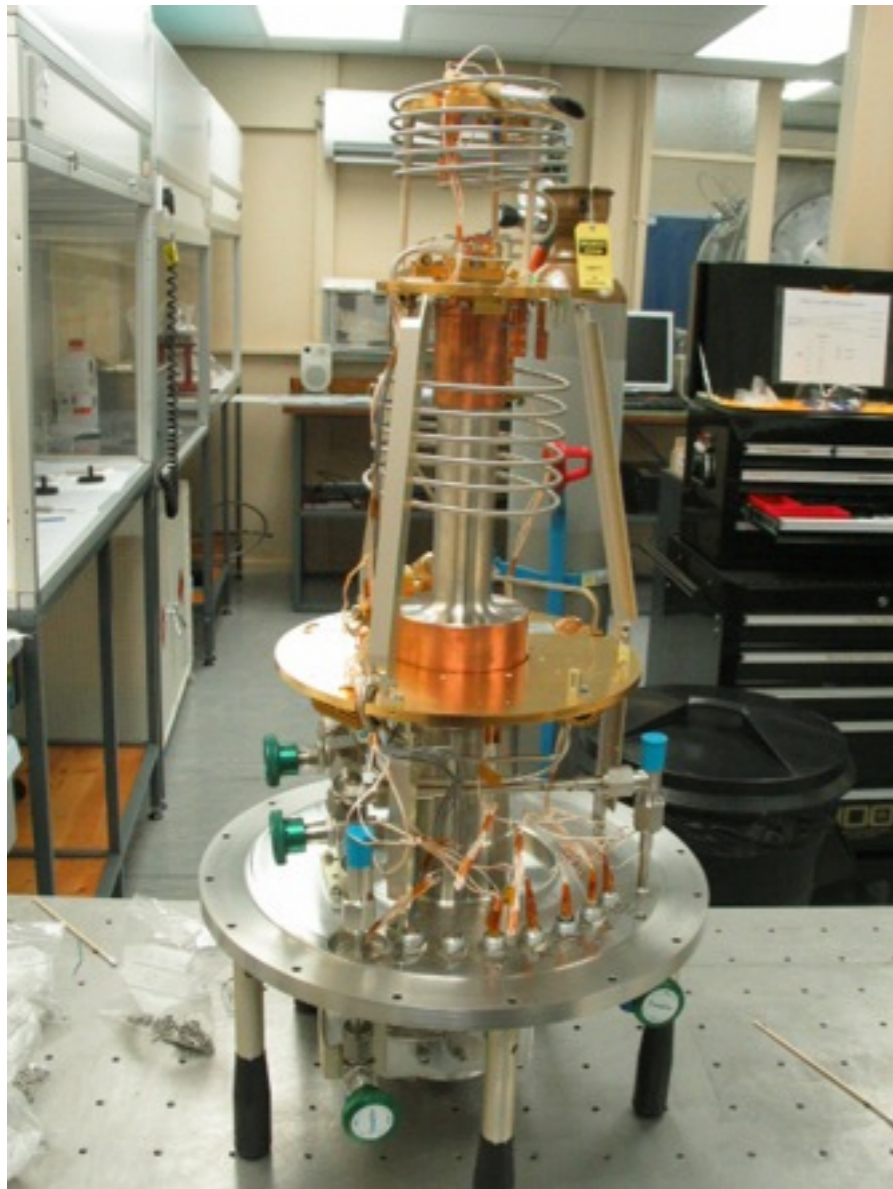


4K Cooler total electrical input power (~110W)

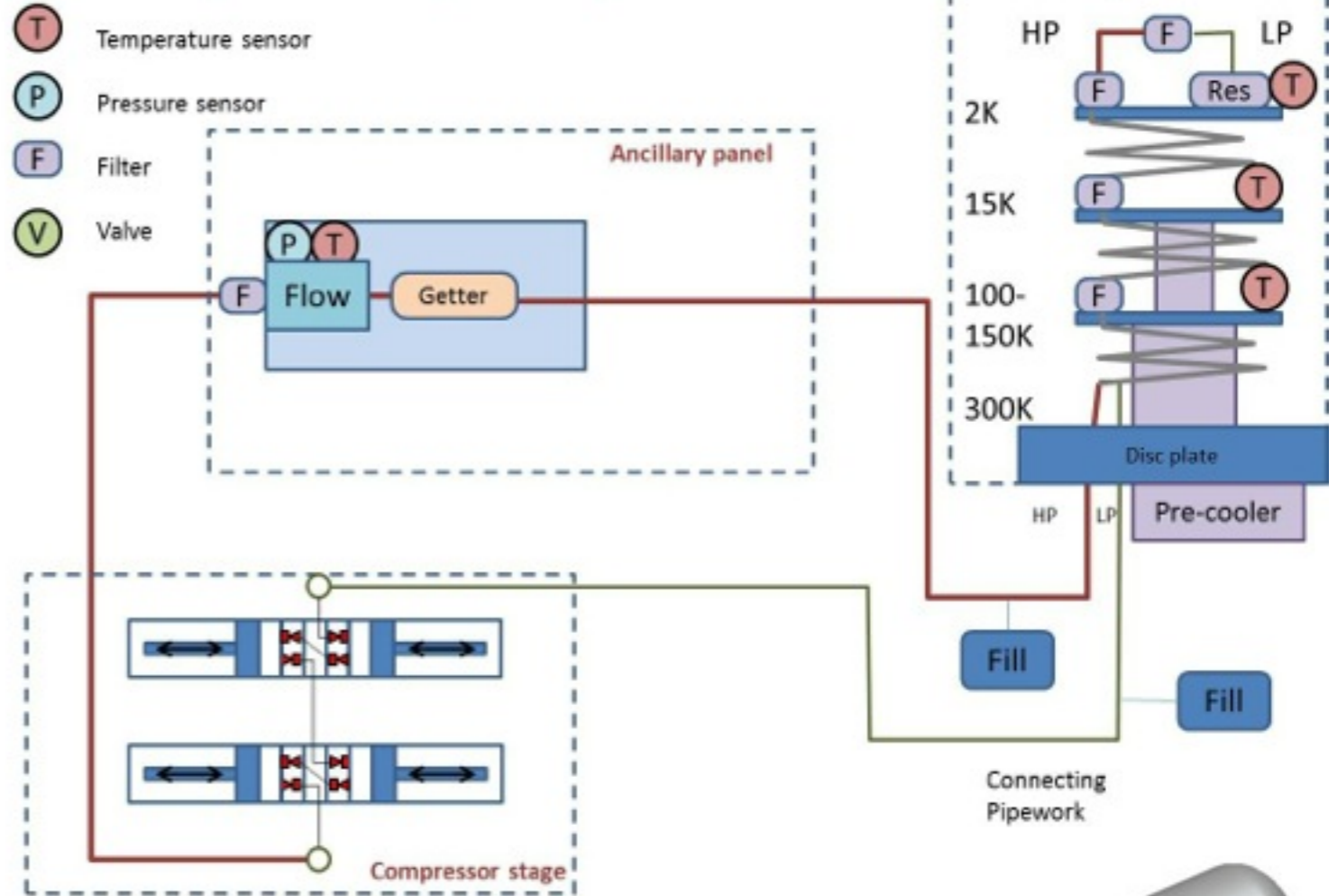


RAL - Astrium 4K JT Loop onboard PLANCK (HFI Instrument)

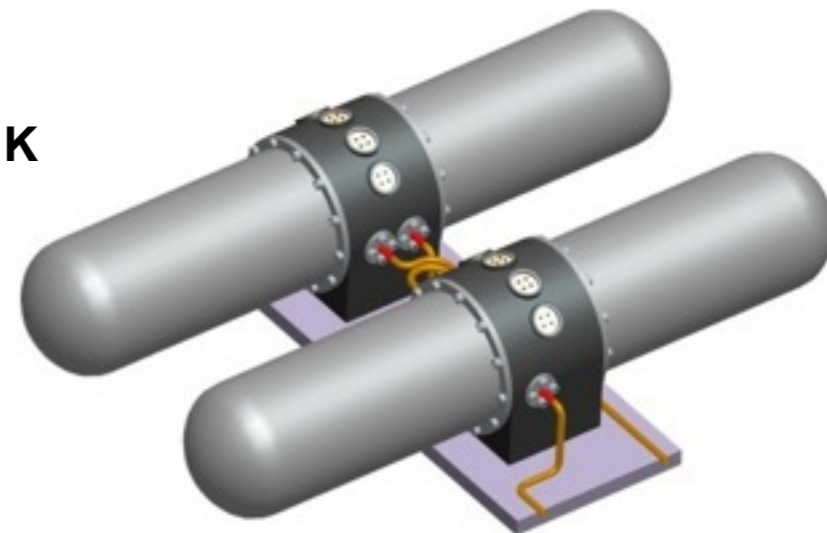
ADVANCED 2 K COOLER PROGRAM



Cooler layout 4 stage Compressor



Objective of 20 mW @ 2 K

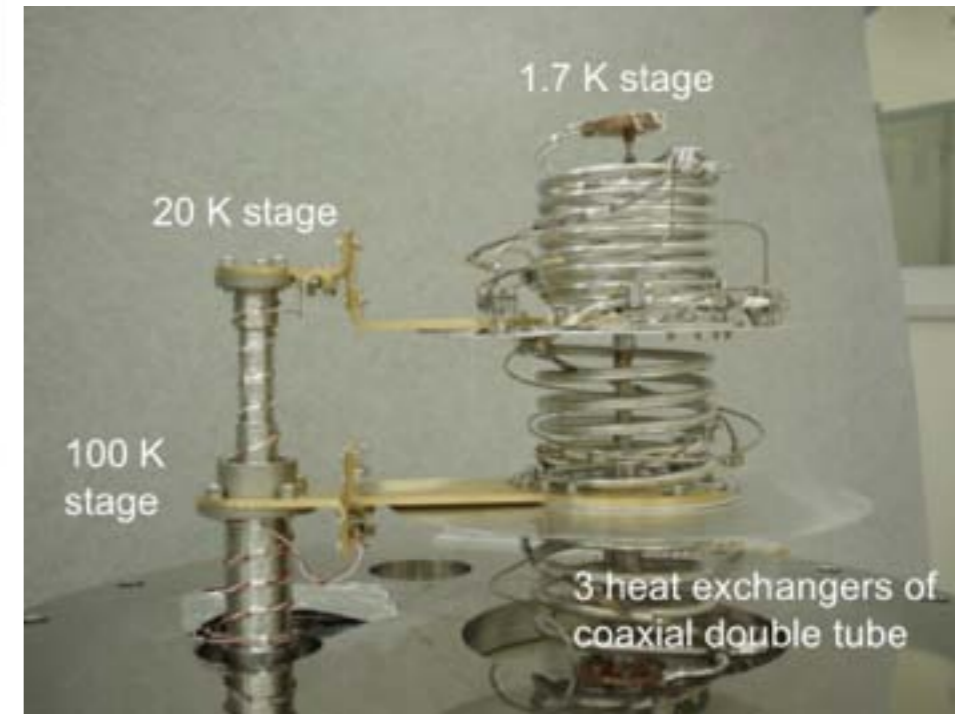


ATHENA Mission

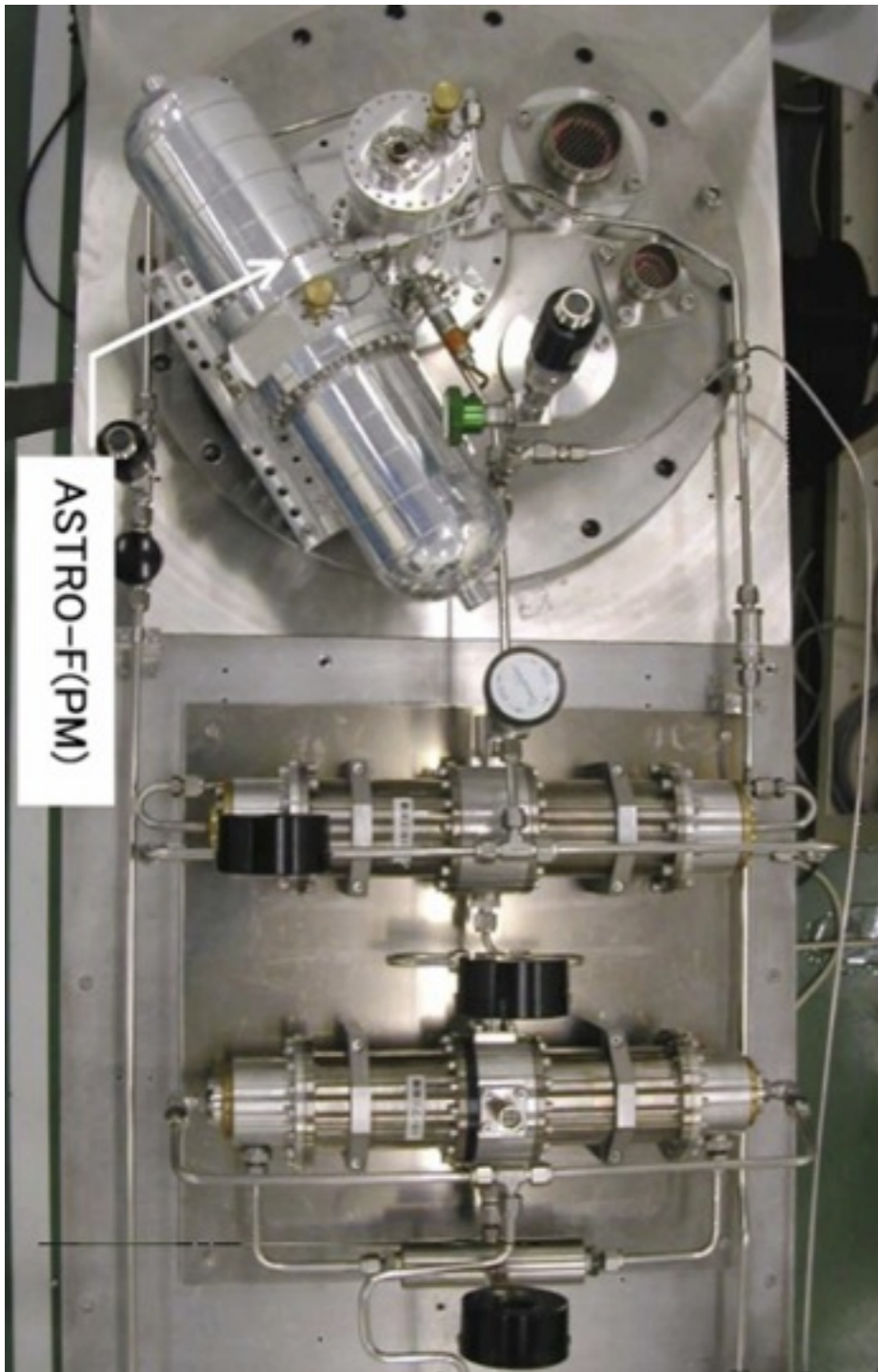
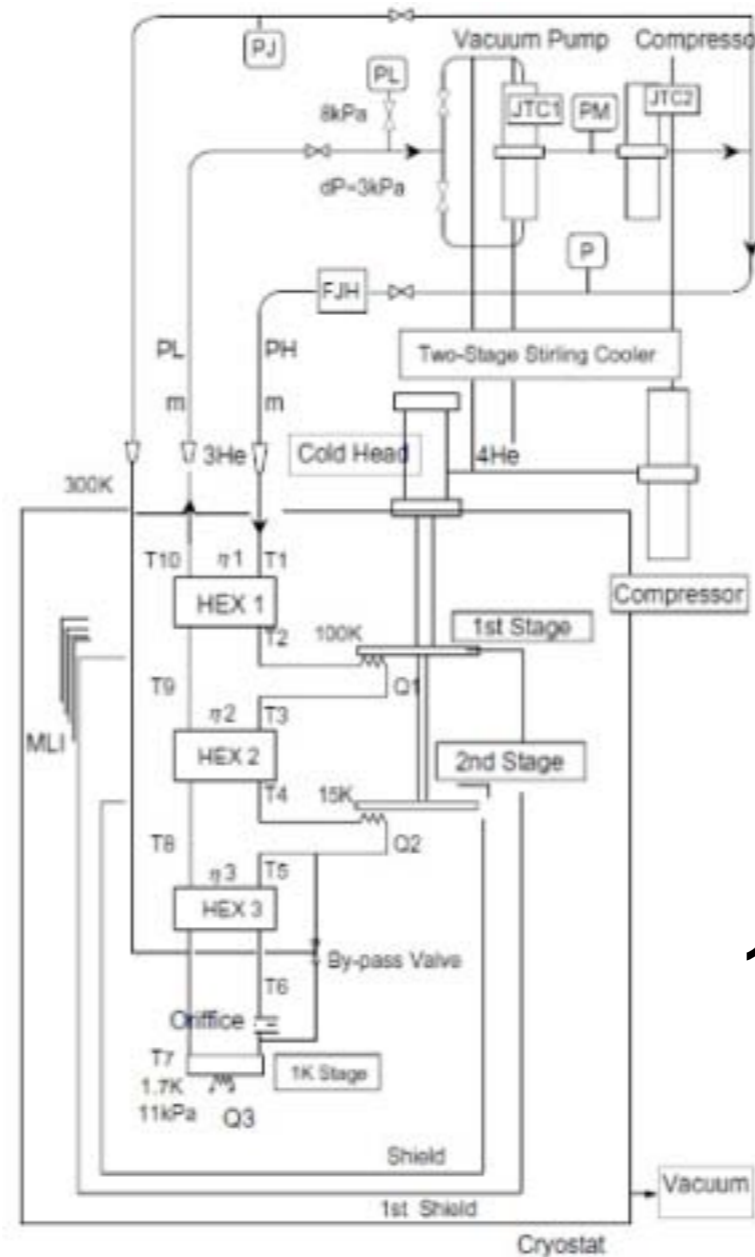
Possible application

JT SYSTEM - JAXA (ASTRO-F, SPICA, ...)

Prototype cooler



16 mW @ 1.7 K with 160 W input

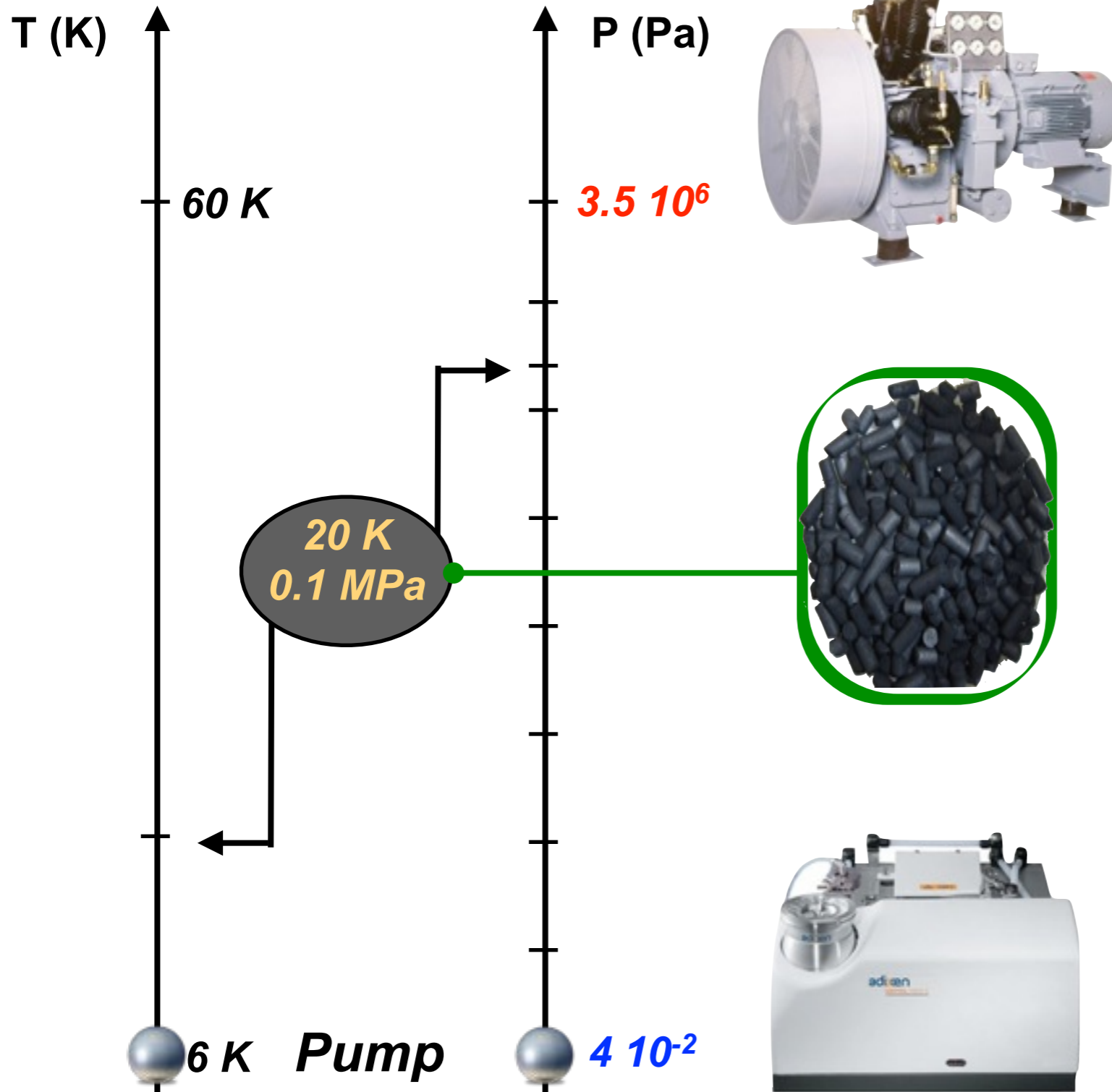


JT compressors
 (4 stages: 8 kPa to 0.7 MPa)
 (ratio 87)

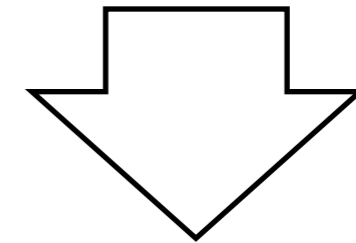
**EM unit (with 4 compressors) under fabrication
 (SPICA/SAFARI mission)**

VIBRATIONLESS ? ALTERNATIVE COMPRESSOR

Compressor



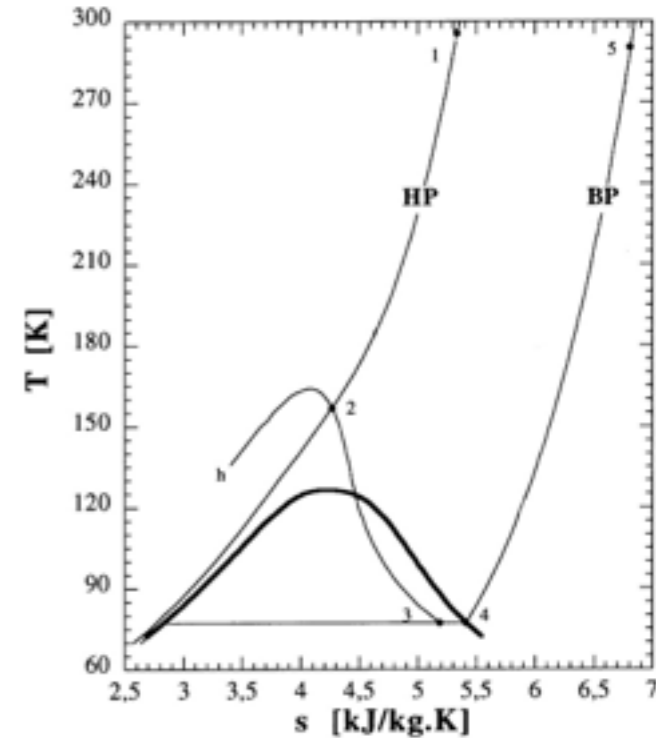
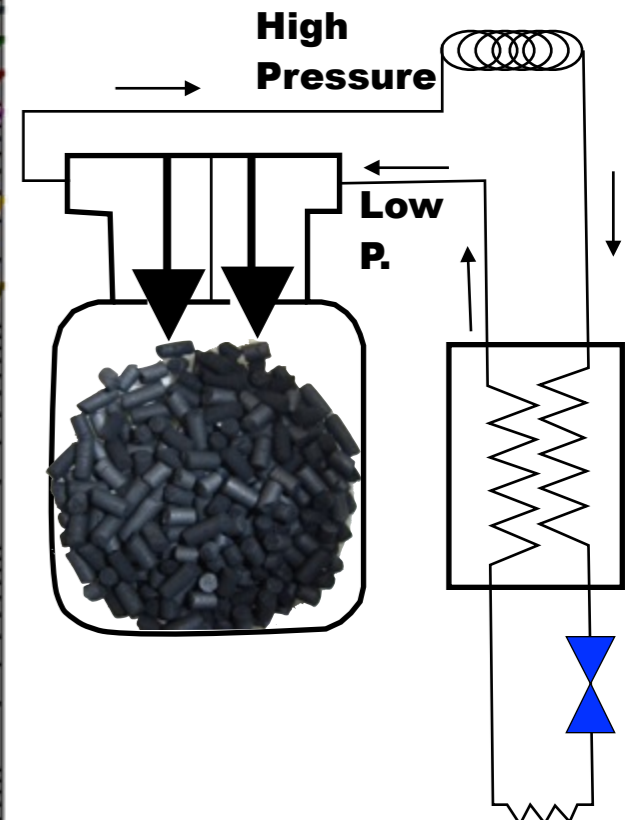
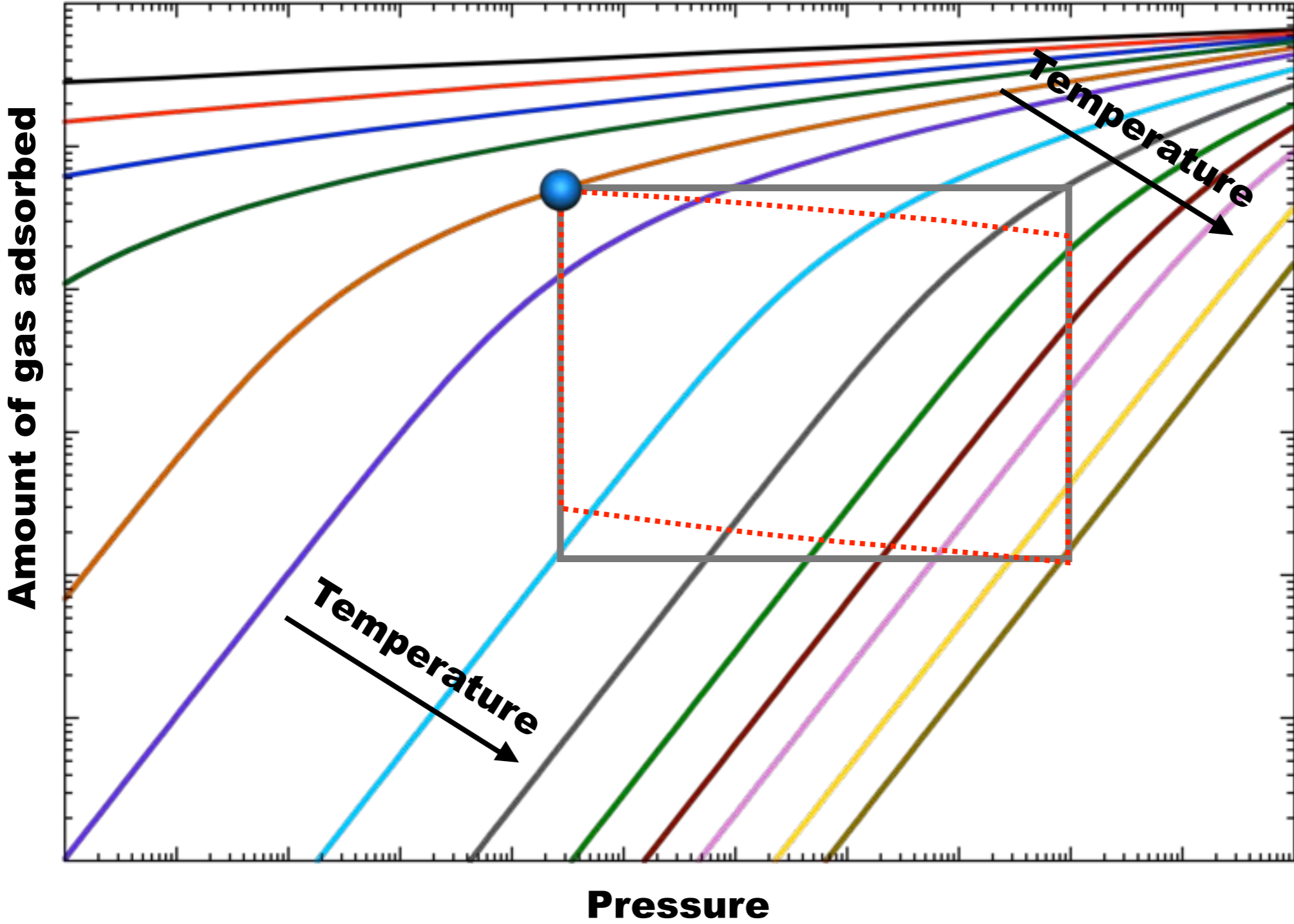
1 order of magnitude in T



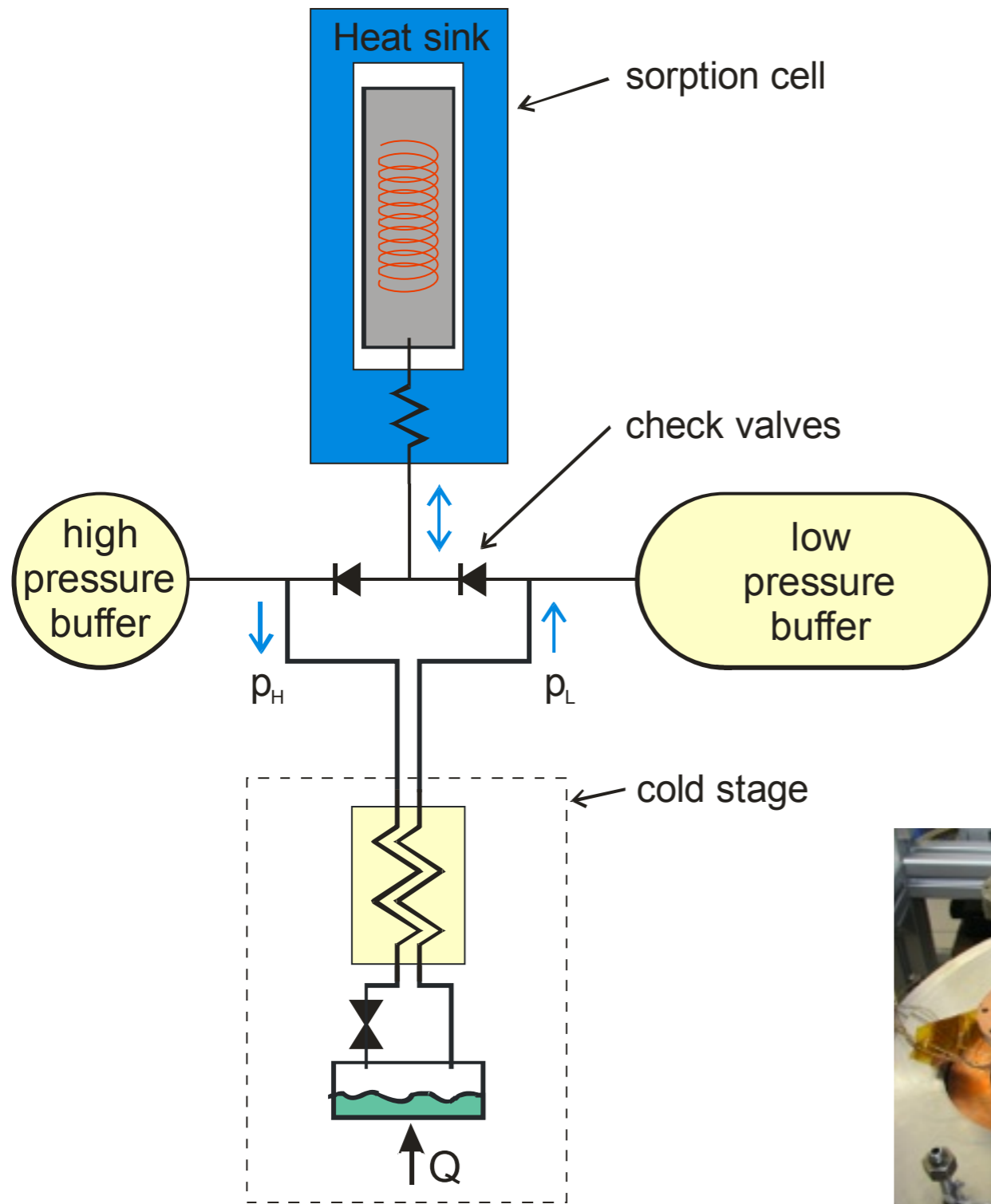
8 orders of magnitude in P

*No mechanical equivalent
 (in 1 stage)*

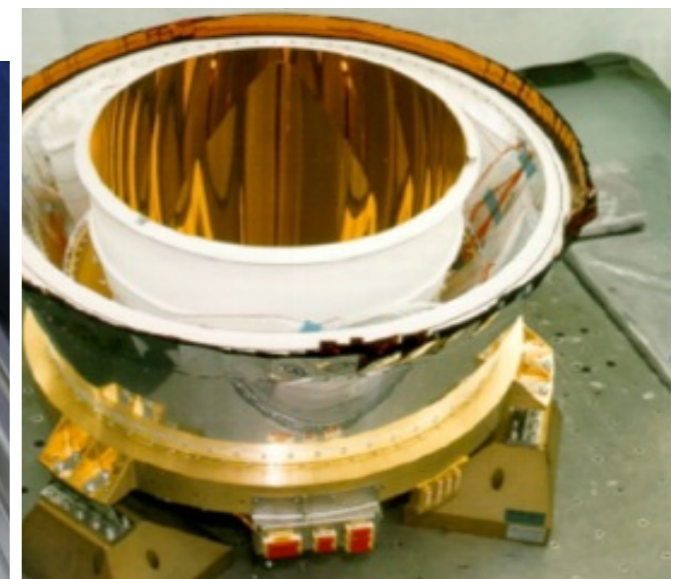
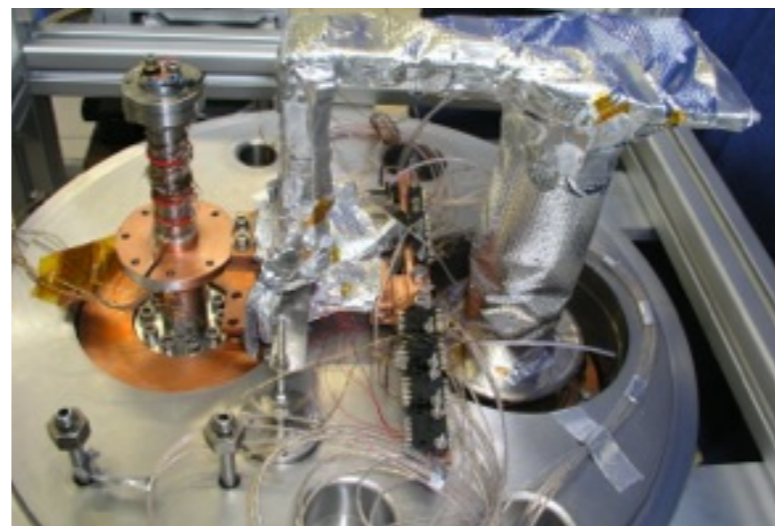
JOULE THOMSON WITH SORPTION COMPRESSOR



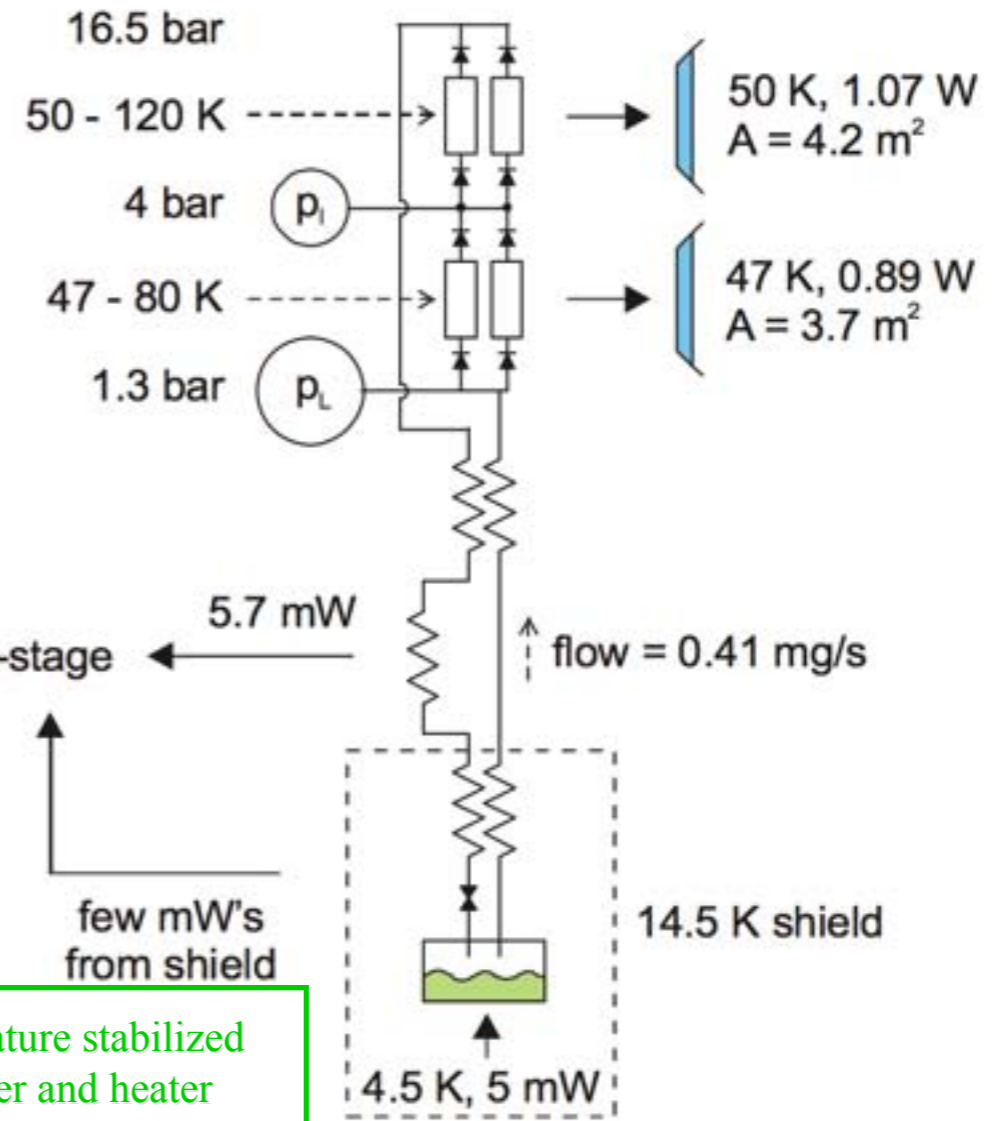
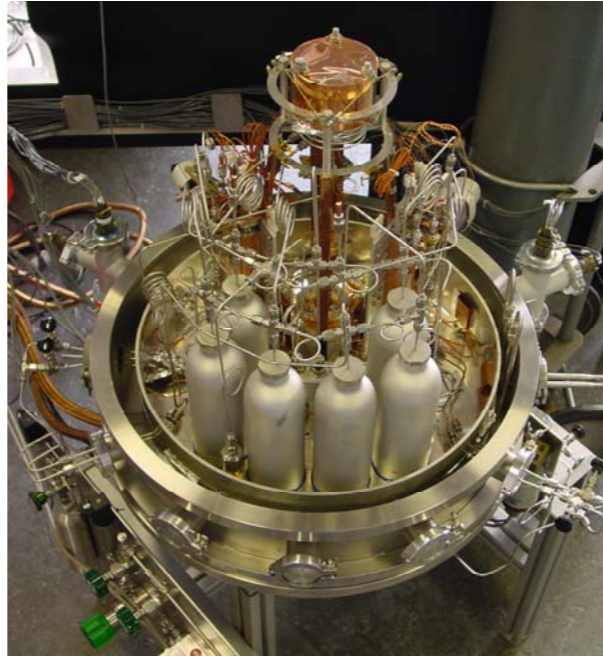
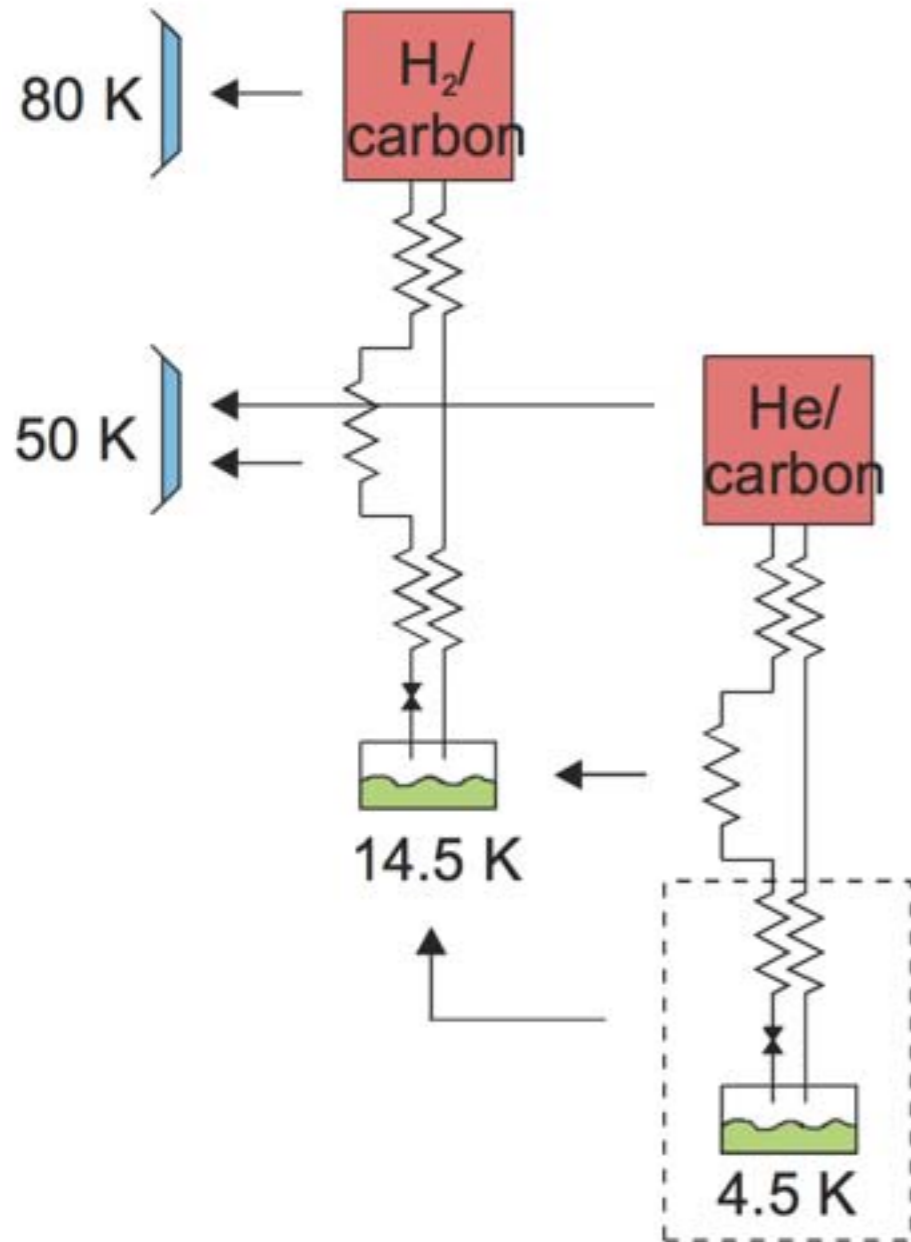
HELIUM & HYDROGEN JT COOLER - VIBRATION FREE



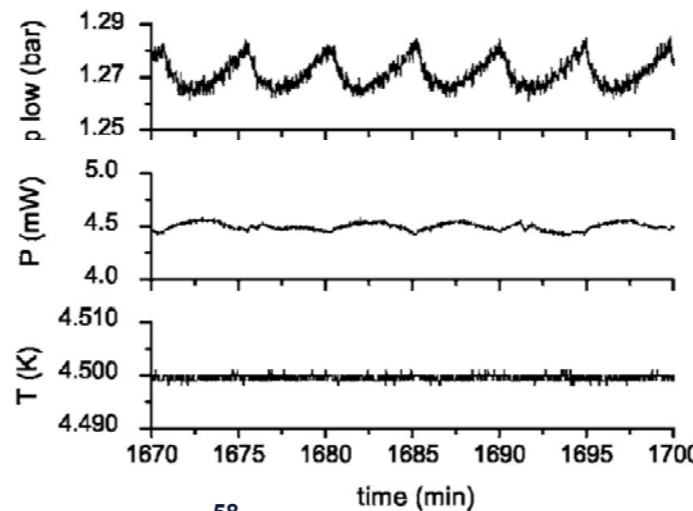
4.5 K helium cooler (ICC14, 2006)



14.5 K hydrogen cooler (ICC17, 2012)



cold tip temperature stabilized by PID controller and heater

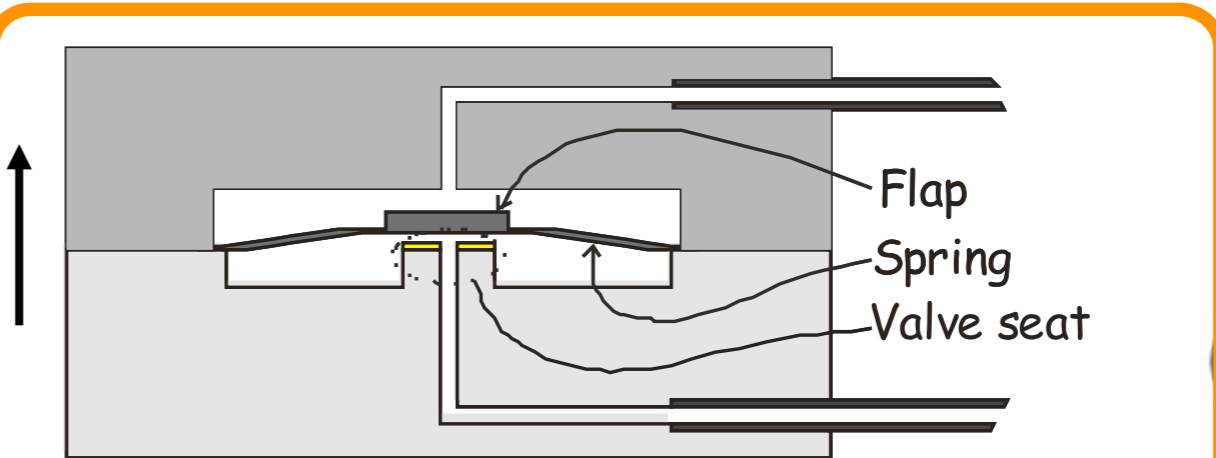
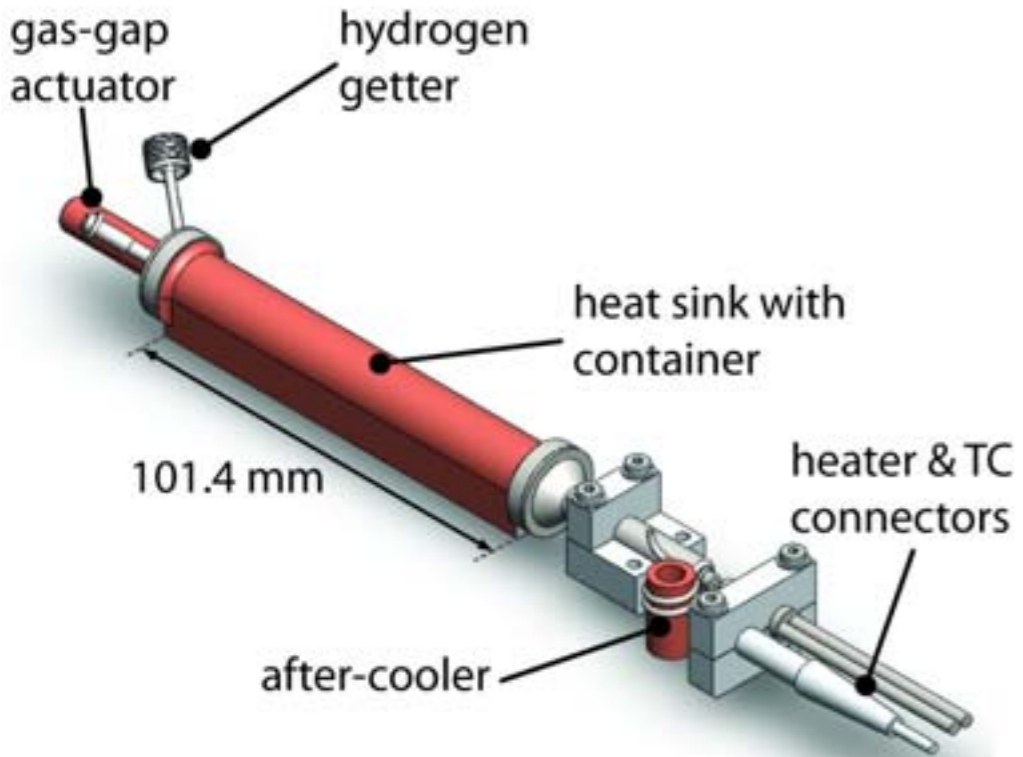
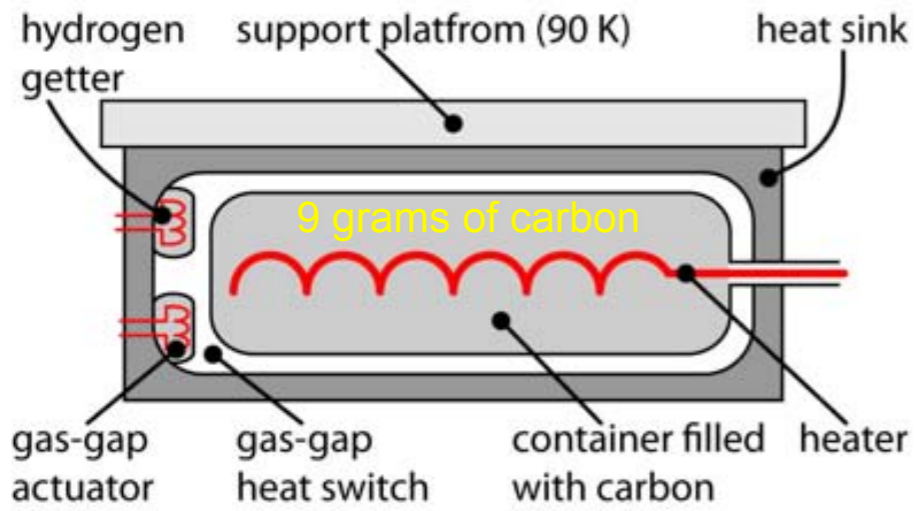


Typical results (Lab prototype)

4.5 mW @ 4.5 K
 $T_{stability} \approx 1 \text{ mK}$



SMALL IS BEAUTIFUL

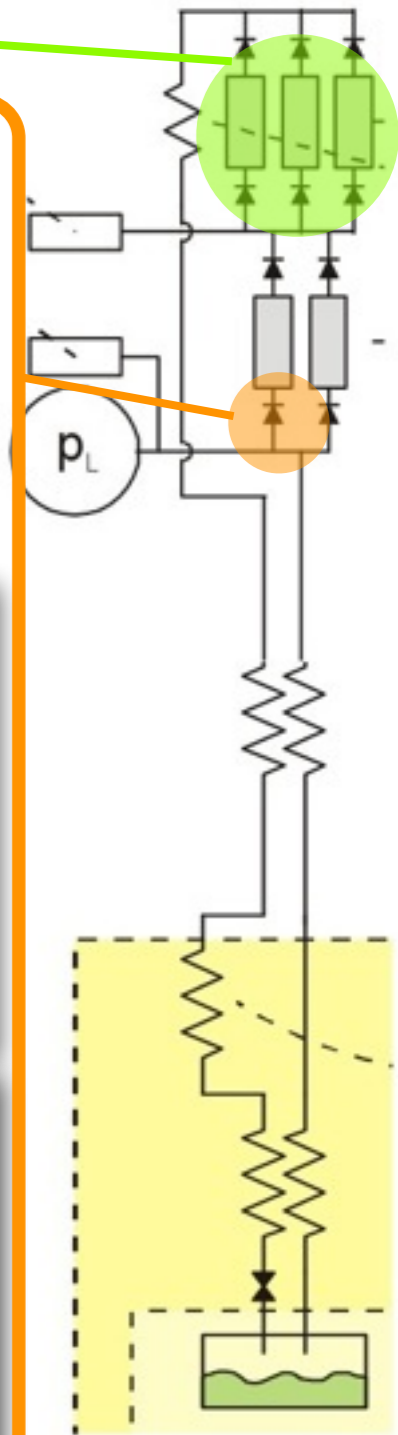
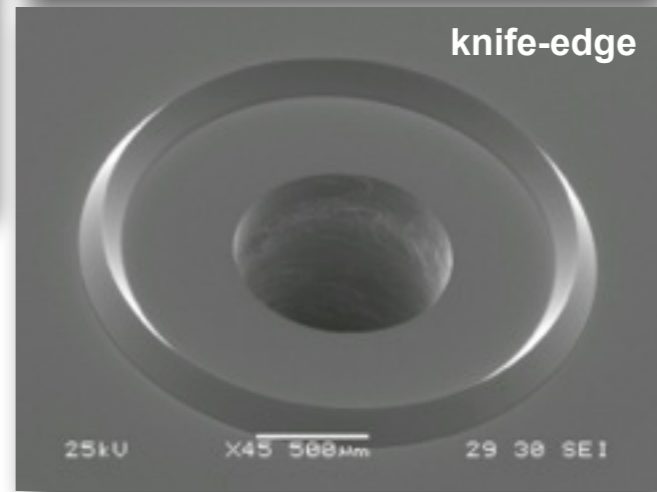
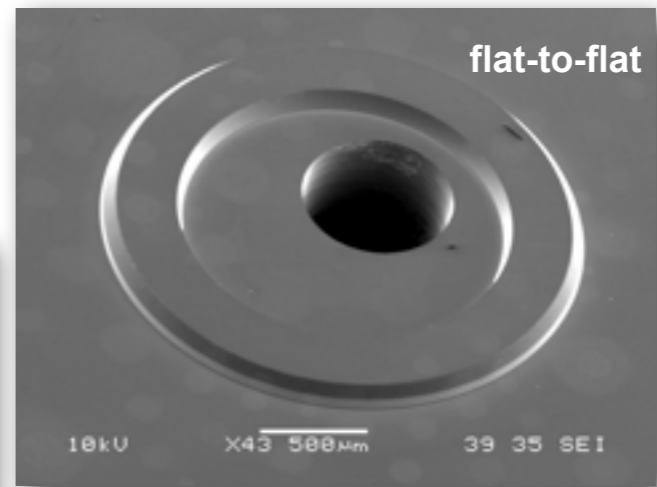


flat-to-flat seal
3 bar to 0.1 bar

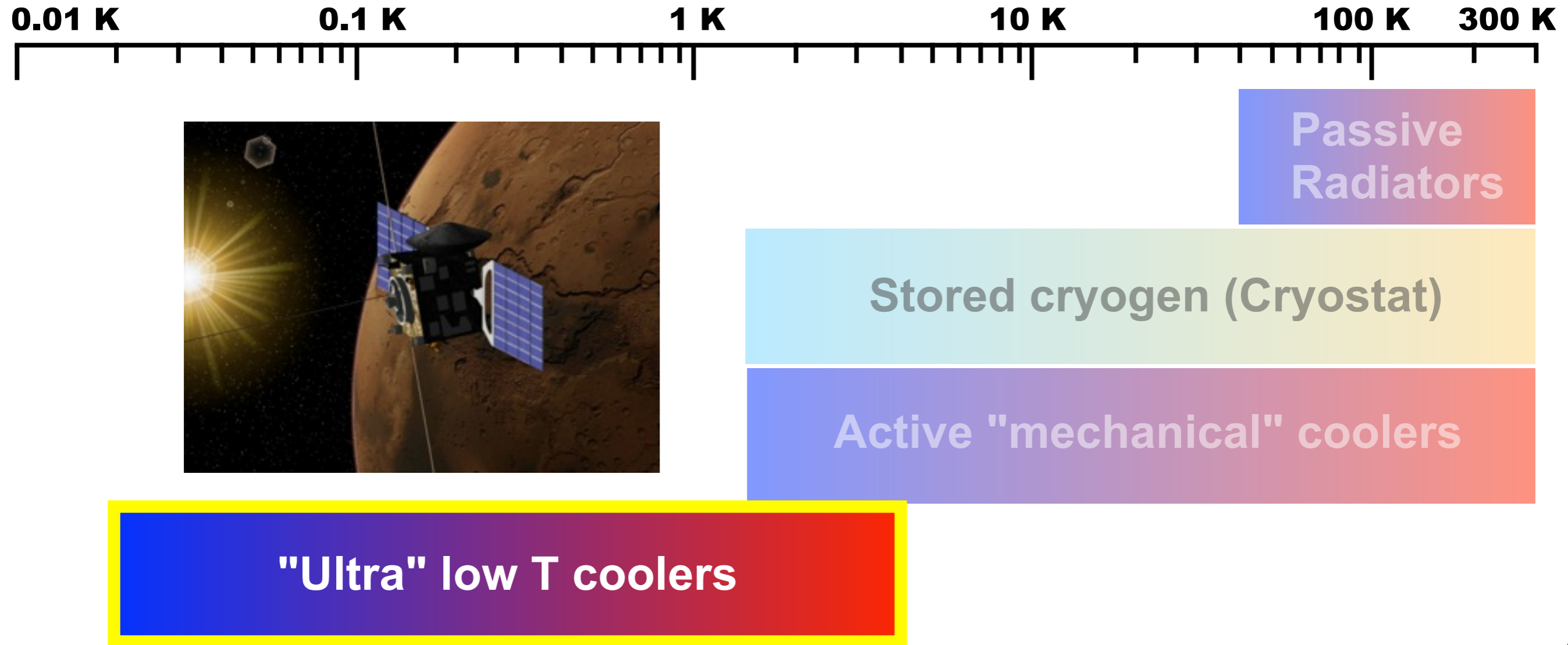


flap: Au layer 10 μm

knife-edge seal
50 bar to 3 bar



SPACE COOLING CHAINS



PAST (LAST DECADE) & FUTURE SUBKELVIN MISSIONS



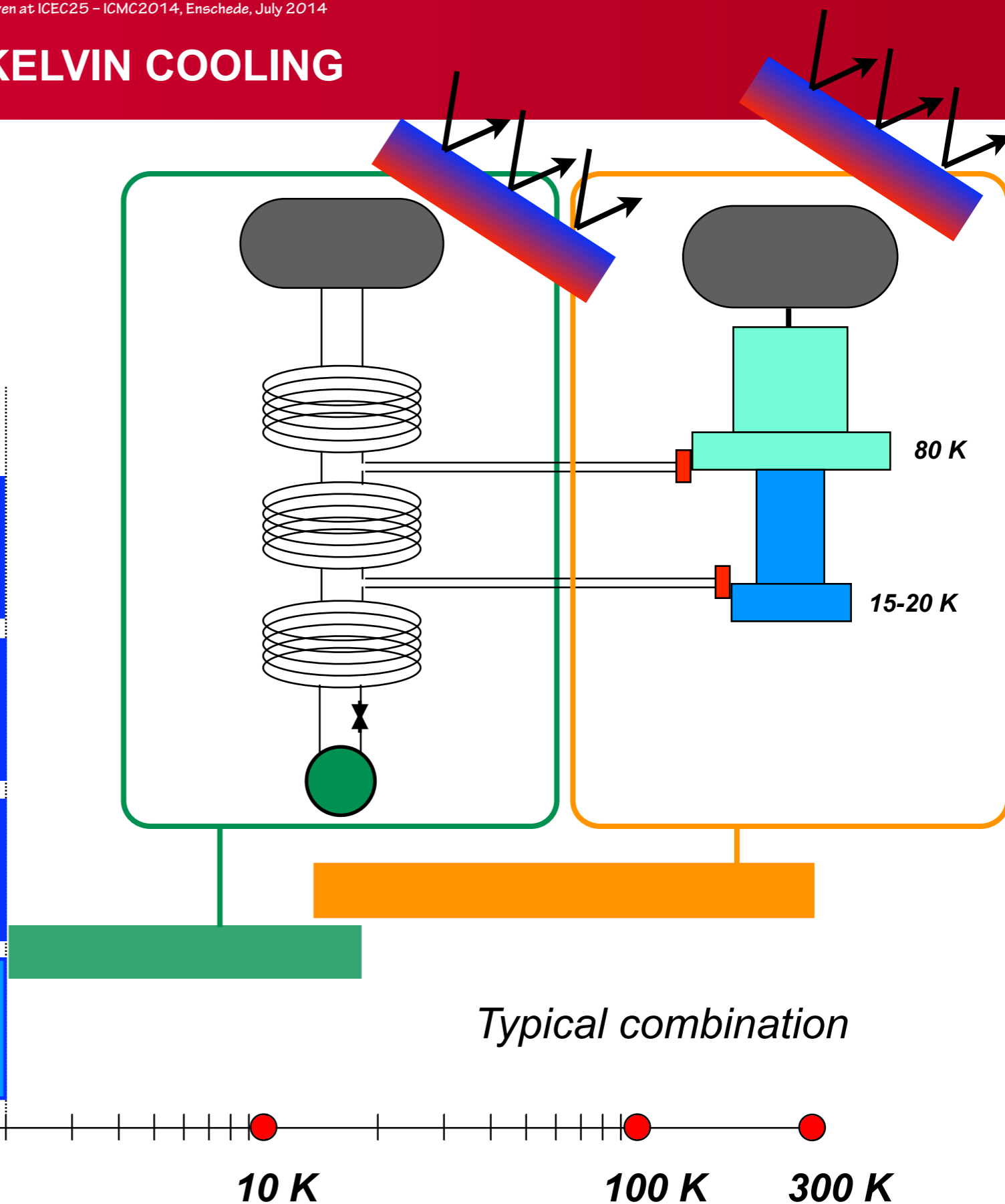
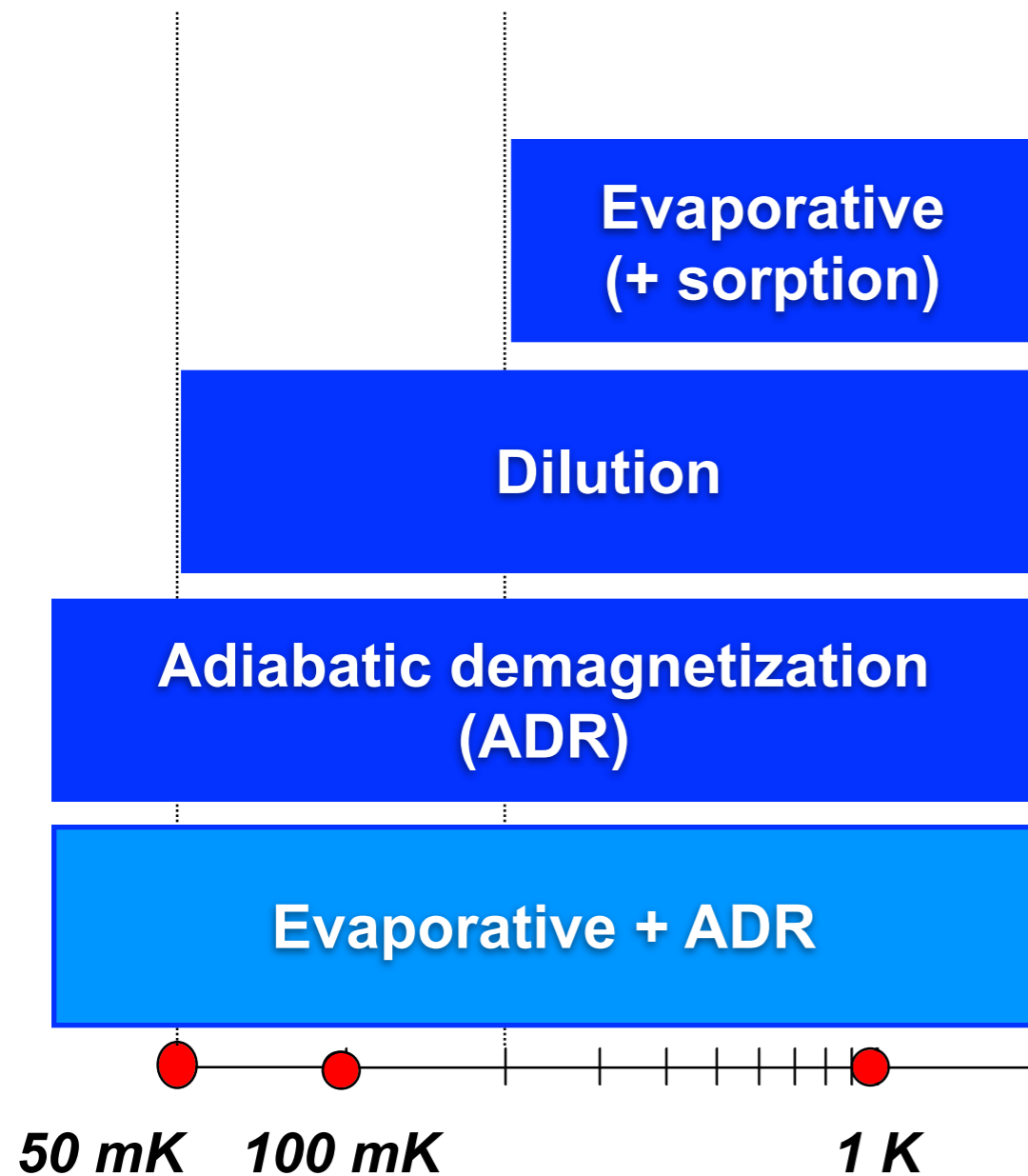
10 mK

100 mK

1 K

SUBKELVIN COOLING

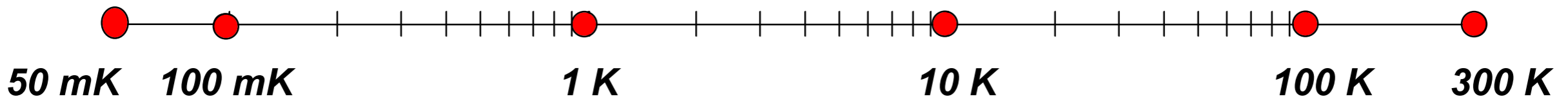
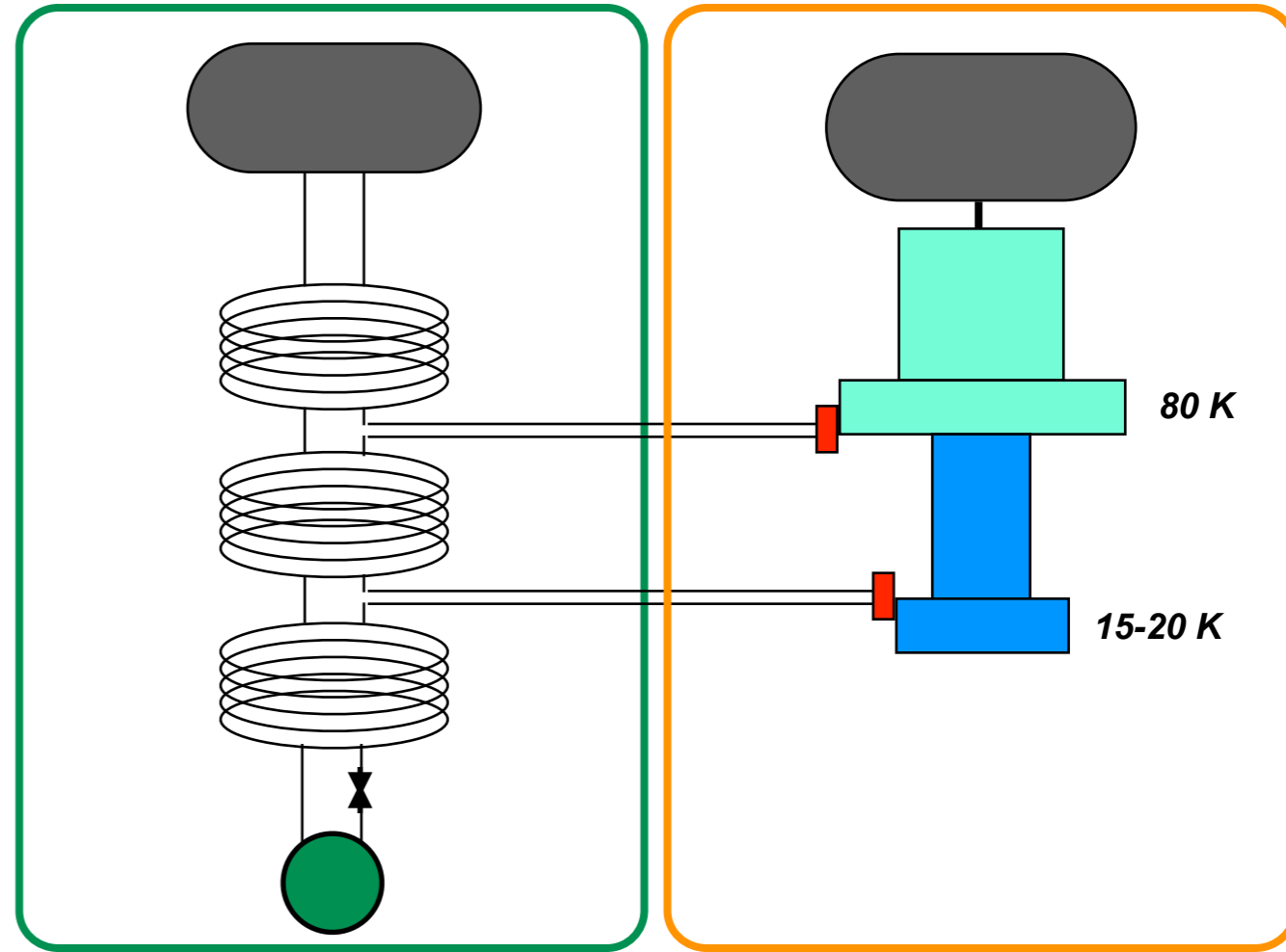
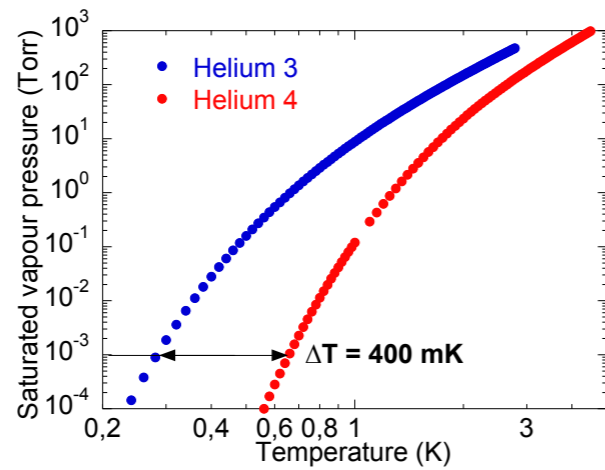
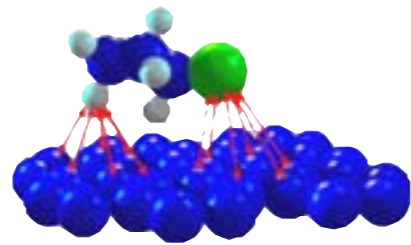
Only 3 (4) options



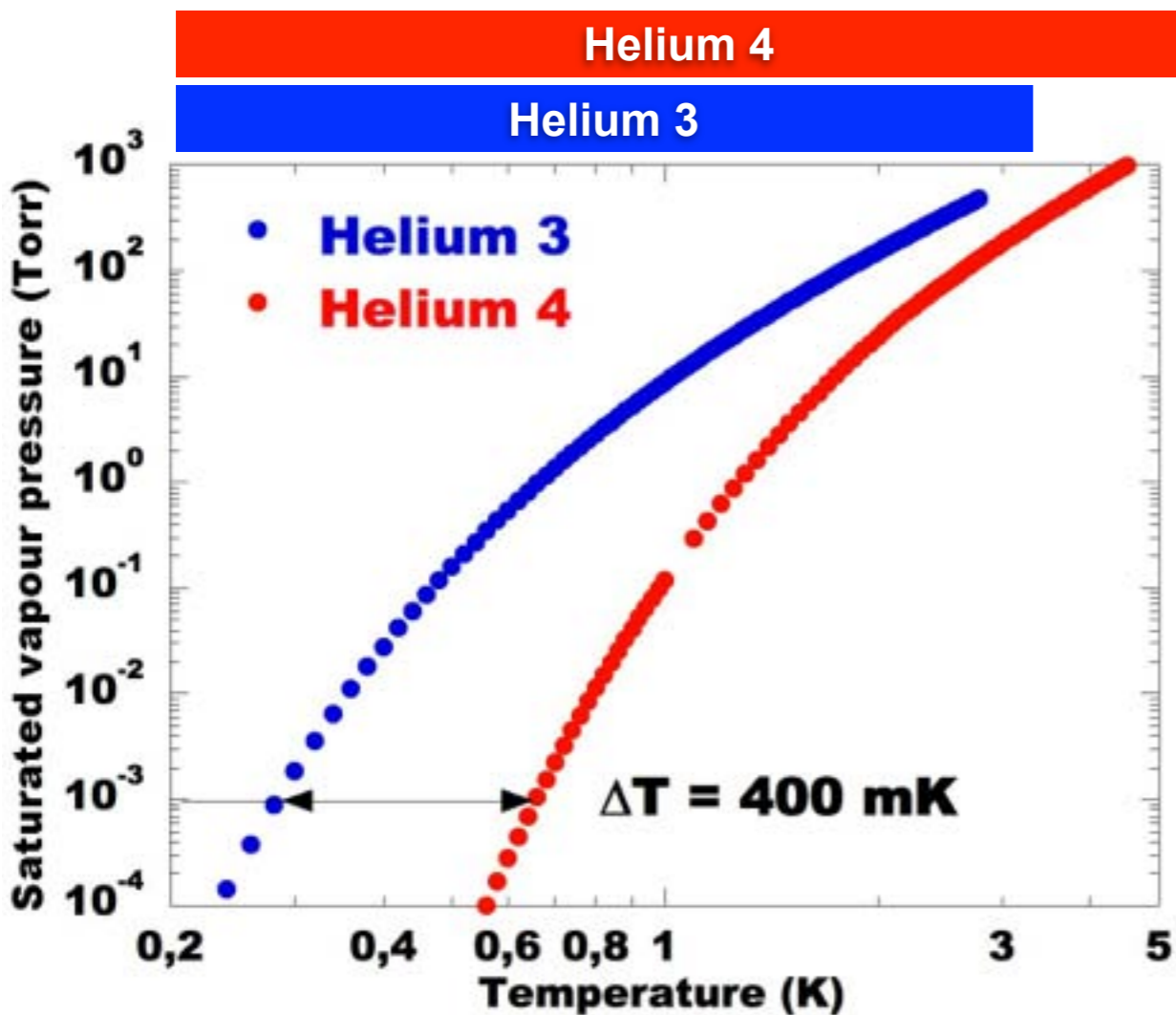
Typical combination

³HE EVAPORATIVE COOLING

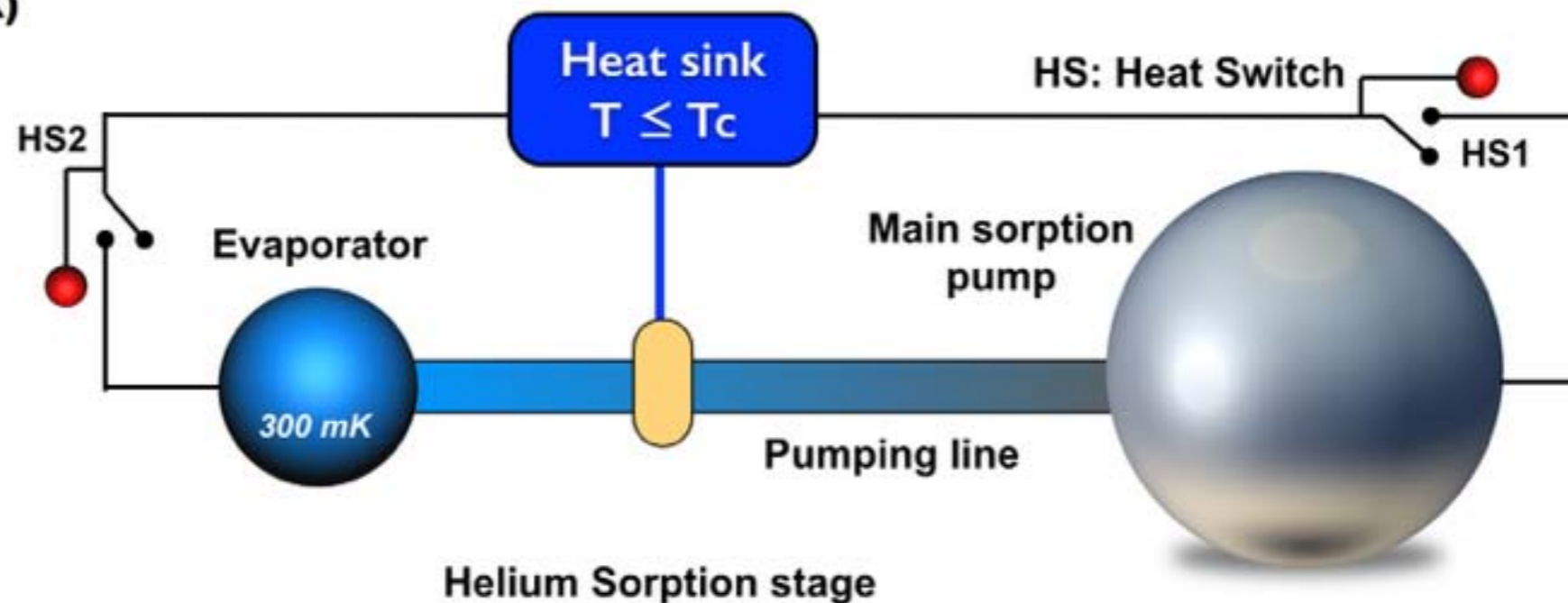
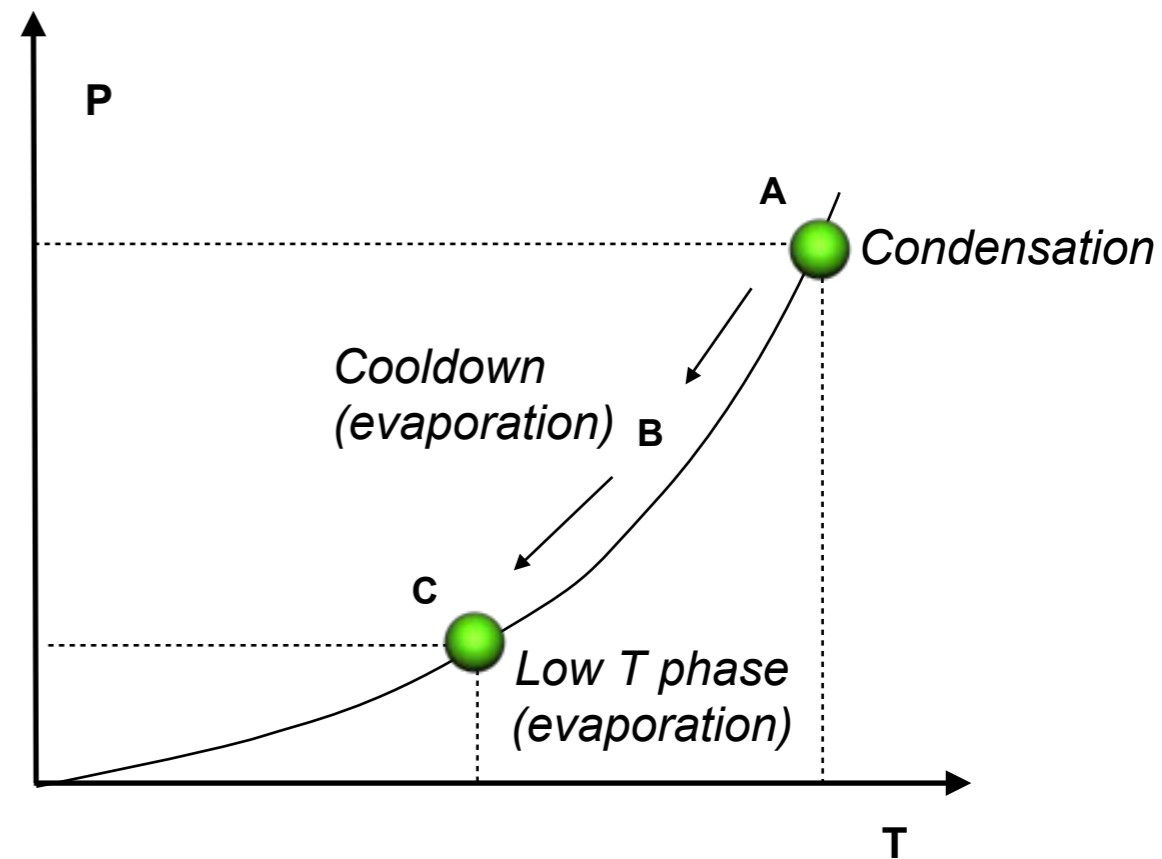
Evaporative cooling



HELIUM EVAPORATIVE COOLING



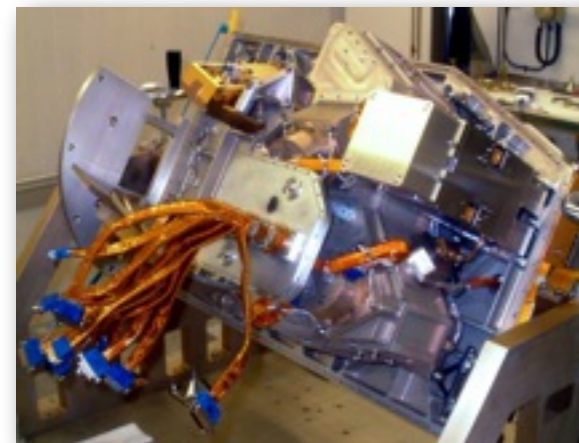
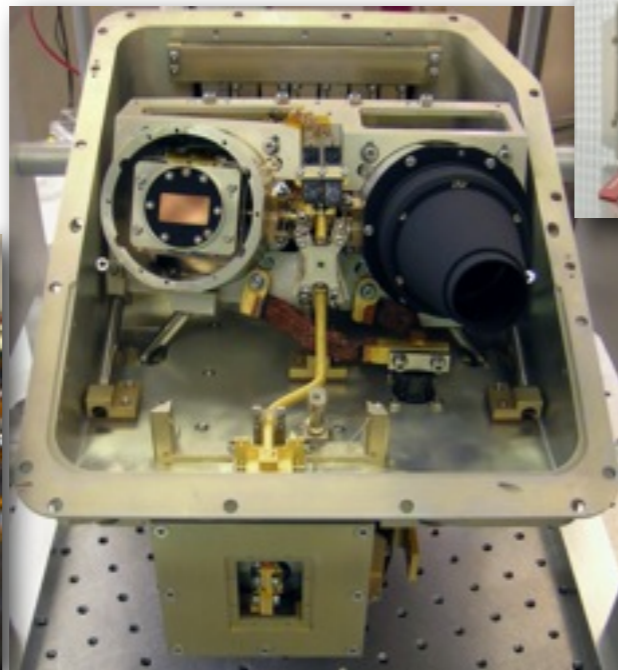
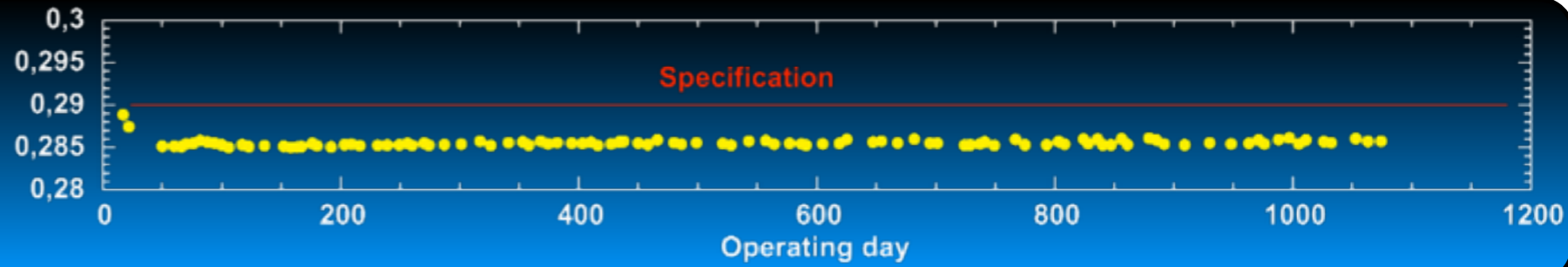
T_{critical}



HERSCHEL

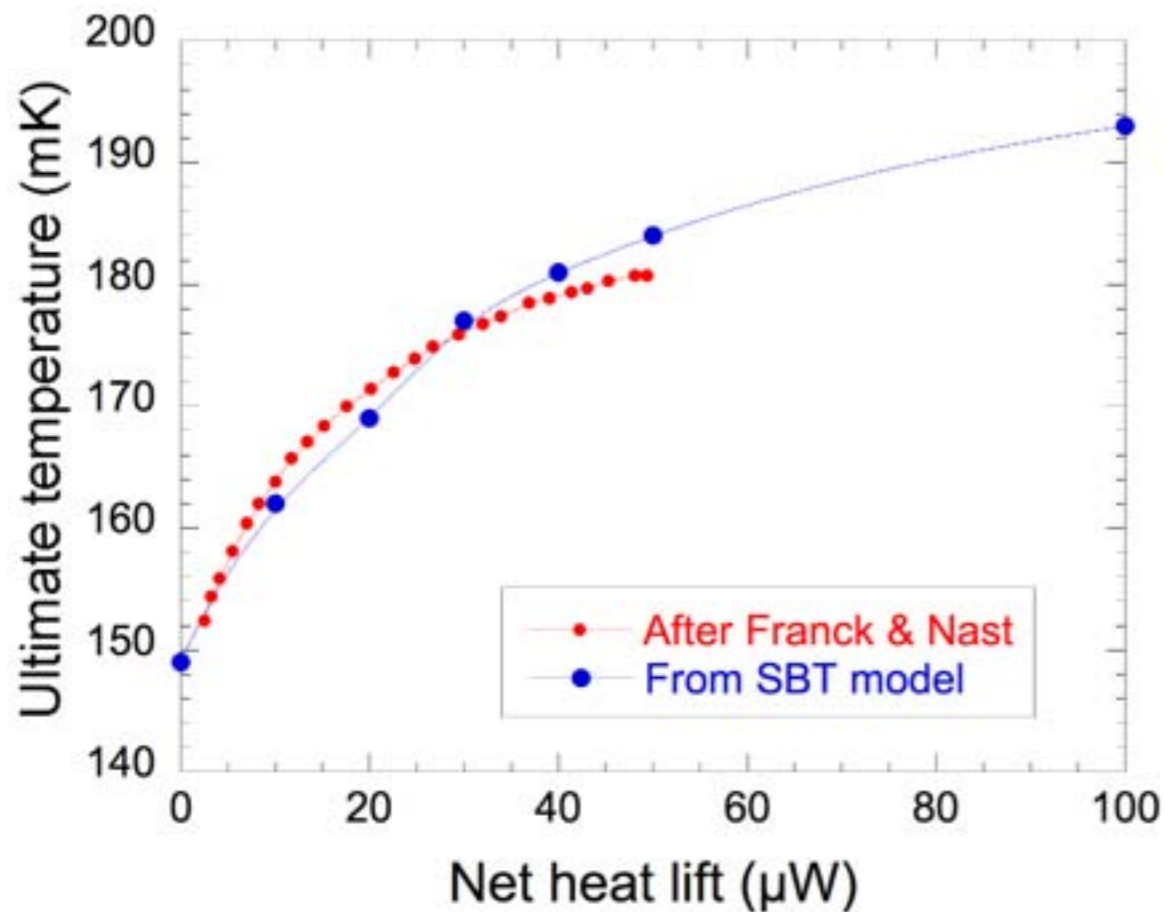
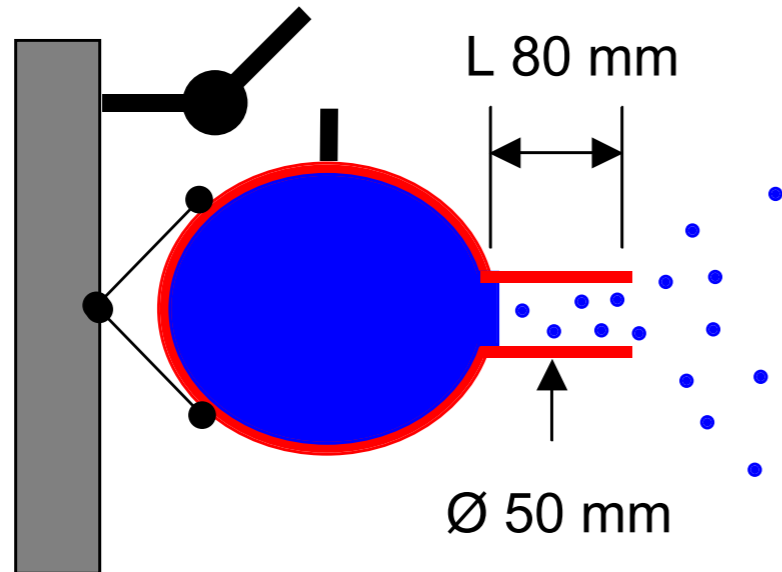
HERSCHEL

SPIRE Sorption unit

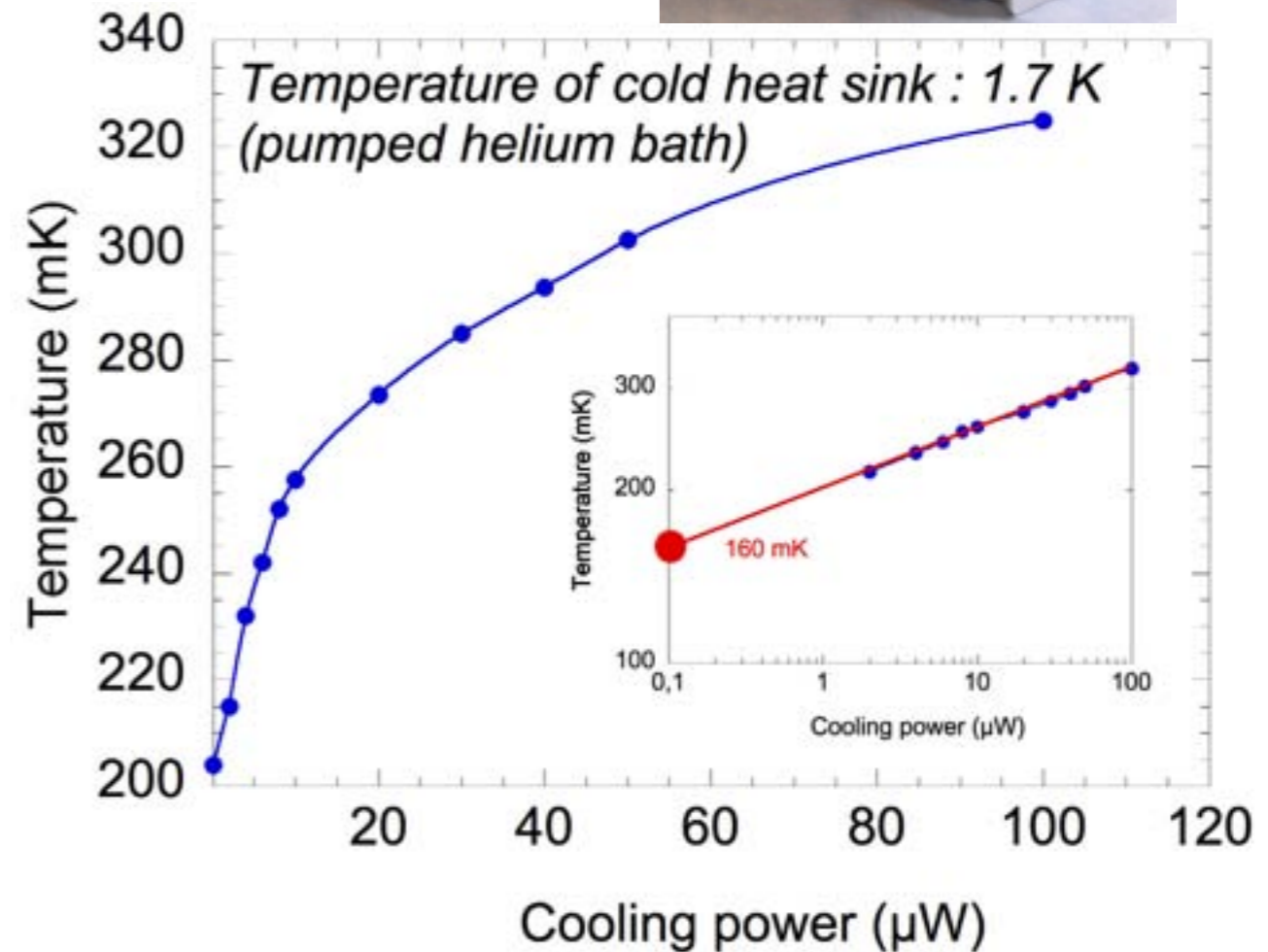


**Two units
(SPIRE and PACS Instruments)
3.8 years in orbit at L2**

3HE SORPTION COOLER: ULTIMATE T ?

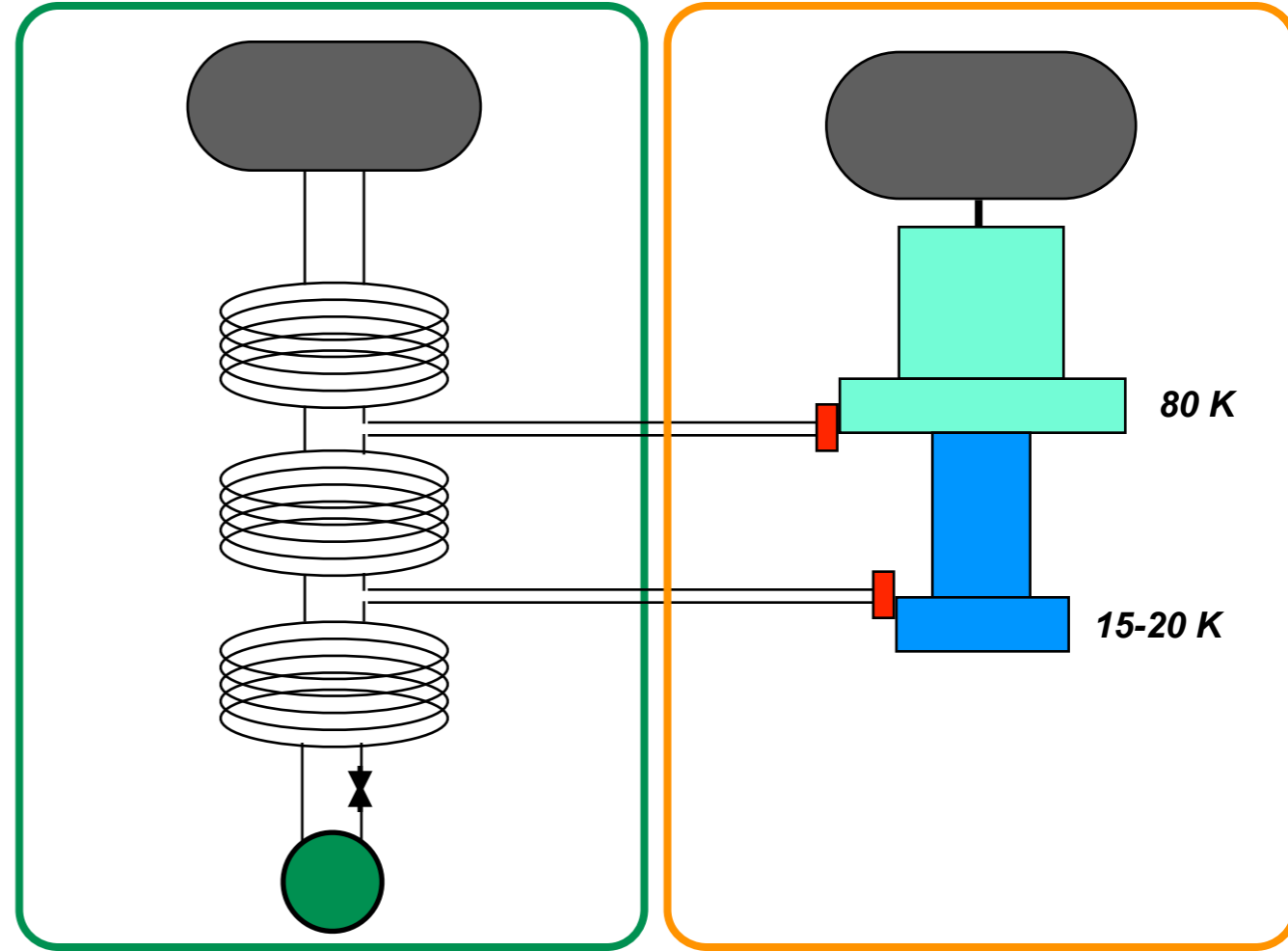
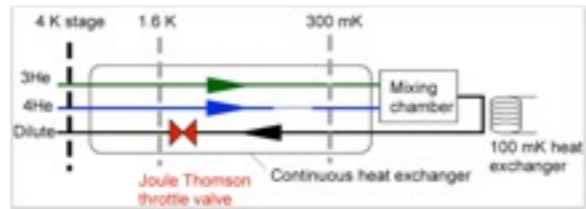


In practice
Multistage system
(⁴He, ³He, ³He)
Ground based



DILUTION COOLER

Dilution



50 mK 100 mK

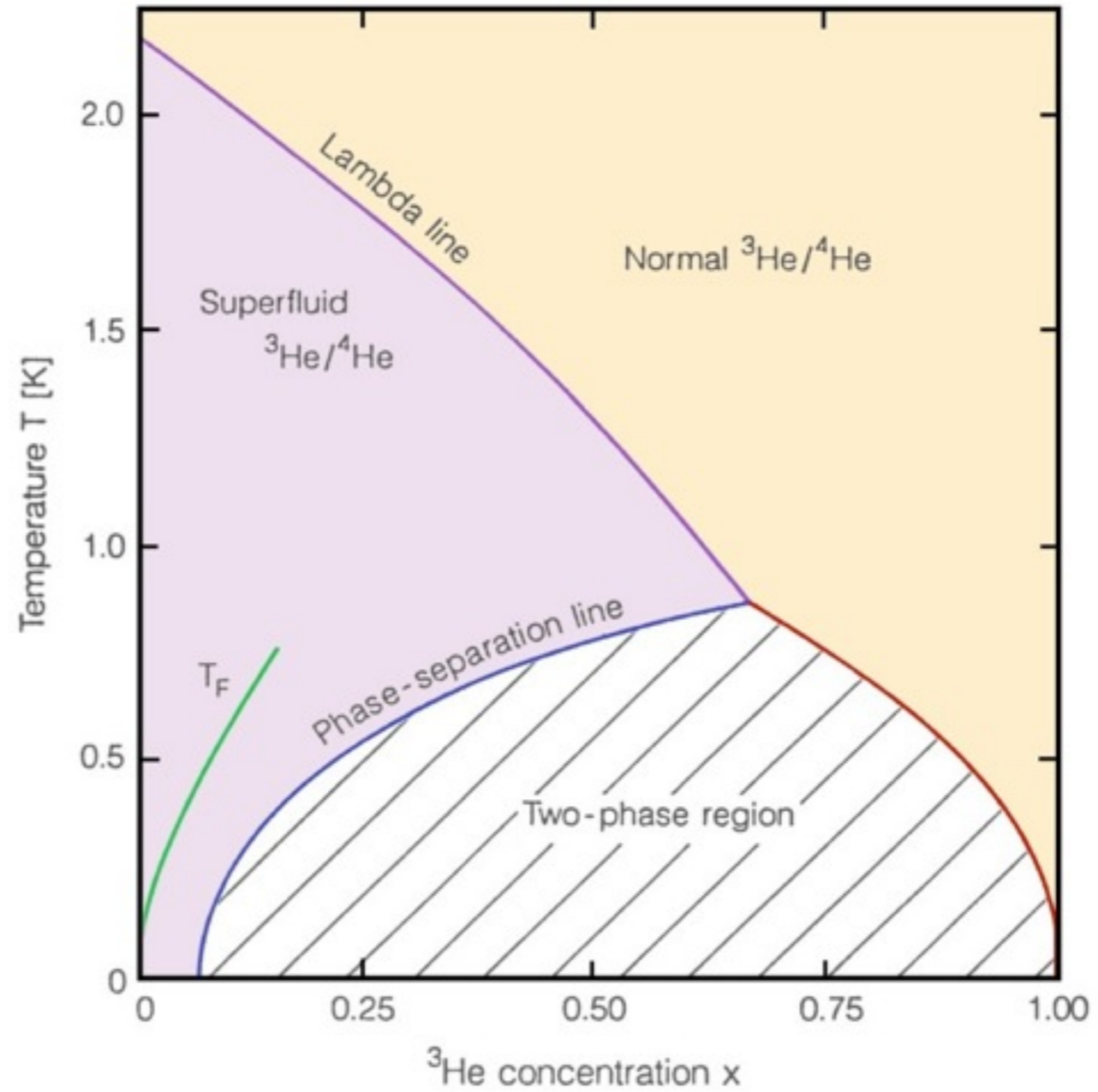
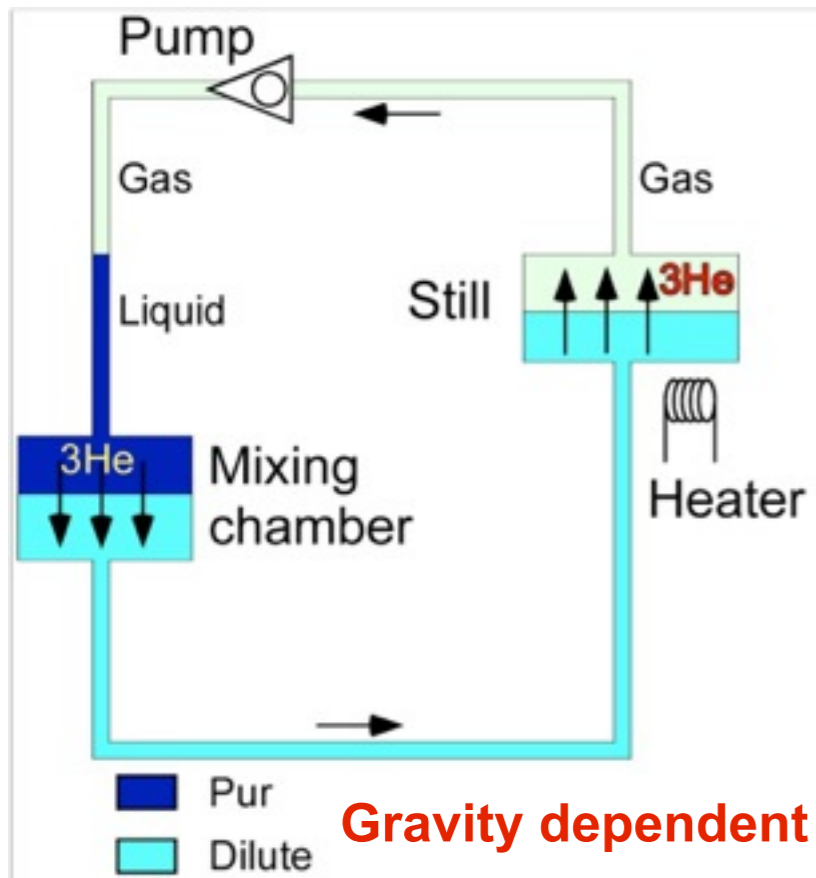
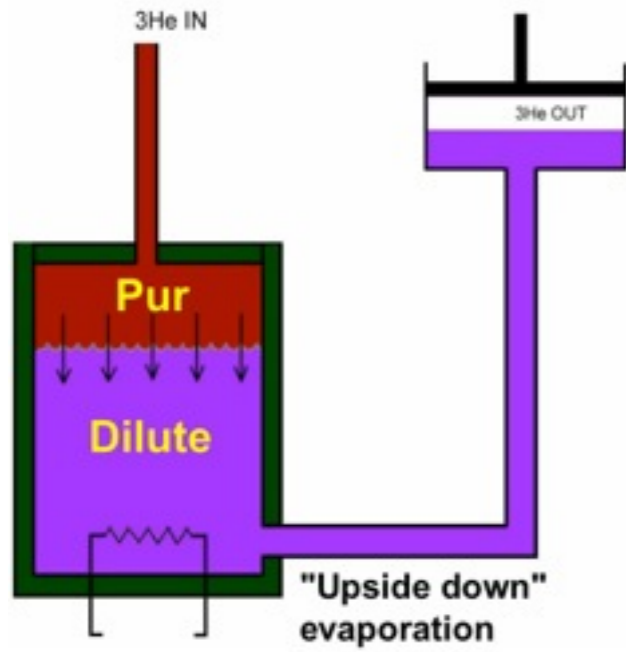
1 K

10 K

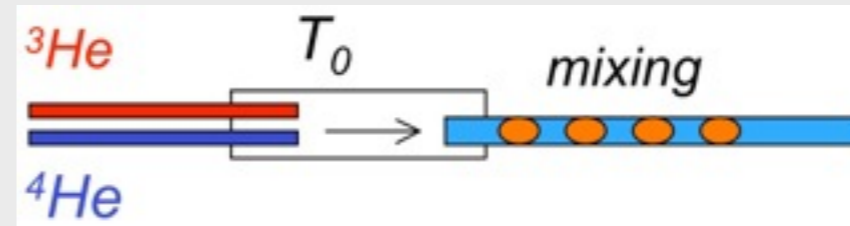
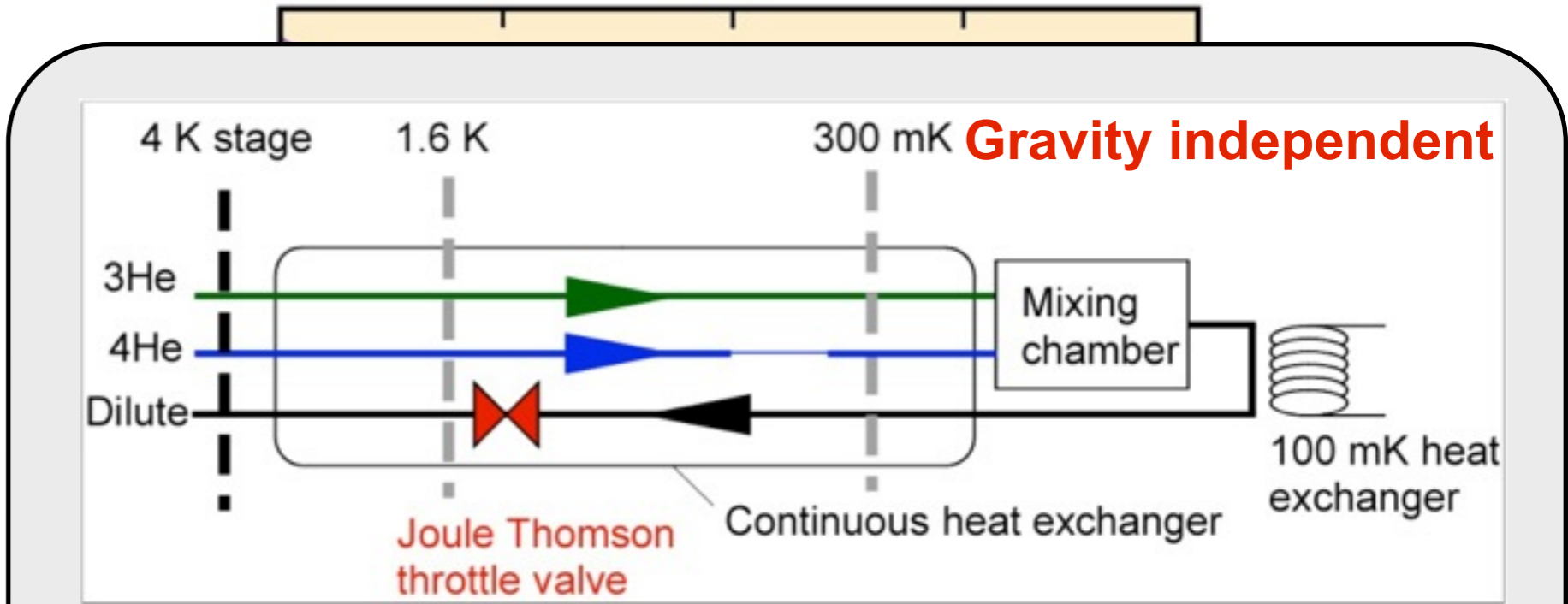
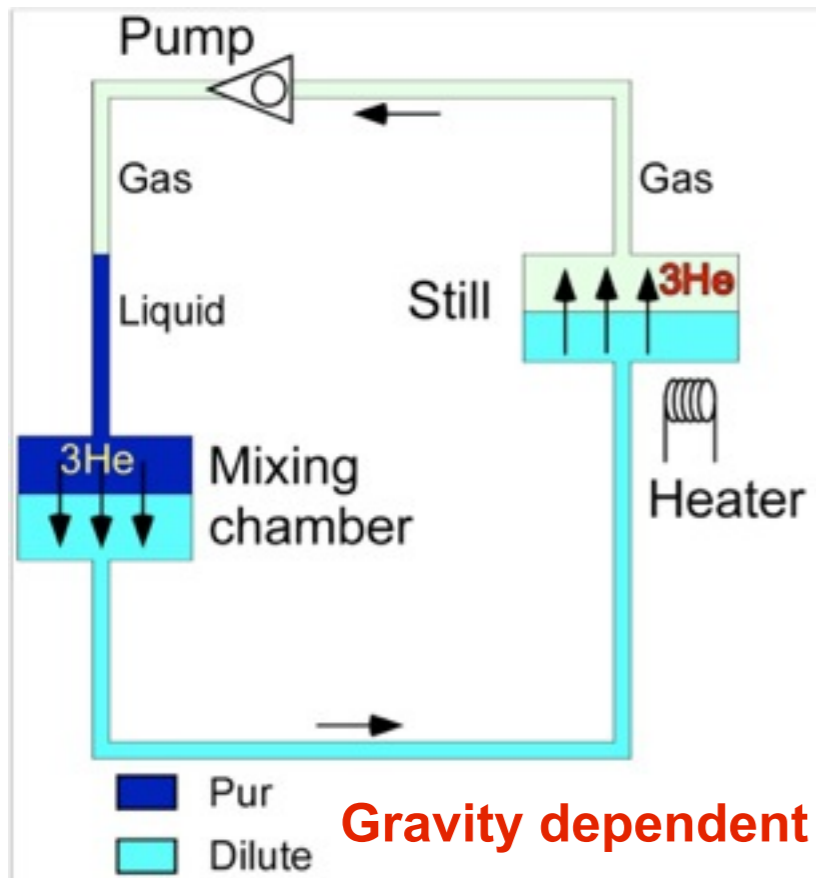
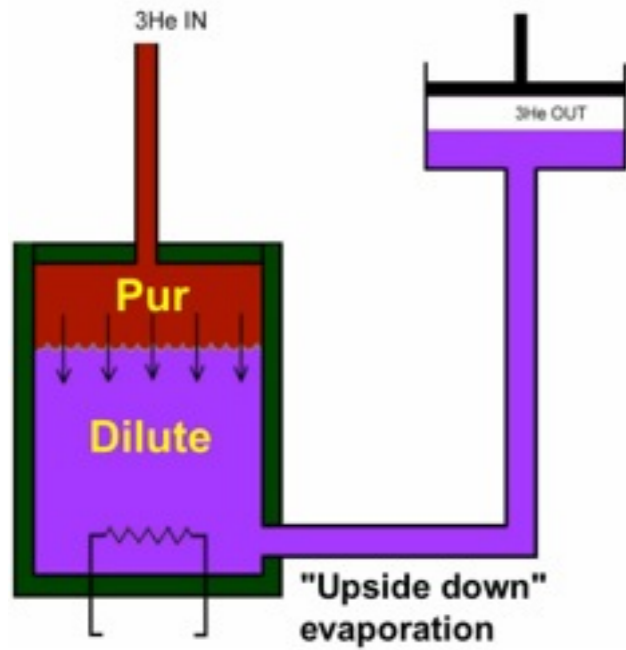
100 K

300 K

DILUTION

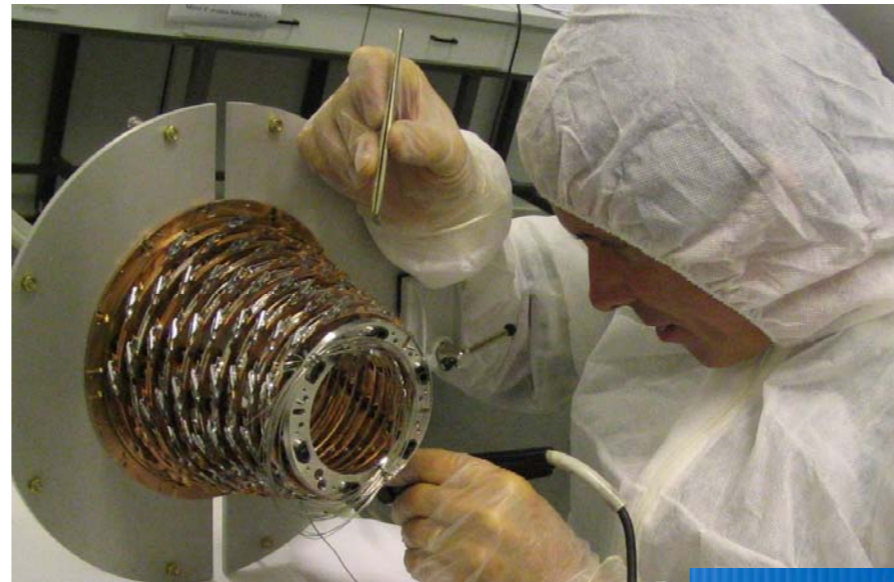
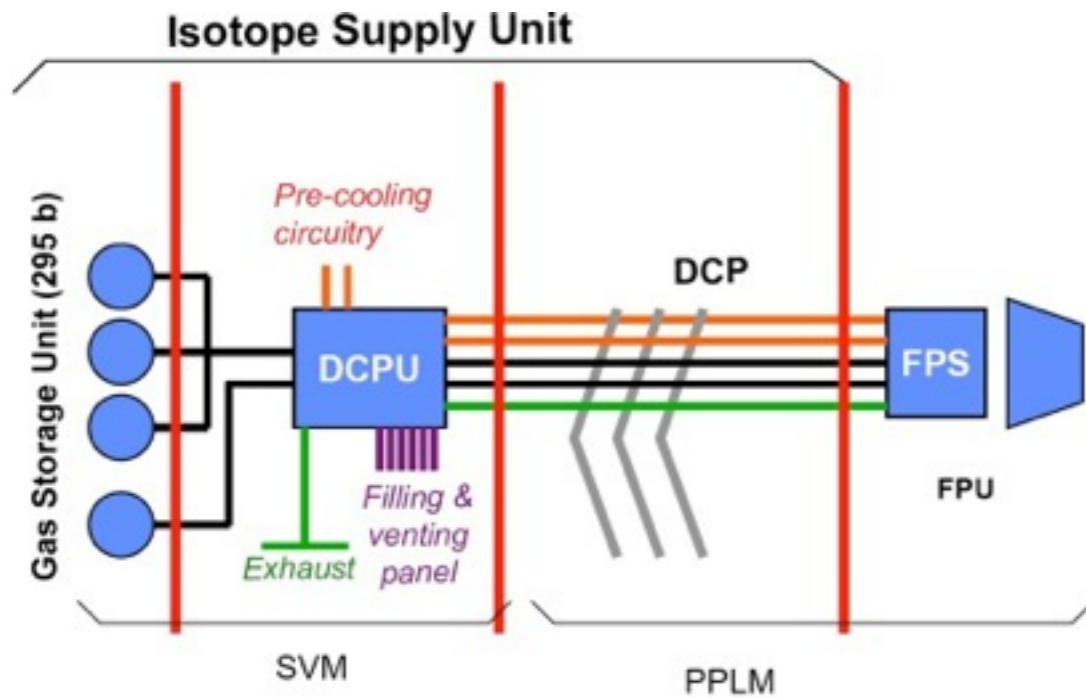


DILUTION



- Can be operated from a 4 K heat sink
- Cooling P @ T ultimate
- + available cooling continuously from 4K down to T ultim.

PLANCK DILUTION COOLER

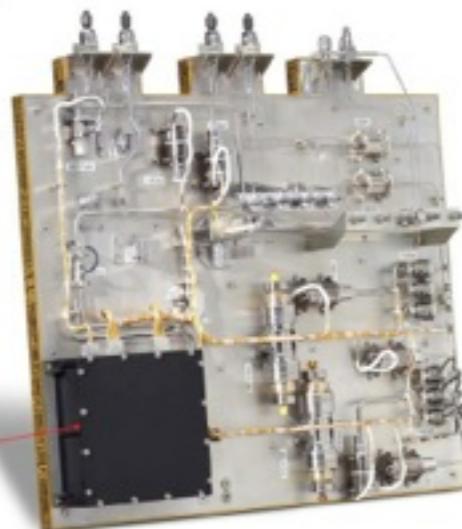


200 nW @ 100 mK

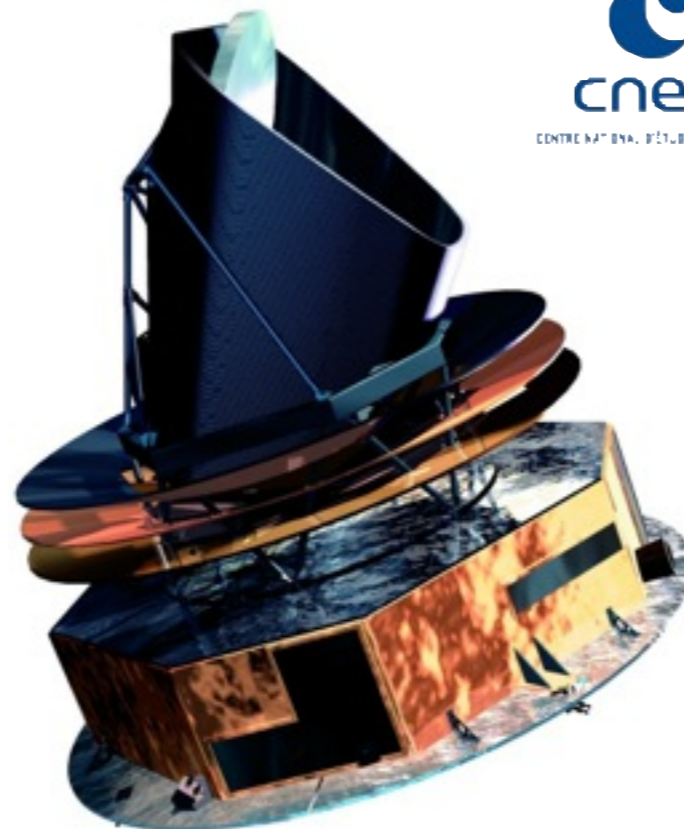
continuous cooling
from 1.6 K ($\approx 8 \mu\text{W}$)

**Open cycle:
Lifetime limited
 ≈ 2 years mission**

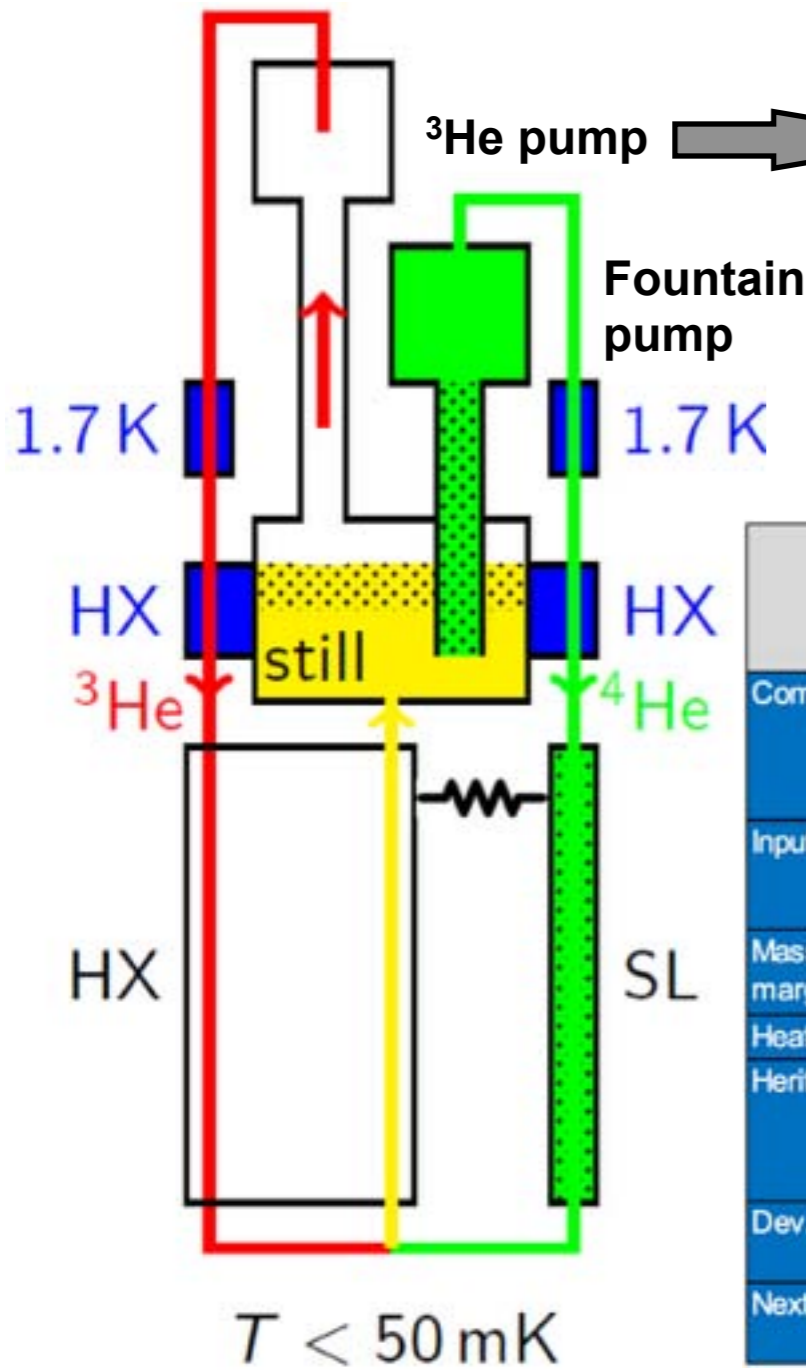
> Control electronics,
valve & Pressure
reduction, and P
measurement panel



Dilution Cooler Electronics



CONTINUOUS DILUTION COOLER



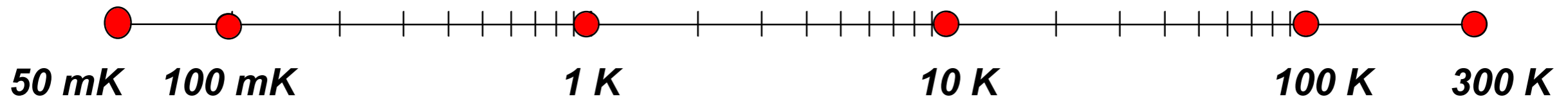
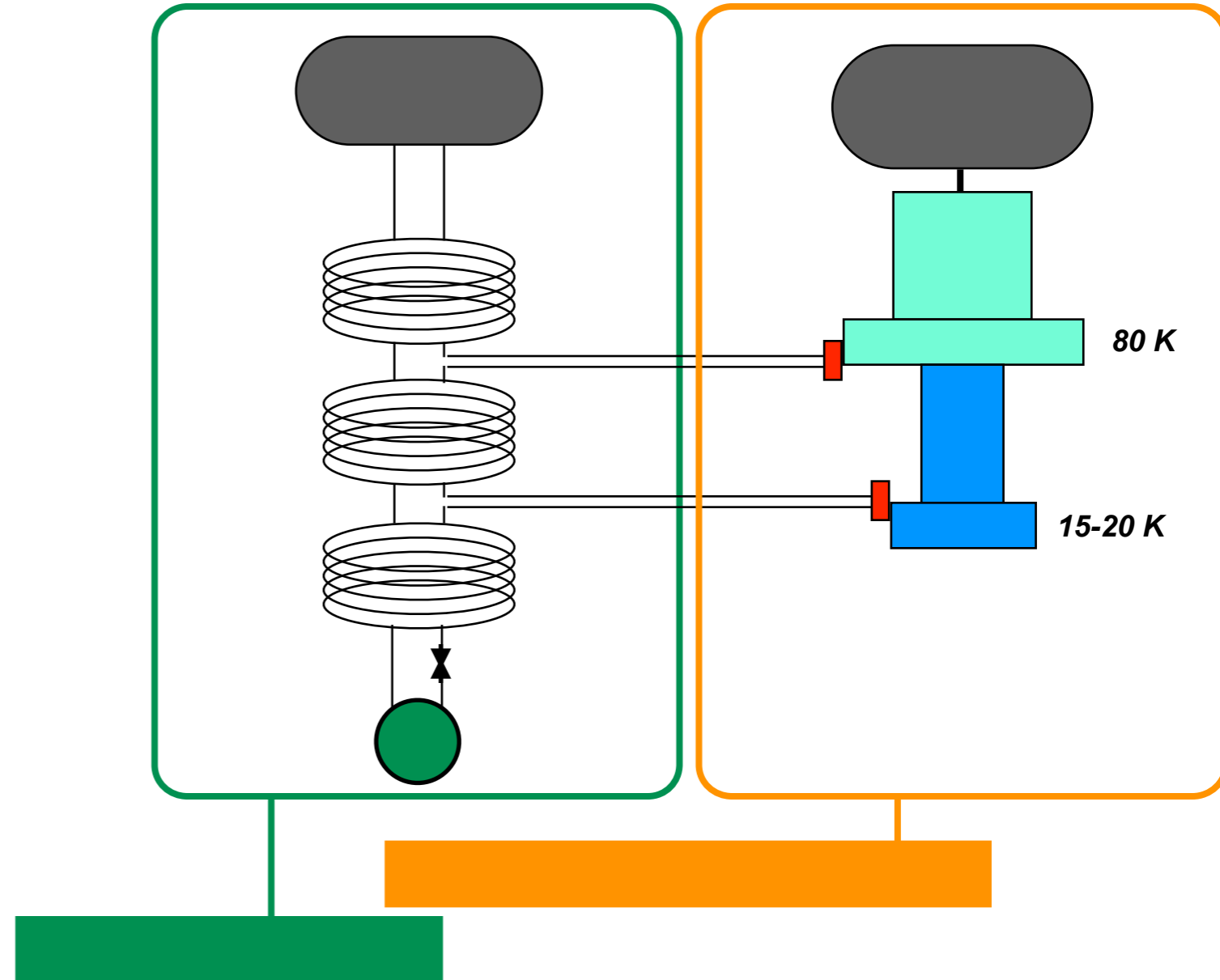
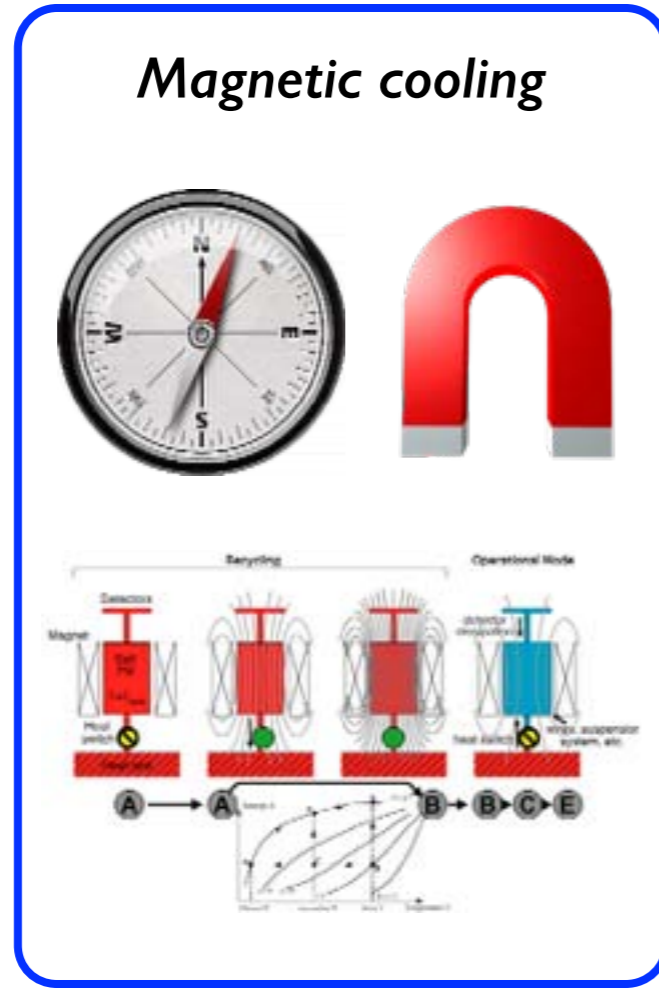
Circulate the ³He

Parameter	Linear Compressor (JAXA + SHI)	Sorption Compressor (COOL + U.Twente)	Holweck Compressor (CNRS + AL)
Compression ratio	5.4mb/140mb (@17 μ mol/s) demonstrated	5mb/200mb (@20 μ mol/s) expected	5mb/200mb (@20 μ mol/s) demonstrated
Input Power (no margin)	<80W	~10W@300K (estimated)	~100W@300K (estimated)
Mass (without electronics & margin)	~20kg	~2.2kg	~2kg
Heat lift below 300K	None	~80mW@15K	None
Heritage	Based on 1K-Class JT Cooler EM compressors	Based on the 4K JT Sorption Cooler EM for Darwin	No space heritage except for gas bearings (MELFI)
Dev. Status	BBM under assembly	Check valves demonstrated	Compression ratio demonstrated
Next development steps	BBM evaluation	15K demonstrator	Demonstrate gas bearings

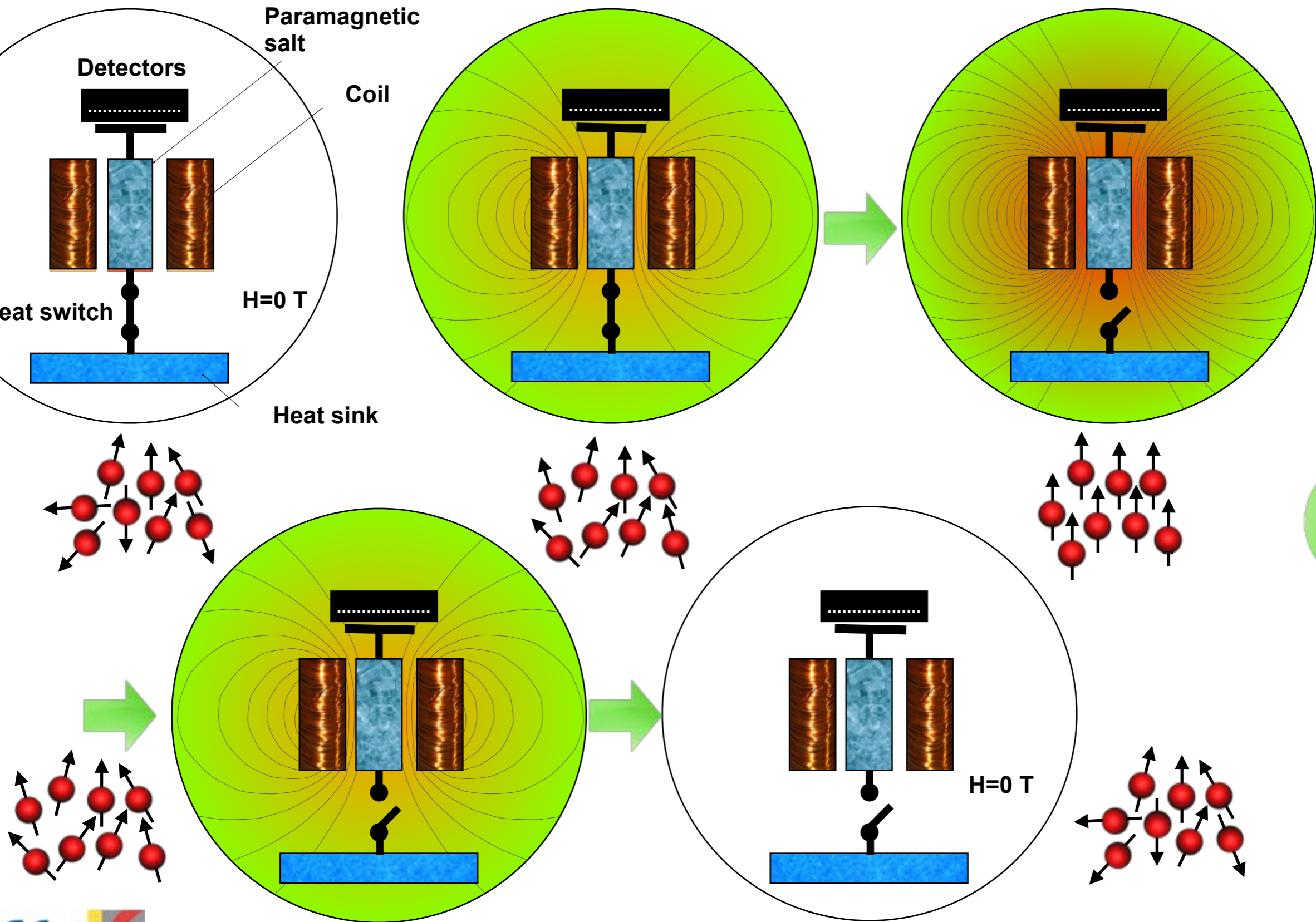


Latest result:
1 μ W @ 51 mK (liquid T !)

ADIABATIC DEMAGNETIZATION



ADIABATIC DEMAGNETIZATION



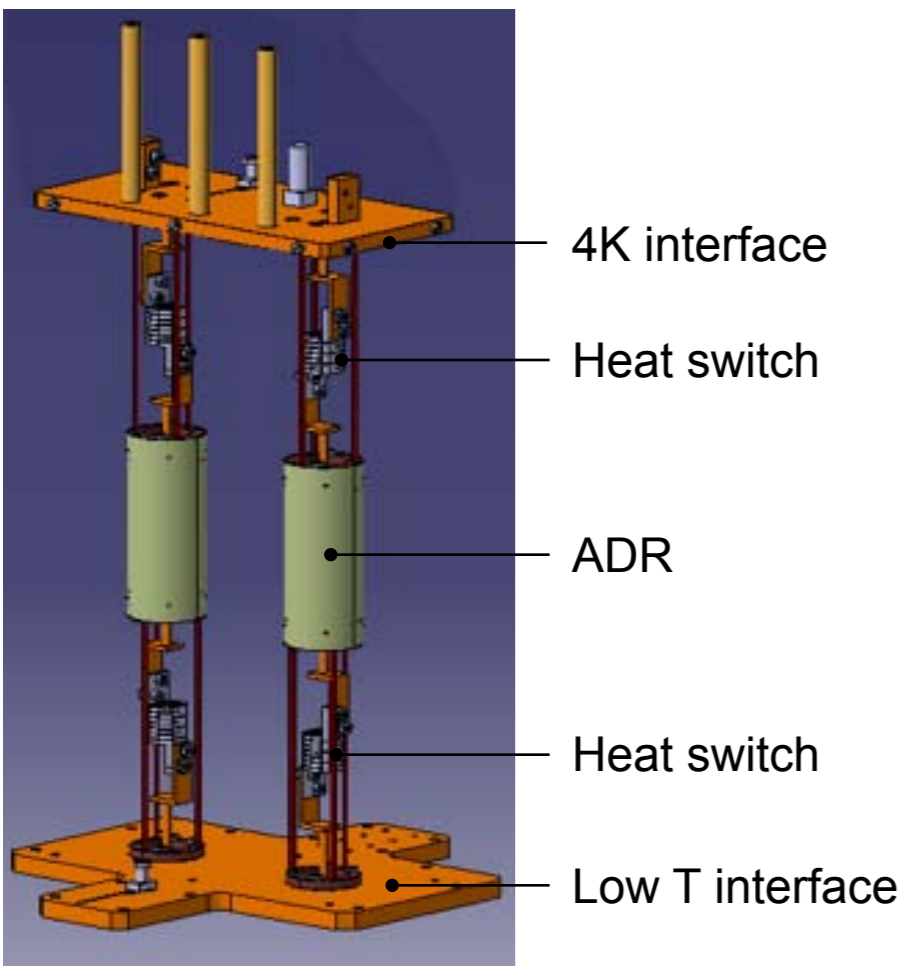
magnetic entropy

thermal entropy

RAPID RECYCLING TANDEM ADR (CONTINUOUS)

- Tandem magnetic refrigerators
- Utilizes two magnetic cooling chains
- Provides continuous cooling
- Magnets are shielded
- Single thermal interface (4 K or lower)

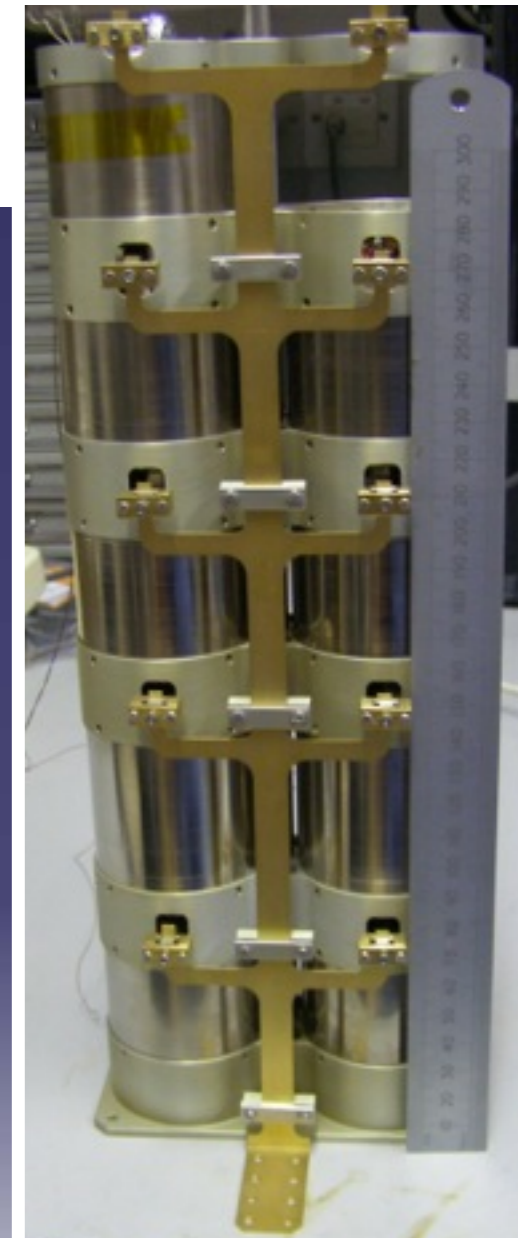
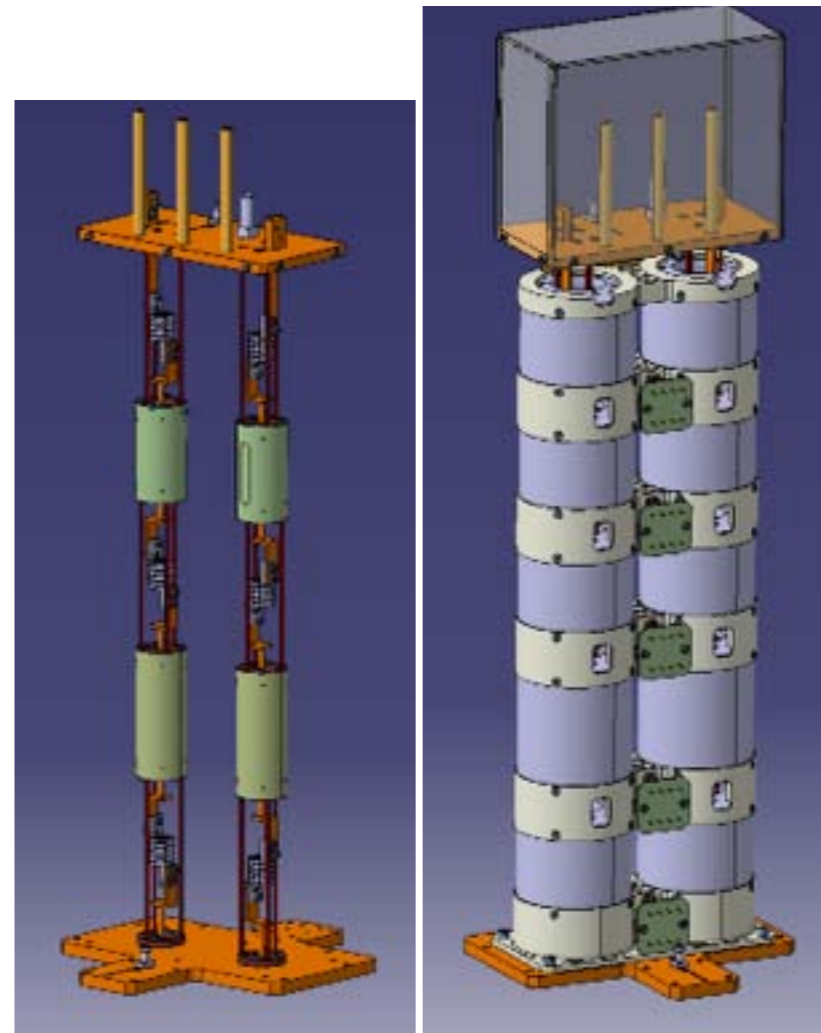
Preliminary results on single stage



200 mK - 4 K:
 Recycling time \approx 2.5 minutes
 (4K to 170 mK in 30 sec.)



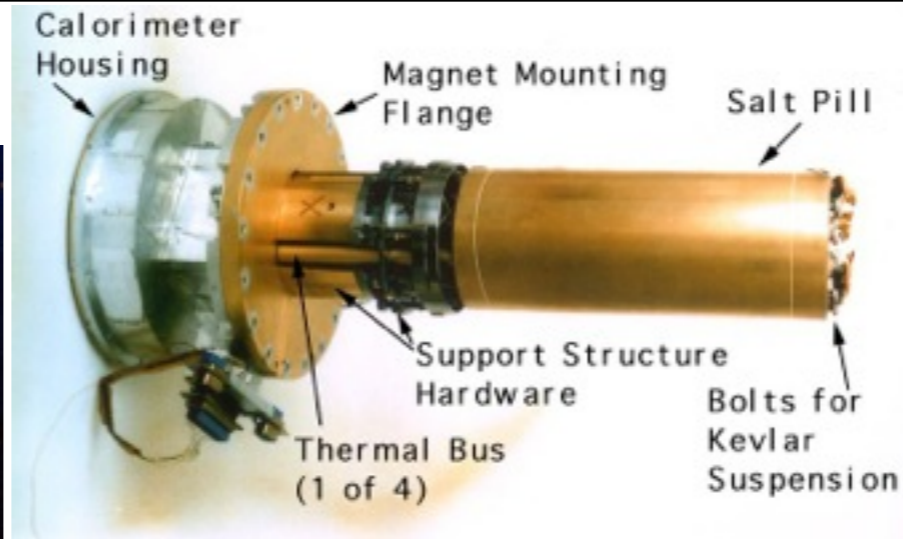
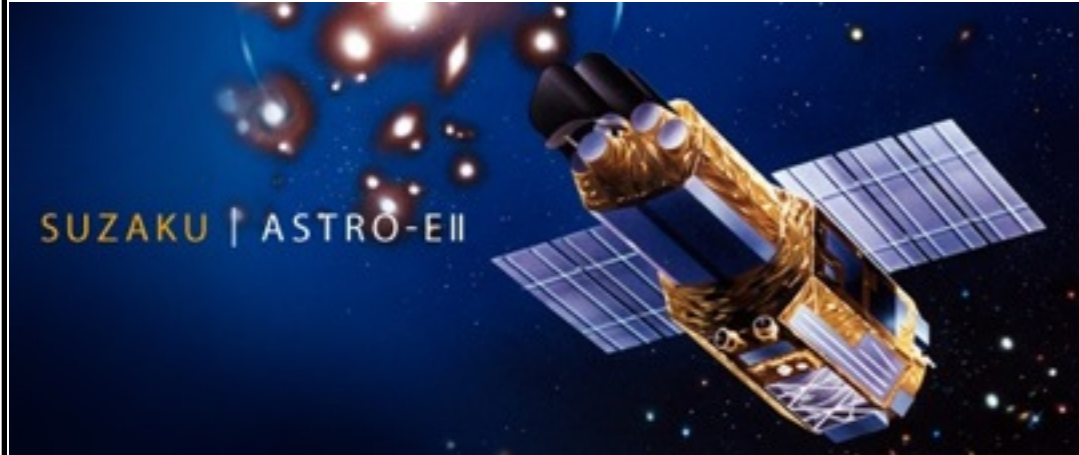
Next step



Estimated base operating T:
80 mK from 4 K
50 mK when operated from 2 K
H 355 x W 120 x D 56 mm
8.3 kg

Predicted cooling powers based on hold time of 10 minutes and operated from 4 K:
 1 μ W at 80 mK
 5 μ W at 100 mK
 41 μ W at 300 mK

ASTRO E2 (SUZAKU) ADR & CADR

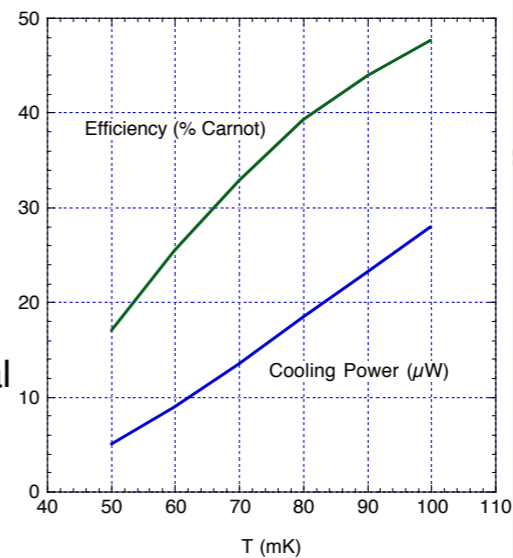
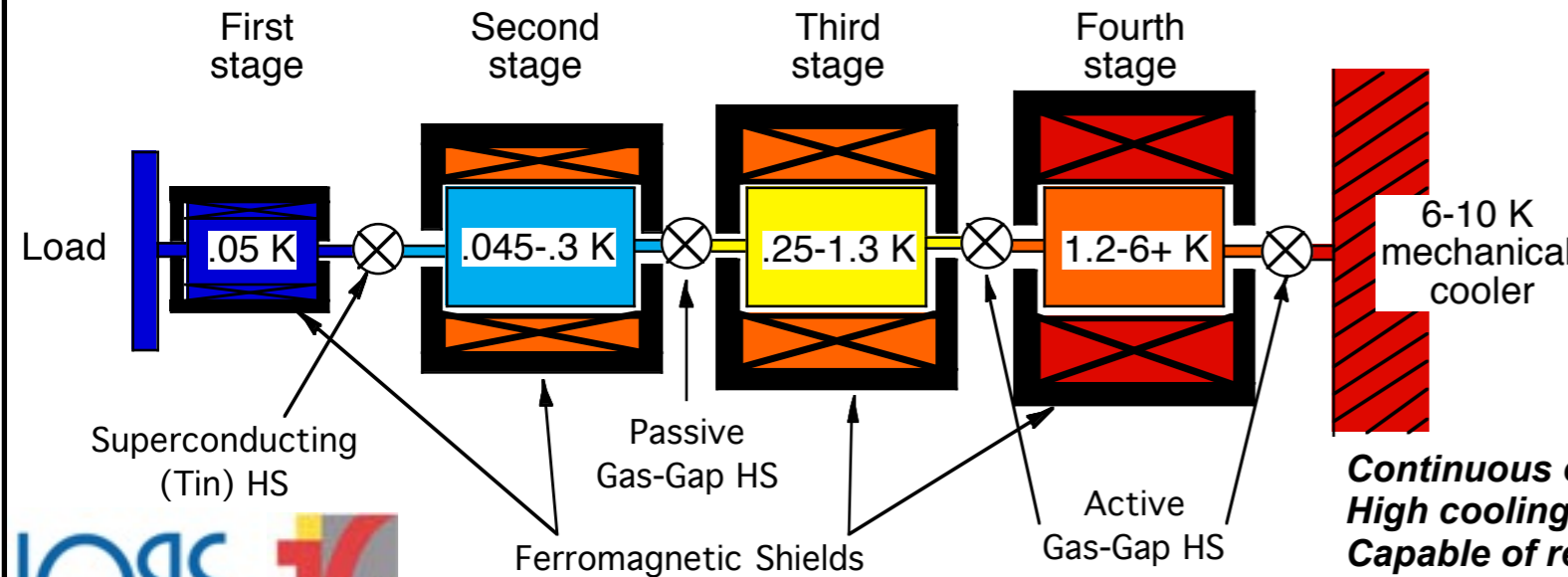


**60 mK reached on orbit !
Lowest T of (known) universe**

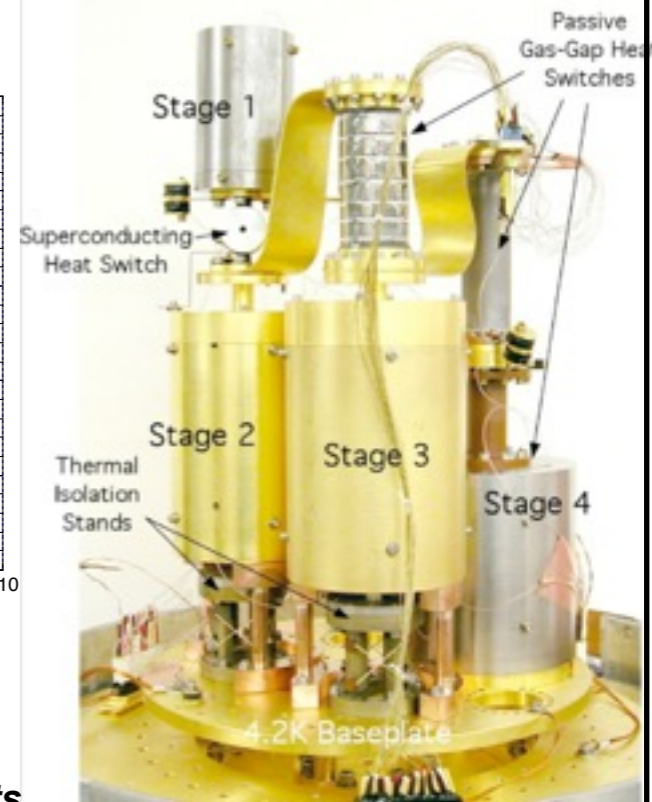


Continuous ADR

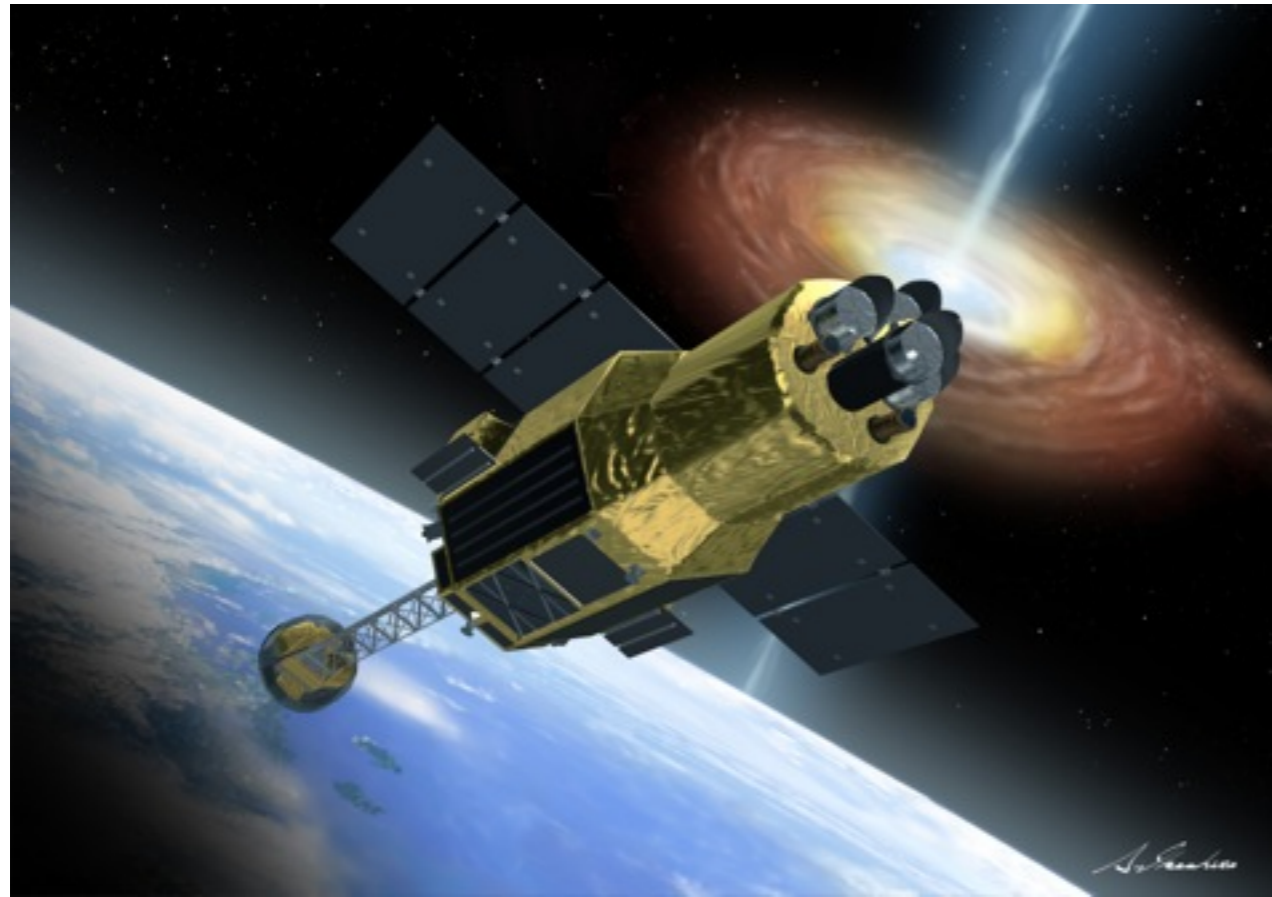
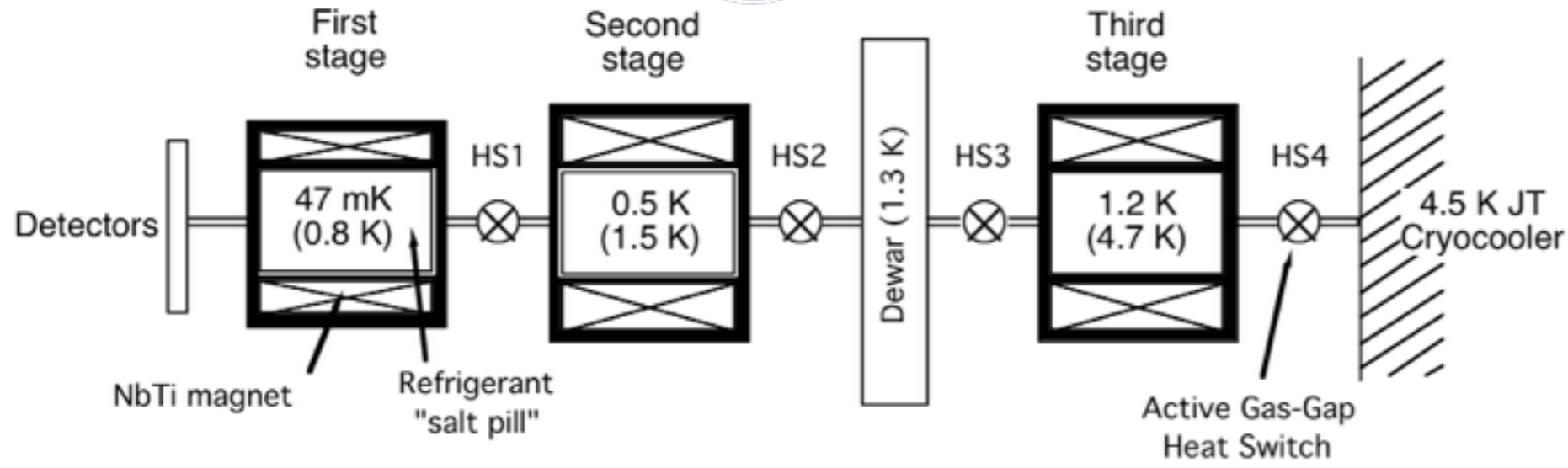
*Load is directly cooled by one stage
Upper stages cascade heat to the heat sink*



**Continuous operation
High cooling power
Capable of rejecting heat at 6+ K
Requires development of Nb₃Sn magnets**

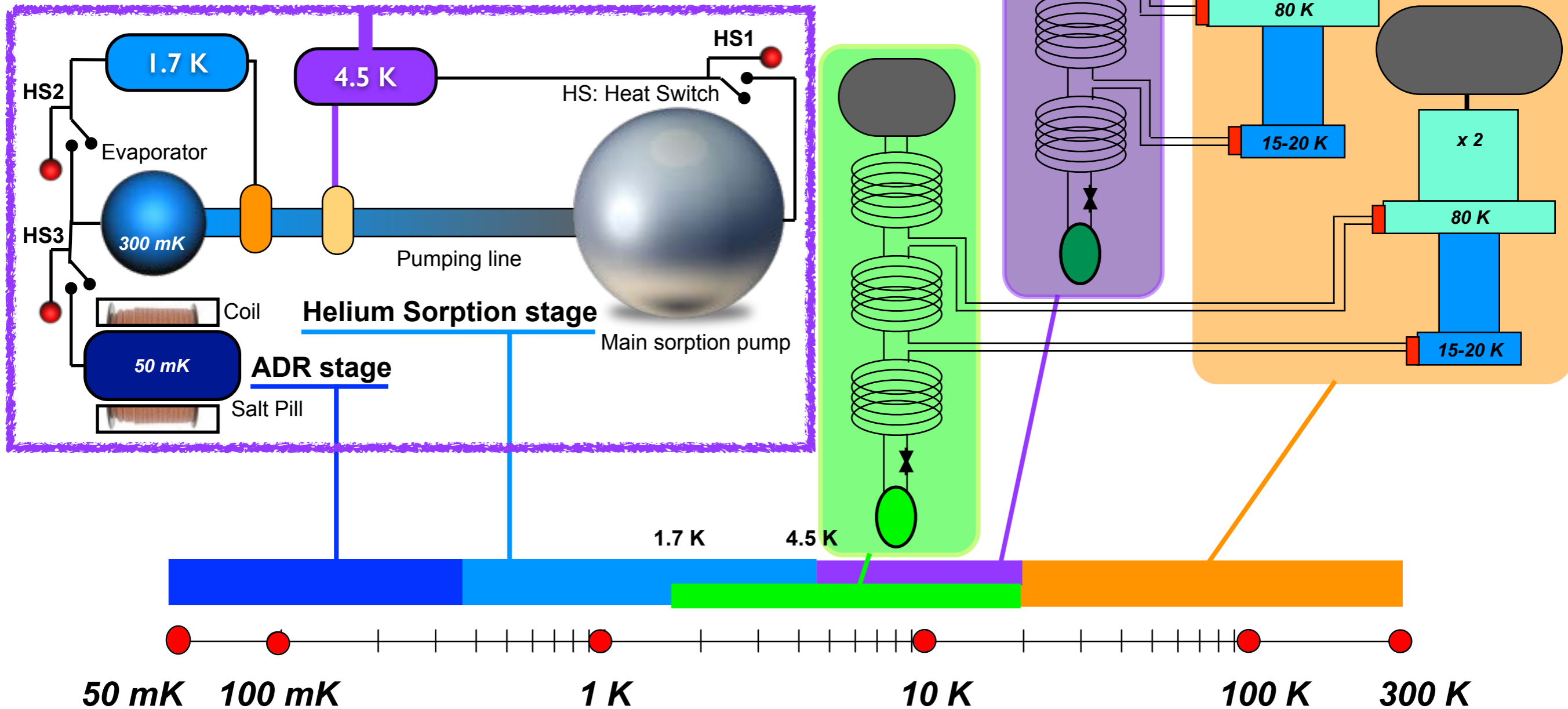


ASTRO-H: 3 STAGES ADR



50 mK COOLER HYBRID CONCEPT

Hybrid cooler: combination of a 300 mK sorption and miniature ADR stages



TWO EM MODEL DEVELOPED



ESA TRP (IXO/ATHENA)
To specifications



123 x 185 x 300 mm

5846 gr

**10 μ W @ 300 mK
(37 hours)**

**1 μ W @ 50 mK
(34 hours)**



JAXA 次世代赤外線天文衛星
esa **SPICA**
Space Infrared Telescope for Cosmology and Astrophysics
REVEALING THE ORIGINS OF PLANETS AND GALAXIES

Infrared astronomy mission \approx 5 - 210 μ m
New framework and schedule
Launch date \approx 2026

**Mission being re-optimized
(M4 selection)**

TWO EM MODEL DEVELOPED



ESA TRP (IXO/ATHENA)
To specifications



123 x 185 x 300 mm

5846 gr

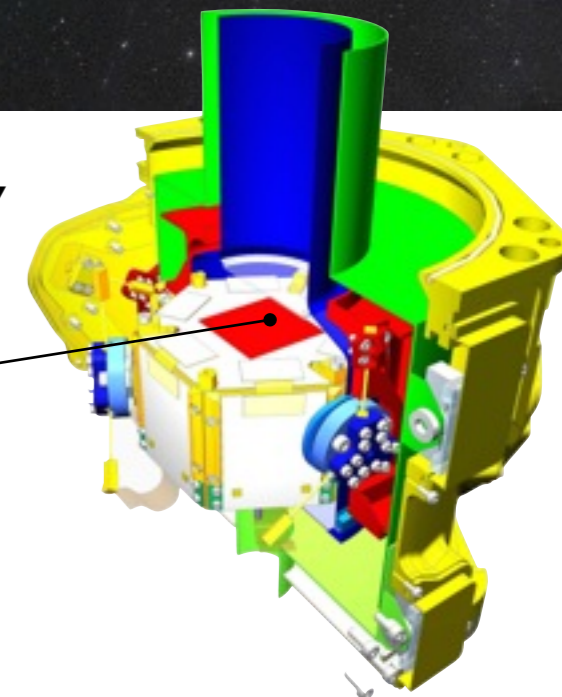
10 μ W @ 300 mK
(37 hours)

1 μ W @ 50 mK
(34 hours)

© ARTECHNOLE - www.artechnique.fr



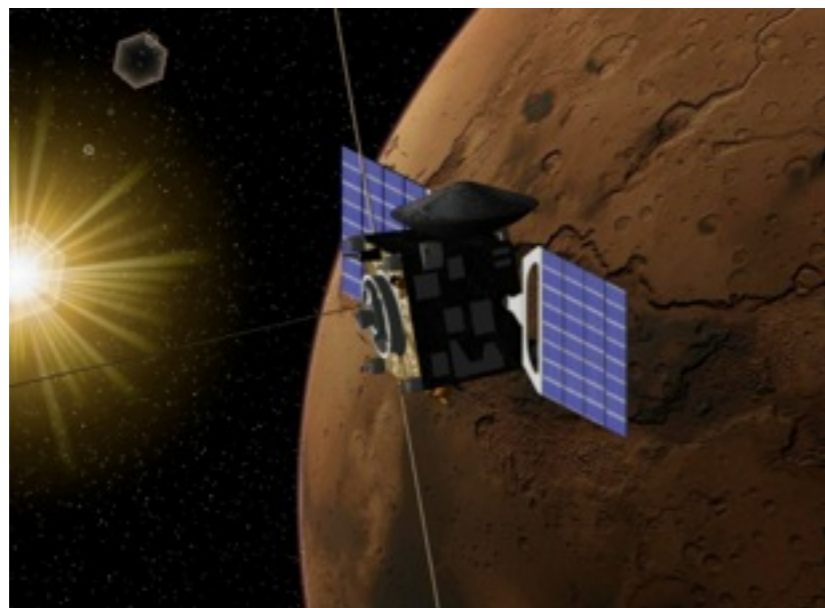
X-IFU Instrument Focal Plane Assembly



50 mK TES sensor array

THE MISSING LINKS

0.01 K 0.1 K 1 K 10 K 100 K 300 K



Passive Radiators

Stored cryogen (Cryostats)

Active "mechanical" coolers

"Ultra" low T coolers
(3 technologies: sorption, magnetic, dilution)

- Distance to detectors could be meters
- Temperature stability
- Induced vibrations
- Absorption of peak powers
- Temperature gradient
- On ground management (time constant)

FLEXIBLE LINKS

≈ Simple but not easy to make

SDL/11-224

FLEXIBLE THERMAL LINKS

Space Dynamics Laboratory (SDL) has the facilities and experience to meet the most stringent link requirements. SDL thermal links have been selected for NASA's JWST program. Full support services include thermal and dynamic testing and certification at cryogenic temperatures.

SDL's Flexible Thermal Links

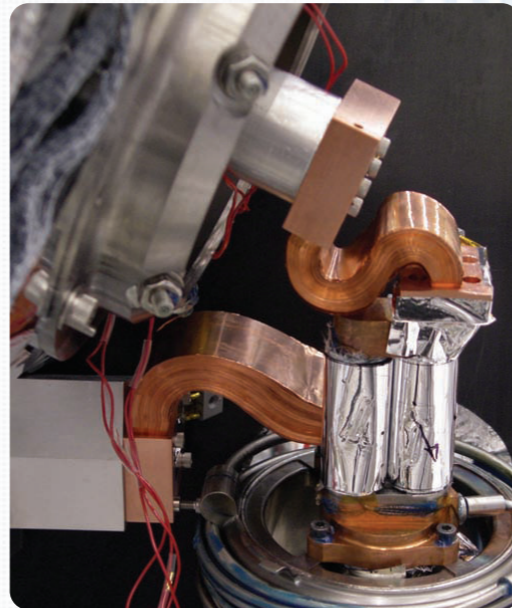
- Eliminate:** Joining materials including solder
- Internal contact resistance
- Wicking into braid/foil
- Outgassing
- Maximize:** Thermal conductance
- Dynamic/mechanical flexibility
- Provide:** High Performance
- Affordable Solutions



Appropriate material types and configurations are available based on customer-specific thermal and mechanical requirements for conductance, mass, and flexibility

SPECIFICATIONS

- CONDUCTANCE:** 0.01 - 10W/K
- STIFFNESS (flexibility):** Typically < 1 N/mm all axes
- MASS:** 5g - 10kg
- MATERIAL:** Copper, Aluminum, etc.
- TYPE:** Foil or Braid
- TRANSFER LENGTHS:** 2mm - 2m



SDL flexible thermal links attach the focal plane assembly to the cryo-coolers on the GIFTS Instrument



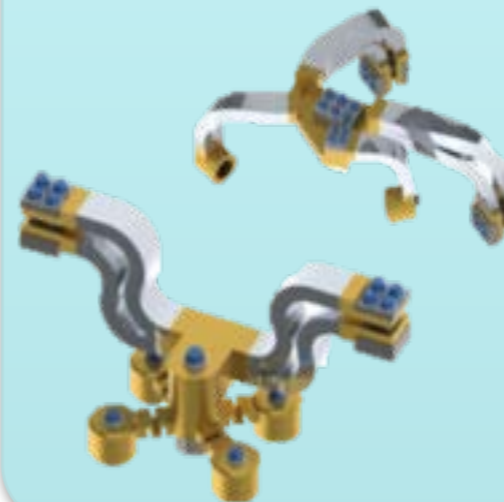
The SDL flexible thermal links are configurable to almost any desired shape and end-block configuration.



1695 North Research Park Way • North Logan, Utah 84341 • Phone 435.713.3400 • www.spacedynamics.org

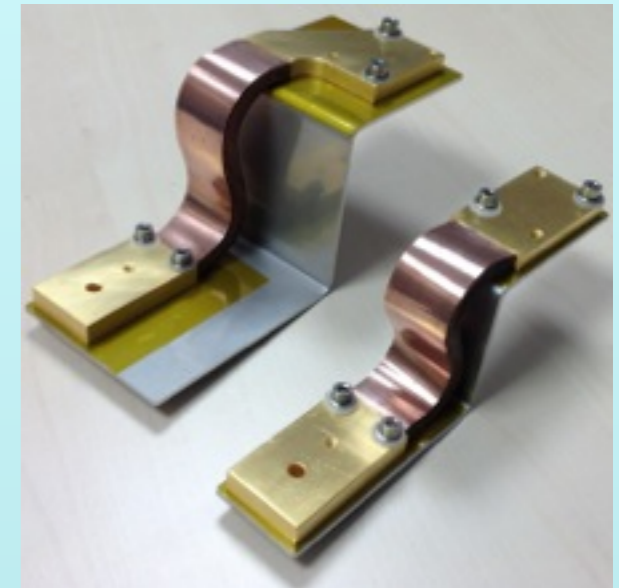


99.999% AL
 High flexibility in all direction (< 5N/mm)
 High K: >1W/K @ 80K
 Mass < 100g for complete thermal link



On-going activities in the frame of CSO and MTG

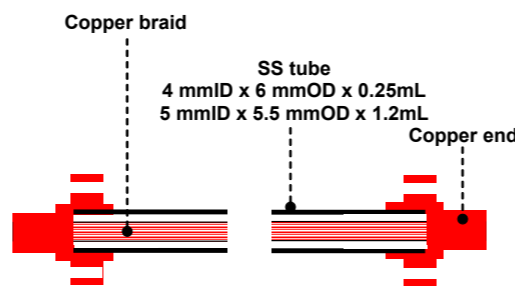
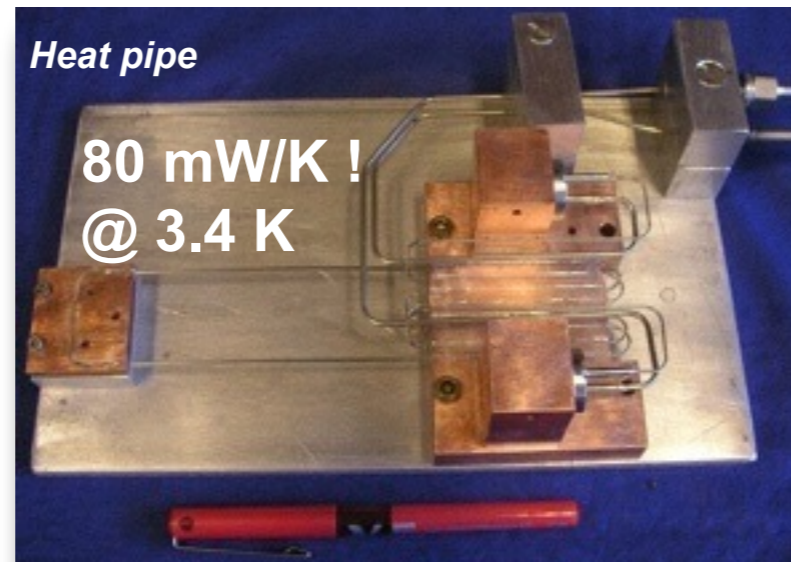
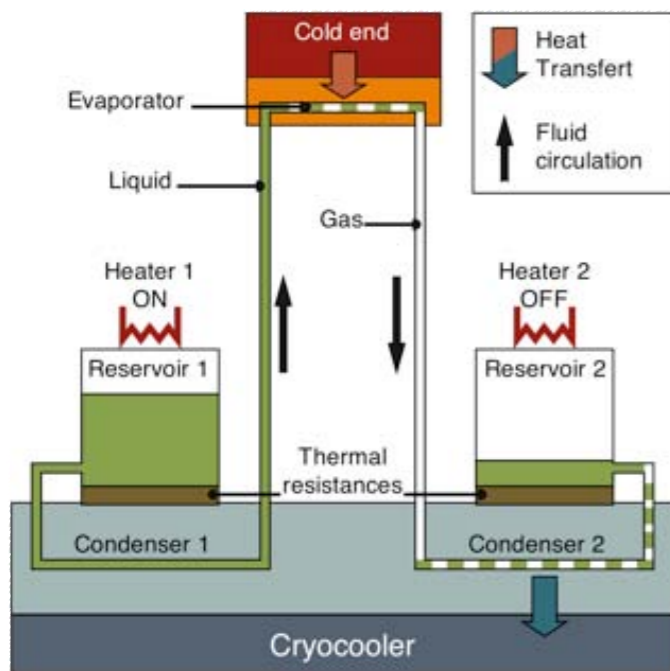
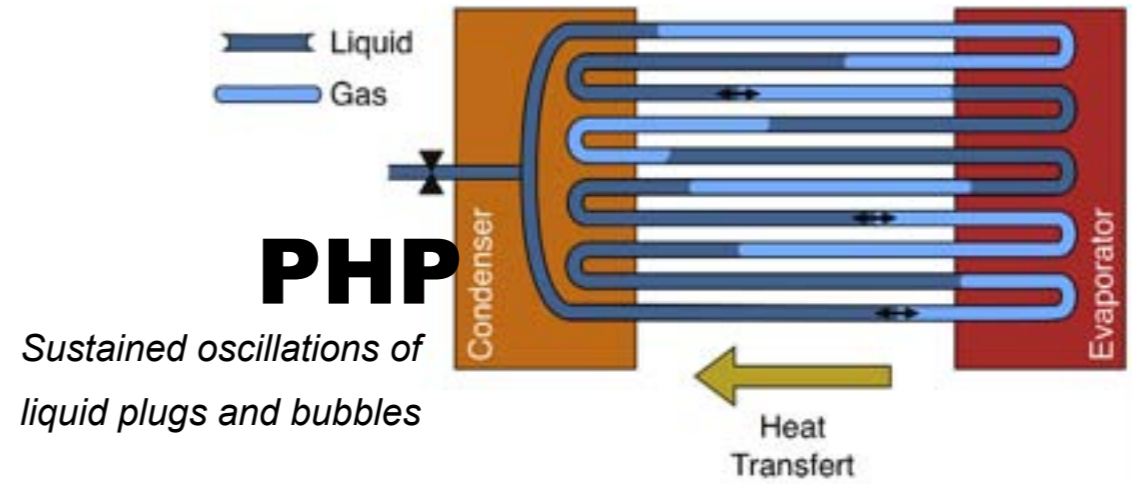
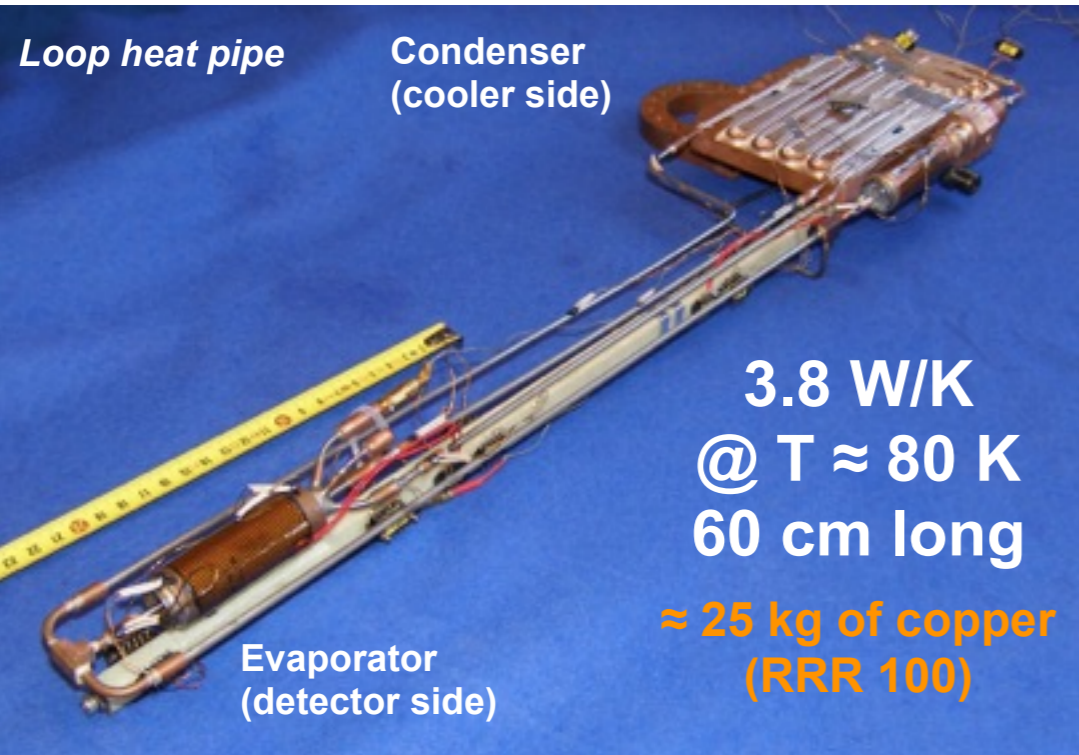
CSO: 2 sets of FMs delivered
MTG: 6 sets delivered



99.99% OFHC
 High flexibility in all direction (< 0.5N/mm)
 High K: >0.5W/K @ 300K
 Mass < 350g for complete thermal link



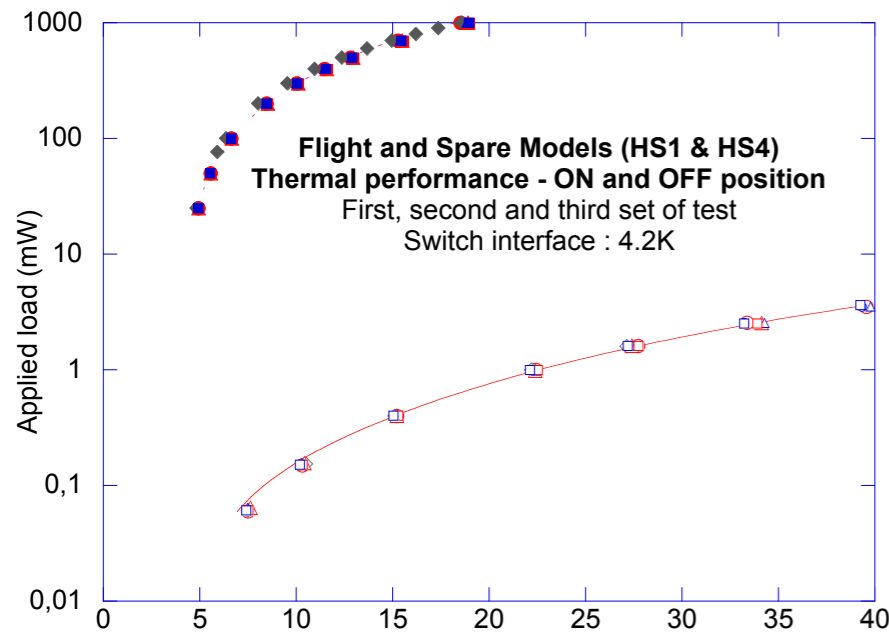
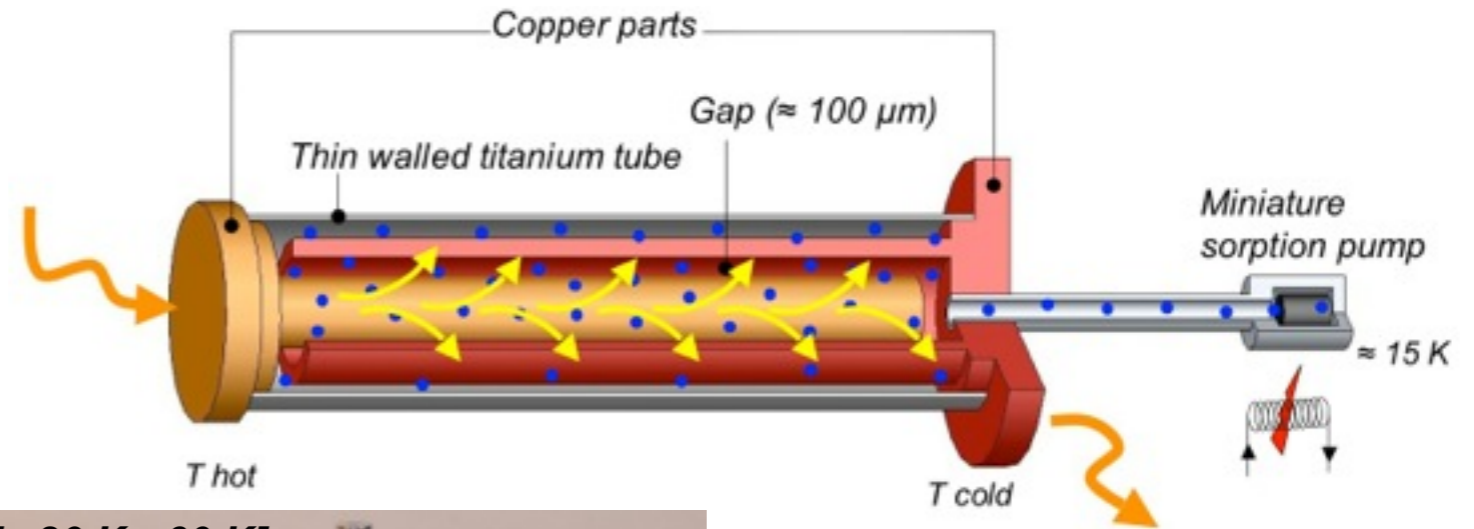
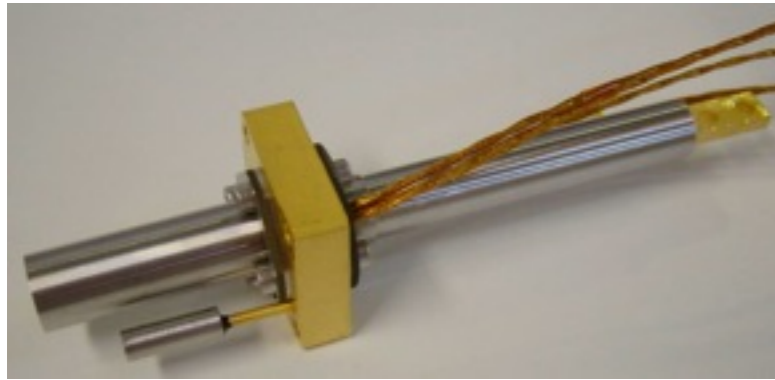
HEAT PIPE & PHP



ANCILLARY EQUIPMENT: HEAT SWITCH

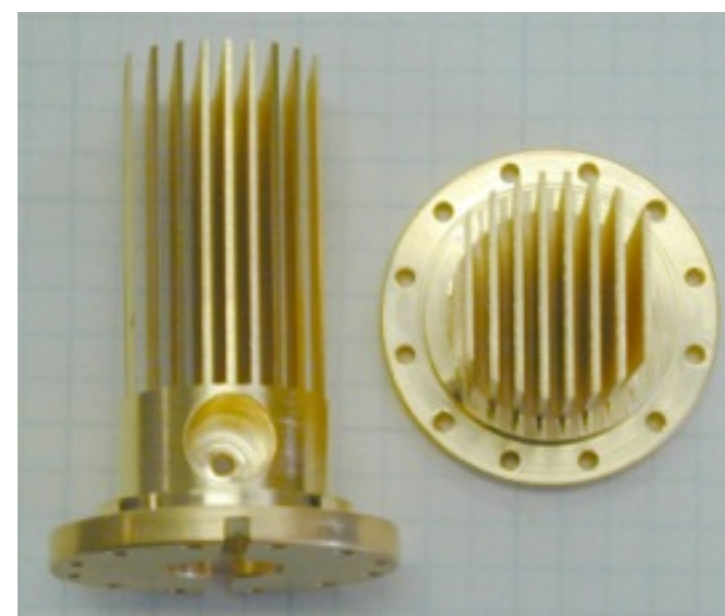
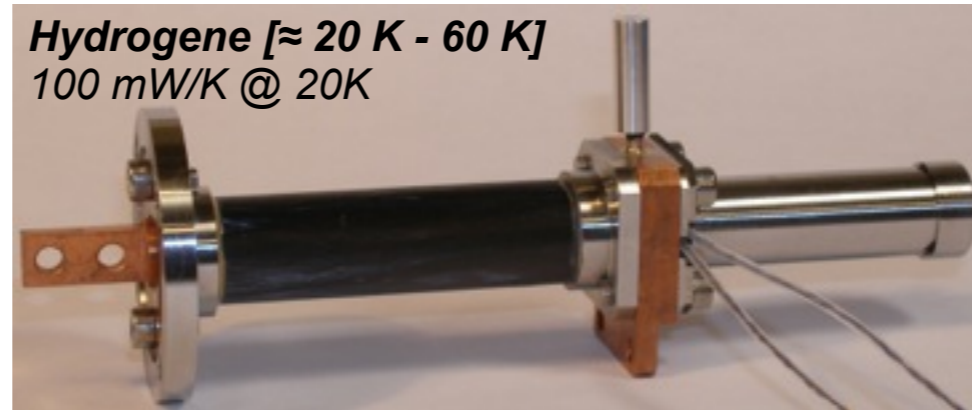
Most commonly used: Gas Gap Heat Switch

PLANCK Helium gas gap heat switch



- OFF theoretical
- HS#1 - OFF - 1st set
- HS#4 - OFF - 1st set
- ◇ HS CQM OFF
- ON theoretical
- HS#1 - ON - 1st set
- HS#4 - ON - 1st set
- ◆ HS CQM ON
- △ HS#1 - OFF - 2nd set
- △ HS#4 - OFF - 2nd set
- ▲ HS#1 - ON - 2nd set
- ▲ HS#4 - ON - 2nd set
- HS#1 - OFF - 3rd set
- HS#4 - OFF - 3rd set
- HS#1 - ON - 3rd set
- HS#4 - ON - 3rd set

- Test #1 : first set of thermal test
- Test #2 : second set performed after 24h 60°C bake out
- Test #3 : third set performed after the vibrations tests

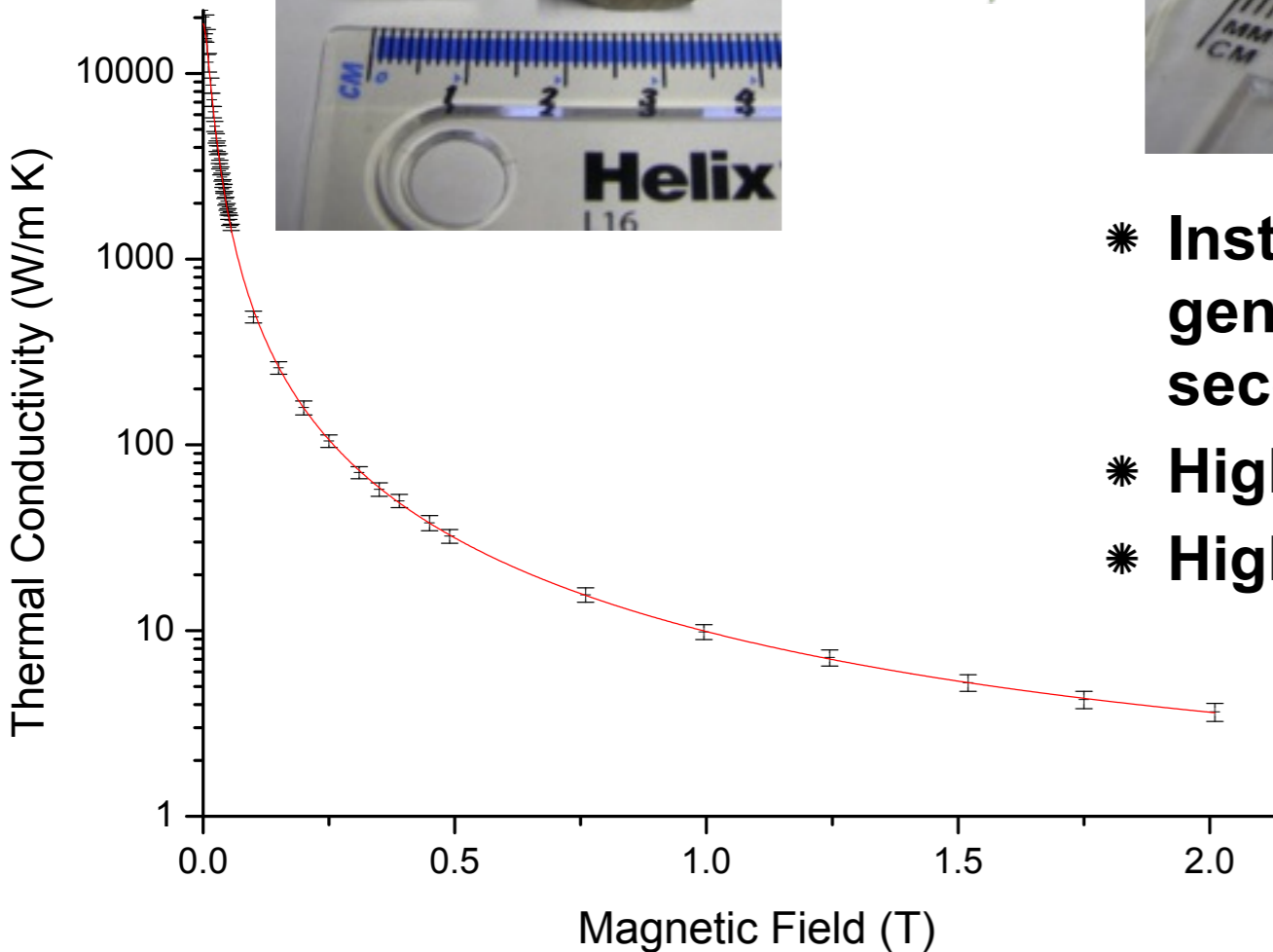
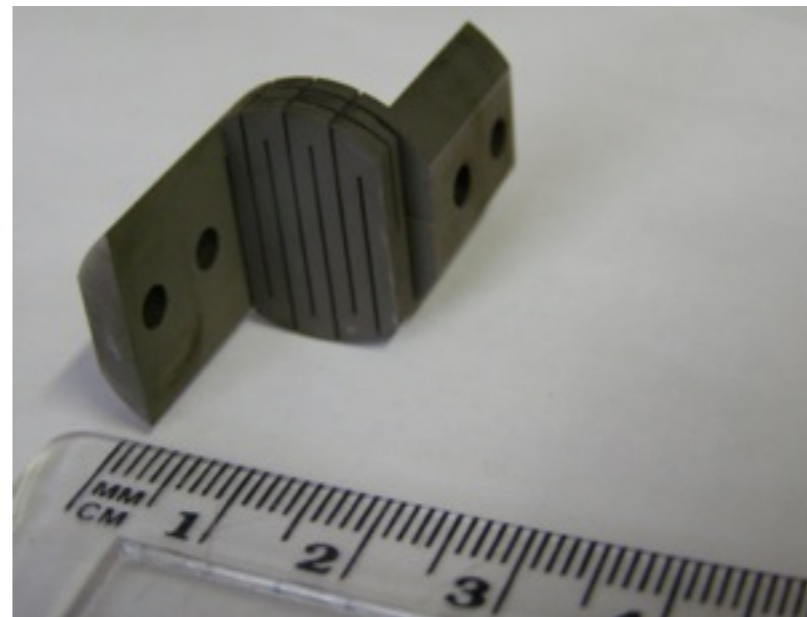
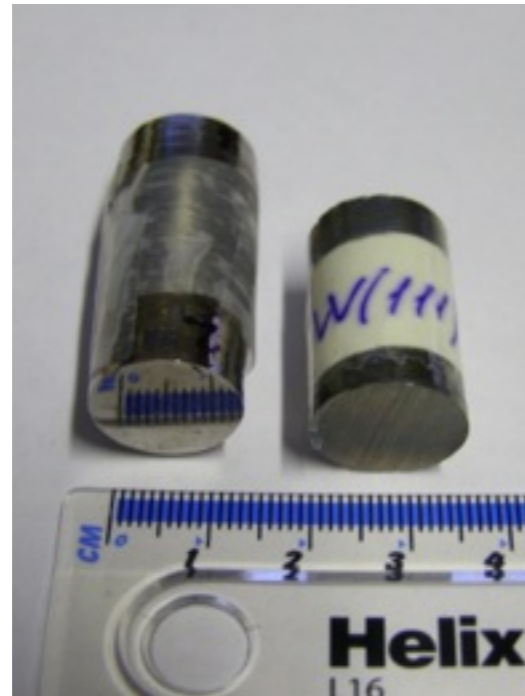


Other geometry (NASA Goddard)

MAGNETORESISTIVE HEAT SWITCH

Magnetoresistive heat switch made from tungsten

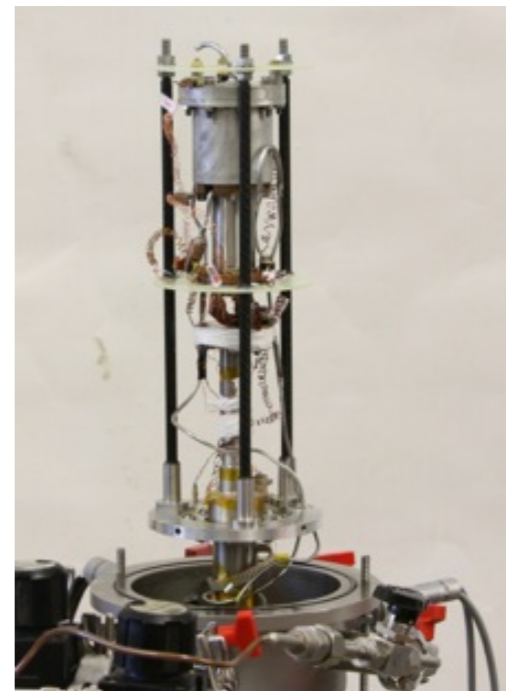
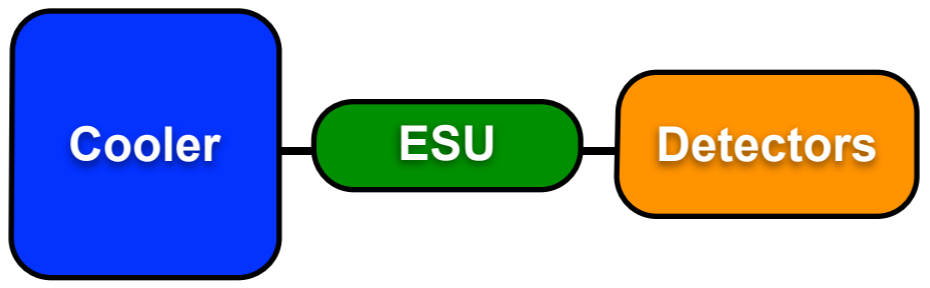
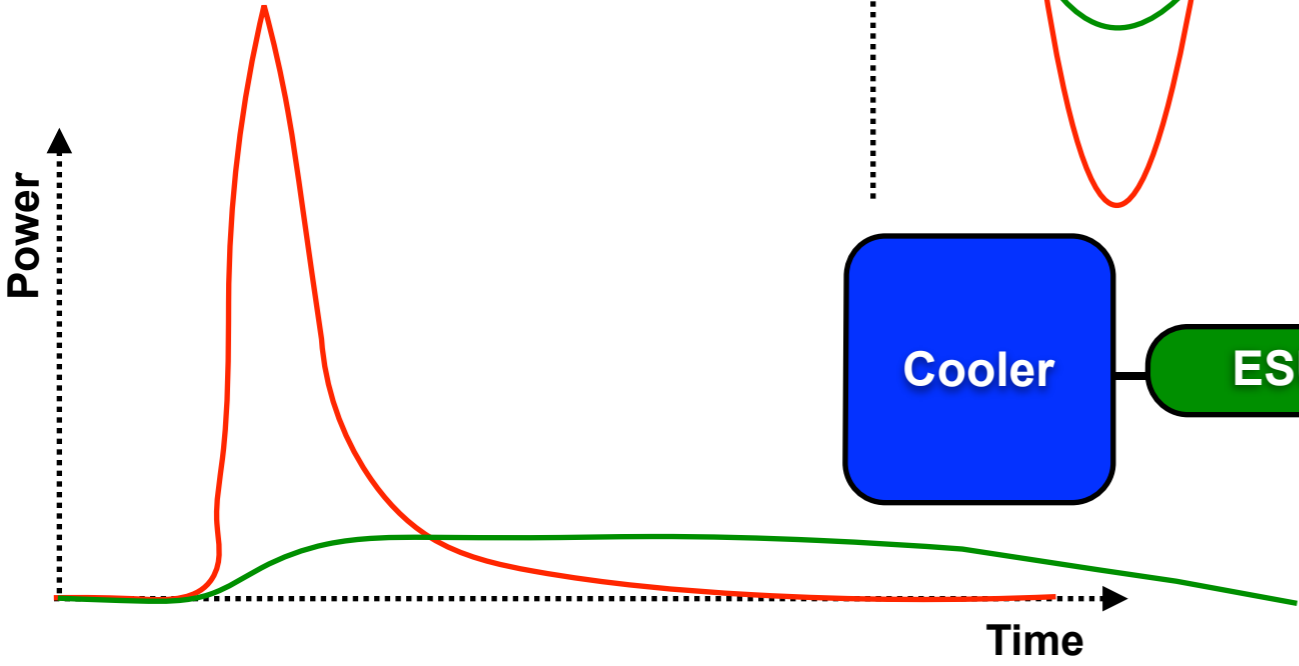
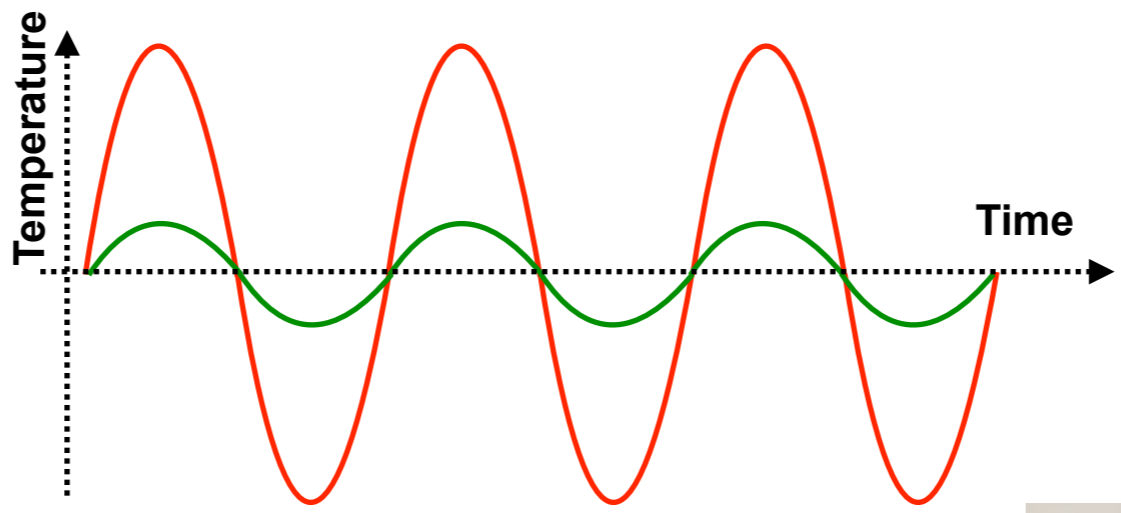
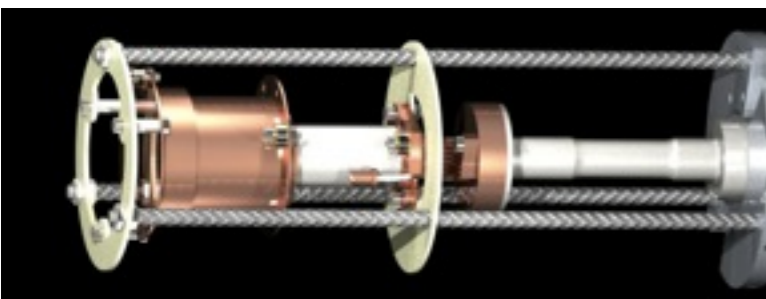
Works by the slowing down of electrons in a metal due to a tangential magnetic field



- * **Instantaneous switching – limited by speed of generating the magnetic field, currently 30 seconds to go from 0 to 2 Tesla**
- * **High thermal conductivity 200 W/cm/K**
- * **High switching ratio $\sim 10^4$ (related to the magnetic field)**



ENERGY STORAGE UNIT



Several applications

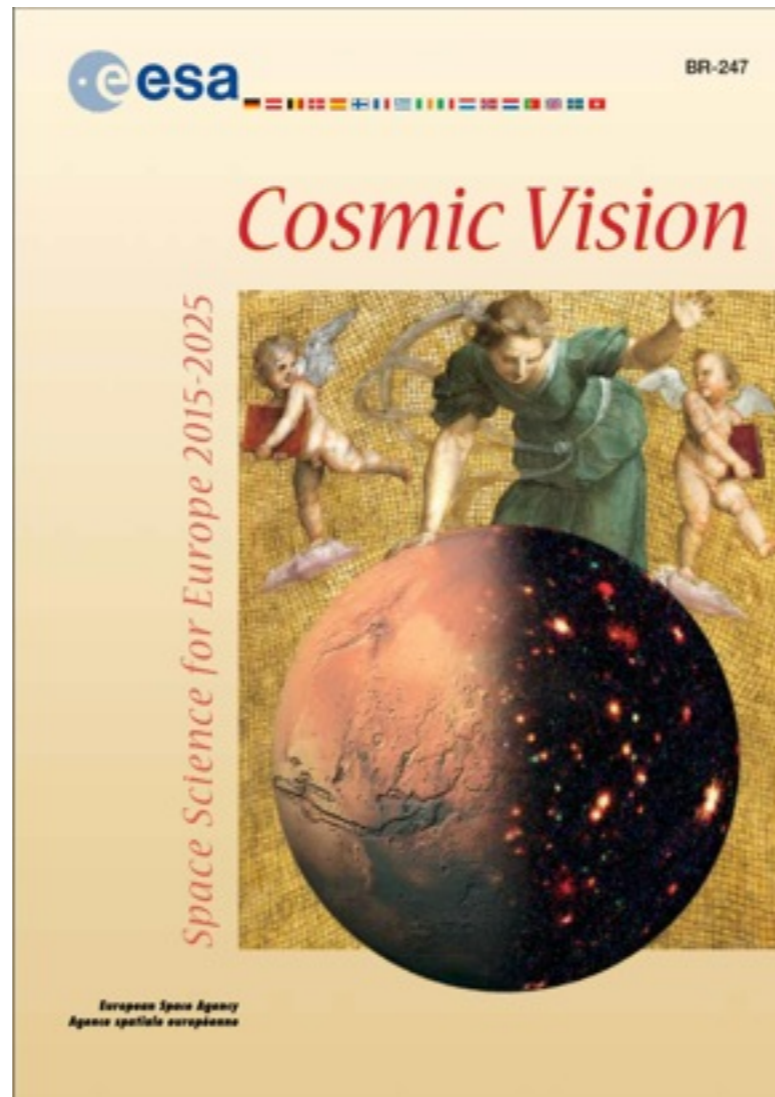
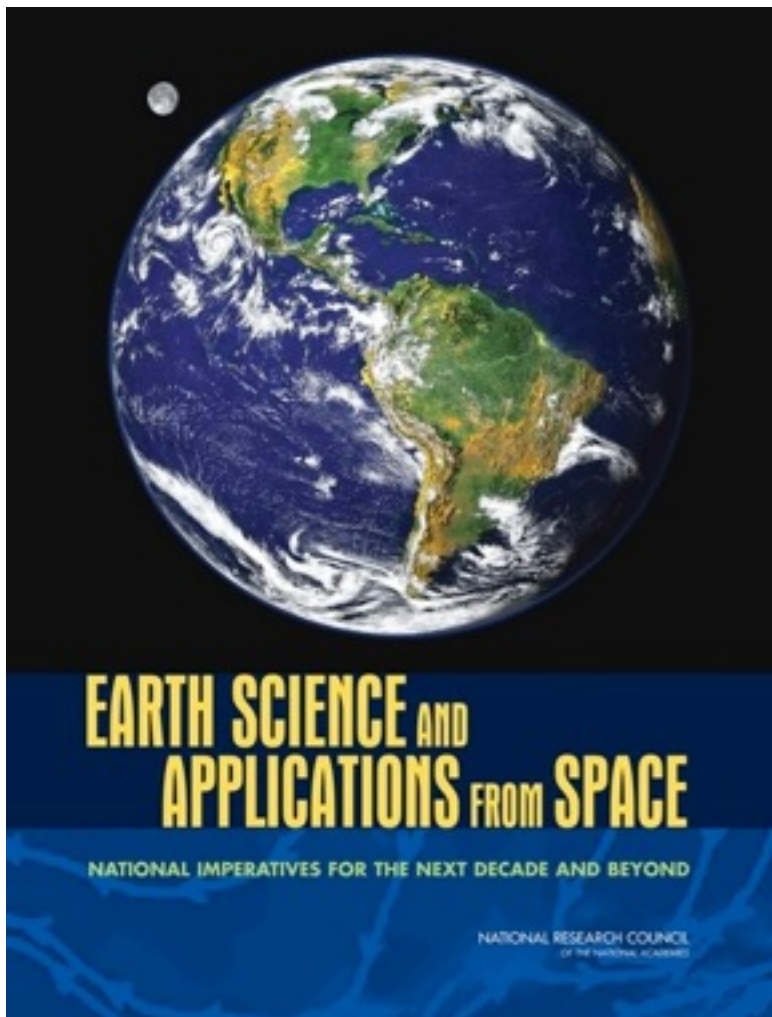
- Attenuate temperature oscillations
- Absorbs peak power (then energy dumped over longer period)
- cooler turned OFF: provides stable operation for a limited time

Example of Thermal Buffers:

Liquid-Gas Hydrogen
 400 J between 15 and 16 K
 $V \approx 20 \text{ cm}^3$

Liquid-Gas Neon
 1000 J at 40 K **+/- 50 mK**
 2000 J between 38 and 40 K
 35 cm³

TO ETERNITY AND BEYOND



- Tonny Benschop
- Grégoire Bonfait
- Tom Bradshaw
- James Butterworth
- Benoit Chidaine
- Bernard Collaudin
- Ian Hepburn
- Dean Johnson
- Sylvain Martin
- Jeff Olson
- Keisuke Shinozaki
- Peter Shirron
- Julien Tanchon
- Thierry Tirolien
- Marcel Ter Brake
- David Valentini
- Stuart Watson

Peter Shirron - Cold facts Feb. 2014

These systems will deliver what the community needs: better access to cold in space. But you know what will happen. Scientists will develop better detectors, and demand lower temperature, longer mission lifetimes, lower cost, mass, size, etc. We should hope for no less. That's what will keep space cryogenics as exciting and relevant into the next 50 years as it has been over the last. 